PHYSICS 140B : STATISTICAL PHYSICS HW ASSIGNMENT #2 SOLUTIONS

(1) Turkey typically cooks at a temperature of 350° F. Calculate the total electromagnetic energy inside an over of volume $V = 1.0 \text{ m}^3$ at this temperature. Compare it to the thermal energy of the air in the oven at the same temperature.

The total electromagnetic energy is

$$E = 3pV = \frac{\pi^2}{15} \frac{V (k_{\rm B}T)^4}{(\hbar c)^3} = 3.78 \times 10^{-5} \,\mathrm{J} \quad .$$

For air, which is a diatomic ideal gas, we have $E = \frac{5}{2}pV$. What do we take for p? If we assume that oven door is closed at an initial temperature of 63° F which is 300 K, then with a final temperature of 350° F = 450 K, we have an increase in the absolute temperature by 50%, hence a corresponding pressure increase of 50%. So we set $p = \frac{3}{2}$ atm and we have

$$E = rac{5}{2} \cdot rac{3}{2} (1.013 imes 10^5 \, {
m Pa}) (1.0 \, {
m m}^3) = 3.80 imes 10^5 \, {
m J} ~,$$

which is about ten orders of magnitude larger.

(2) Let *L* denote the number of single particle energy levels and *N* the total number of particles for a given system. Find the number of possible *N*-particle states $\Omega(L, N)$ for each of the following situations:

(a) Distinguishable particles with L = 3 and N = 3.

 $\Omega_{\rm D}(3,3) = 3^3 = 27.$

(b) Bosons with L = 3 and N = 3.

 $\Omega_{\rm BE}(3,3) = {5 \choose 3} = 10.$

(c) Fermions with L = 10 and N = 3.

 $\Omega_{\rm FD}(10,3) = \binom{10}{3} = 120.$

(d) Find a general formula for $\Omega_{\rm D}(L, N)$, $\Omega_{\rm BE}(L, N)$, and $\Omega_{\rm FD}(L, N)$.

The general results are

$$\Omega_{\rm D}(L,N) = L^N$$
, $\Omega_{\rm BE}(L,N) = \binom{N+L-1}{N}$, $\Omega_{\rm FD}(L,N) = \binom{L}{N}$

(3) A species of noninteracting quantum particles in d = 2 dimensions has dispersion $\varepsilon(\mathbf{k}) = \varepsilon_0 |\mathbf{k}\ell|^{3/2}$, where ε_0 is an energy scale and ℓ a length.

(a) Assuming the particles are S = 0 bosons obeying photon statistics, compute the heat capacity C_V .

The density of states is

$$g(\varepsilon) = \frac{1}{2\pi} \frac{k}{d\varepsilon/dk} = \frac{k^{1/2}}{3\pi\varepsilon_0 \,\ell^{3/2}} = \frac{\varepsilon^{1/3}}{3\pi\ell^2 \varepsilon_0^{4/3}} \quad .$$

The total energy is

$$E = A \int_{0}^{\infty} d\varepsilon \ g(\varepsilon) \ \frac{\varepsilon}{e^{\varepsilon/k_{\rm B}T} - 1} = \frac{A}{3\pi\ell^2} \, \Gamma\left(\frac{7}{3}\right) \zeta\left(\frac{7}{3}\right) \frac{(k_{\rm B}T)^{7/3}}{\varepsilon_0^{4/3}} \quad ,$$

where *A* is the system area. Thus,

$$C_A(T) = \left(\frac{\partial E}{\partial T}\right)_{\!\!A} = \frac{Ak_{\scriptscriptstyle\rm B}}{3\pi\ell^2}\,\Gamma\!\left(\frac{10}{3}\right)\zeta\!\left(\frac{7}{3}\right)\left(\frac{k_{\scriptscriptstyle\rm B}T}{\varepsilon_0}\right)^{\!\!4/3} \quad . \label{eq:CA}$$

(b) Assuming the particles are S = 0 bosons, is there an Bose condensation transition? If yes, compute the condensation temperature $T_c(n)$ as a function of the particle density. If no, compute the low-temperature behavior of the chemical potential $\mu(n, T)$.

The following integral may be useful:

$$\int_{0}^{\infty} \frac{u^{s-1} du}{e^u - 1} = \Gamma(s) \sum_{n=1}^{\infty} n^{-s} \equiv \Gamma(s) \zeta(s) \quad ,$$

where $\Gamma(s)$ is the gamma function and $\zeta(s)$ is the Riemann zeta-function.

The condition for Bose-Einstein condensation is

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$$n = \int_{0}^{\infty} d\varepsilon \, g(\varepsilon) \, \frac{1}{e^{\varepsilon/k_{\rm B}T_{\rm c}} - 1} = \frac{1}{3\pi\ell^2} \, \Gamma\left(\frac{4}{3}\right) \zeta\left(\frac{4}{3}\right) \left(\frac{k_{\rm B}T_{\rm c}}{\varepsilon_0}\right)^{4/3} \quad .$$

Thus,

$$T_{\rm c} = \frac{\varepsilon_0}{k_{\rm B}} \left(\frac{3\pi\ell^2 n}{\Gamma(\frac{4}{3})\,\zeta(\frac{4}{3})} \right)^{3/4} \quad . \label{eq:Tc}$$

(4) Hydrogen (H_2) freezes at 14 K and boils at 20 K under atmospheric pressure. The density of liquid hydrogen is 70 kg/m^3 . Hydrogen molecules are bosons. No evidence has been found for Bose-Einstein condensation of hydrogen. Why not?

If we treat the H_2 molecules as bosons, and we ignore the rotational freedom, which is appropriate at temperatures below $\Theta_{rot} = 85.4$ K, we have

$$T_{\rm c} = \frac{2\pi\hbar^2}{mk_{\rm B}} \left(\frac{n}{\zeta(\frac{3}{2})}\right)^{2/3} = 6.1\,{\rm K} \quad . \label{eq:Tc}$$

Thus, the critical temperature for ideal gas Bose-Einstein condensation is significantly below the freezing temperature for H_2 . The freezing transition into a regular solid preempts any BEC phenomena.

(5) (Difficult) Consider a three-dimensional Bose gas of particles which have two internal polarization states, labeled by $\sigma = \pm 1$. The single particle energies are given by

$$arepsilon(m{k},\sigma) = rac{\hbar^2m{k}^2}{2m} + \sigma\Delta$$
 ,

where $\Delta > 0$.

(a) Find the density of states per unit volume $g(\varepsilon)$.

Let $g_0(\varepsilon)$ be the DOS per unit volume for the case $\Delta = 0$. Then

$$g_0(\varepsilon) \, d\varepsilon = \frac{d^3\!k}{(2\pi)^3} = \frac{k^2 \, dk}{2\pi^2} \quad \Rightarrow \quad g_0(\varepsilon) = \frac{2}{\sqrt{\pi}} \left(\frac{m}{2\pi\hbar^2}\right)^{\!\!3/2} \!\!\varepsilon^{1/2} \, \Theta(\varepsilon)$$

For finite Δ , the single particle energies are shifted uniformly by $\pm \Delta$ for the $\sigma = \pm 1$ states, hence

$$g(\varepsilon) = g_0(\varepsilon + \Delta) + g_0(\varepsilon - \Delta)$$
 .

(b) Find an implicit expression for the condensation temperature $T_c(n, \Delta)$. When $\Delta \to \infty$, your expression should reduce to the familiar one derived in class.

For Bose statistics, we have in the uncondensed phase,

$$\begin{split} n &= \int_{-\infty}^{\infty} d\varepsilon \, \frac{g(\varepsilon)}{e^{(\varepsilon-\mu)/k_{\rm B}T} - 1} \\ &= {\rm Li}_{3/2} \big(e^{(\mu+\Delta)/k_{\rm B}T} \big) \, \lambda_T^{-3} + {\rm Li}_{3/2} \big(e^{(\mu-\Delta)/k_{\rm B}T} \big) \, \lambda_T^{-3} \end{split}$$

In the condensed phase, $\mu = -\Delta - O(N^{-1})$ is pinned just below the lowest single particle energy, which occurs for $\mathbf{k} = 0$ and $\sigma = -1$. We then have

$$n = n_0 + \zeta(3/2) \, \lambda_T^{-3} + \mathrm{Li}_{3/2} \bigl(e^{-2\Delta/k_\mathrm{B}T} \bigr) \, \lambda_T^{-3} \quad . \label{eq:n_alpha}$$

To find the critical temperature, set $n_0 = 0$ and $\mu = -\Delta$:

$$n\,\lambda_{T_{\rm c}}^3 = \zeta(3/2) + {\rm Li}_{3/2} \bigl(e^{-2\Delta/k_{\rm B}T_{\rm c}} \bigr) \quad . \label{eq:Linear}$$

This is a nonlinear and implicit equation for $T_c(n, \Delta)$. When $\Delta = \infty$, we have

$$k_{\rm B}T_{\rm c}(n,\infty) = \frac{2\pi\hbar^2}{m} \left(\frac{n}{\zeta(3/2)}\right)^{\!\!2/3} \quad . \label{eq:kBTc}$$

(c) When $\Delta = \infty$, the condensation temperature should agree with the familiar result for three-dimensional Bose condensation. Assuming $\Delta \gg k_{\rm B}T_{\rm c}(n, \Delta = \infty)$, find analytically the leading order difference $\delta T_{\rm c}(n, \Delta) \equiv T_{\rm c}(n, \Delta) - T_{\rm c}(n, \Delta = \infty)$.

For finite Δ , we still have the implicit nonlinear equation to solve, but in the limit $\Delta \gg k_{\rm B}T_{\rm c}$, we can expand $T_{\rm c}(n,\Delta) = T_{\rm c}(n,\infty) + \delta T_{\rm c}(n,\Delta)$. We may then set $T_{\rm c}(n,\Delta)$ to $T_{\rm c}(n,\infty)$ in the second term of our nonlinear implicit equation, move this term to the LHS, whence

$$\left(\frac{T_{\rm c}(n,\Delta)}{T_{\rm c}(n,\infty)}\right)^{3/2} = 1 - \frac{1}{\zeta(3/2)} \operatorname{Li}_{3/2} \left(e^{-2\Delta/k_{\rm B}T_{\rm c}(n,\infty)} \right) \quad ,$$

which is a simple algebraic equation for $T_{\rm c}(n, \Delta)$. The second term on the RHS is tiny since $\Delta \gg k_{\rm \scriptscriptstyle B} T_{\rm c}(n, \infty)$. We then find

$$T_{\rm c}(n,\Delta) = T_{\rm c}(n,\infty) \left\{ 1 - \frac{2}{3\zeta(3/2)} e^{-2\Delta/k_{\rm B}T_{\rm c}(n,\infty)} + \mathcal{O}\left(e^{-4\Delta/k_{\rm B}T_{\rm c}(n,\infty)}\right) \right\} \quad,$$

and thus the shift in the condensation temperature is

$$\delta T_{\rm c}(n,\Delta) = -\frac{2}{3\,\zeta(3/2)}\,e^{-2\Delta/k_{\rm B}T_{\rm c}(n,\infty)}\,T_{\rm c}(n,\infty) \quad . \label{eq:deltaTc}$$

Note that

$$\frac{2\Delta}{k_{\rm\scriptscriptstyle B}T_{\rm c}(n,\infty)} = \frac{m\Delta}{\pi\hbar^2} \bigg(\frac{\zeta(3/2)}{n}\bigg)^{2/3} \label{eq:kappa}$$