

Problem 1

electron has kinetic energy  $511,000 \text{ eV} = m_e c^2$

$$K = (\gamma - 1) m_e c^2 \Rightarrow m_e c^2 \Rightarrow \gamma m_e c^2 = 2 m_e c^2 \Rightarrow$$

$$\Rightarrow \boxed{\gamma = 2} \quad \gamma = \frac{1}{\sqrt{1 - \frac{v^2}{c^2}}} \Rightarrow \frac{1}{\gamma^2} = 1 - \frac{v^2}{c^2} \Rightarrow \frac{v^2}{c^2} = 1 - \frac{1}{\gamma^2}$$

$$\Rightarrow \frac{v^2}{c^2} = 1 - \frac{1}{4} = \frac{3}{4} \Rightarrow \boxed{\frac{v}{c} = \frac{\sqrt{3}}{2} = 0.866} \quad (a)$$

$$p = \gamma m_e v \Rightarrow p = \gamma m_e c^2 \frac{v}{c} \cdot \frac{1}{c} = 2 \times 511,000 \text{ eV} \times \frac{\sqrt{3}}{2} \times \frac{1}{c}$$

$$\Rightarrow \boxed{p = 885,078 \text{ eV}/c} \quad (b)$$

(c) Let  $d$  be the distance earth-sun. Time that the light takes is  $\Delta t_{\text{light}} = \frac{d}{c} = 500 \text{ s}$

Time that electron takes is  $\Delta t = \frac{d}{v} = \frac{d}{c} \cdot \frac{c}{v} = 577 \text{ s}$  in earth's frame

In the rest frame of the electron  $\Delta t$  is the proper time, i.e.

$$\Delta t_p = \frac{\Delta t}{\gamma} = \frac{d}{v} \frac{1}{\gamma} = \frac{d}{c} \cdot \frac{c}{v} \cdot \frac{1}{\gamma} \Rightarrow$$

$$\Rightarrow \Delta t_p = \Delta t_{\text{light}} \cdot \frac{2}{\sqrt{3}} \cdot \frac{1}{2} = \frac{\Delta t_{\text{light}}}{\sqrt{3}}$$

$$\Rightarrow \boxed{\Delta t_p = 288.7 \text{ s}} \quad (c)$$

## Problem 2

Wien's law,  $\lambda_m T = \frac{hc}{4.9648}$ .  $T = 300 \text{ K} \Rightarrow$

$$\Rightarrow \lambda_m = \frac{12,400 \times 11,600 \text{ \AA}}{4.96 \times 300} \Rightarrow \boxed{\lambda_m^{\text{emit}} = 96,667 \text{ \AA}} \quad (a)$$

(b) The earth absorbs maximum radiation at wavelength where the suns emits it, i.e. same formula as above with

$$T = 5800 \text{ K} \Rightarrow \boxed{\lambda_m^{\text{absorb}} = \lambda_m^{\text{emit}} \times \frac{300}{5800} \text{ \AA} = 5000 \text{ \AA}} \quad (b)$$

(c) The total radiation emitted by the earth is the same as what is incident on the earth. So

$$P_{\text{tot}} = \sigma T^4 \times 4\pi R_e^2$$

$$P_{\text{tot}} = 5.67 \times 10^{-8} \times 300^4 \times 4\pi \times 6378^2 \times 10^6 \text{ W}$$

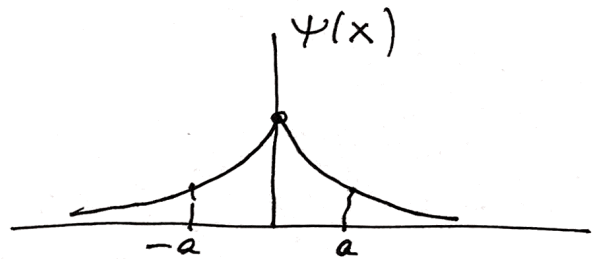
$$\boxed{P_{\text{tot}} = 2.35 \times 10^{17} \text{ W}}$$

### Problem 3

$\Psi(x) = C e^{-|x|/a}$ . Use that  $\Psi$  is even.

$$\frac{1}{2} = \int_0^{\infty} dx |\Psi(x)|^2 = C^2 \int_0^{\infty} dx e^{-2x/a} = -C^2 \frac{a}{2} e^{-2x/a} \Big|_0^{\infty} = C^2 \frac{a}{2}$$

$$\Rightarrow \frac{1}{2} = C^2 \frac{a}{2} \Rightarrow \boxed{C = \frac{1}{\sqrt{a}}} \quad (a)$$



$$(b) \Delta x = \sqrt{\langle x^2 \rangle - \langle x \rangle^2}$$

$\langle x \rangle = 0$  by symmetry

$$\langle x^2 \rangle = \frac{2}{a} \int_0^{\infty} dx x^2 |\Psi(x)|^2 = \frac{2}{a} \int_0^{\infty} dx x^2 e^{-2x/a}$$

Use formula  $\int_0^{\infty} dx x^n e^{-\lambda x} = \frac{n!}{\lambda^{n+1}} \Rightarrow$

$$\Rightarrow \langle x^2 \rangle = \frac{2}{a} \cdot \frac{2!}{\left(\frac{2}{a}\right)^3} = \frac{2a^3}{a} \cdot \frac{2}{2^3} = \frac{a^2}{2}$$

$$\Rightarrow \boxed{\Delta x = \frac{a}{\sqrt{2}} = 0.707a} \quad (b)$$

$$(c) P(|x| < a) = 2 \int_0^a dx |\Psi(x)|^2 = 2 \int_0^a dx \frac{1}{a} e^{-2x/a} =$$
$$= \frac{2}{a} \left( -\frac{a}{2} \right) e^{-2x/a} \Big|_0^a = 1 - e^{-2} = 0.865$$

$$\boxed{P(|x| < a) = 0.865} \quad (c)$$

## Problem 4

energies for harmonic oscillator are

$$E_n = \hbar\omega \left( n + \frac{1}{2} \right)$$

$n$  has to be  $\pm$  for absorption or emission. So

$$\hbar\omega = \frac{hc}{\lambda} = \frac{12,400}{5000} \text{ eV} = 2.48 \text{ eV}$$

$\Rightarrow$  ground state energy  $\boxed{E_0 = \frac{\hbar\omega}{2} = 1.24 \text{ eV}} \quad (a)$

(b) Classical amplitude  $A$  obeys

$$\frac{1}{2} m_e \omega^2 A^2 = \frac{\hbar\omega}{2} \Rightarrow A^2 = \frac{\hbar}{m_e \omega} = \frac{\hbar^2}{m_e \hbar\omega}$$

using  $\frac{\hbar^2}{m_e} = 7.62 \text{ eV} \text{ \AA}^2 \Rightarrow \boxed{A = \sqrt{\frac{7.62}{2.48}} \text{ \AA} = 1.75 \text{ \AA}} \quad (b)$

(c) Wave function for ground state is

$$\psi(x) = C e^{-\frac{m\omega}{2\hbar} x^2} = C e^{-\frac{x^2}{2A^2}}$$

$$\frac{|\psi(x=A)|^2}{|\psi(0)|^2} = e^{-1} = 0.368 \Rightarrow$$

$$\boxed{\frac{|\psi(0)|^2}{|\psi(A)|^2} = 2.72 \text{ times more likely}} \quad (c)$$

## Problem 5

$$(a) E_1 = \frac{\hbar^2 \pi^2}{2m_e L^2} = \frac{3.81 \pi^2}{4^2} eV = 2.35 eV$$

$$\Rightarrow \boxed{E_1 = 2.35 eV \text{ assuming } \infty \text{ well}} \quad (a)$$

(b) Wavefunction penetrates on right side  
- a distance  $\delta = 1/\alpha$

$$\alpha = \sqrt{\frac{2m}{\hbar^2} (U - E)} = \sqrt{\frac{1}{3.81} (20 - 2.35) \text{ \AA}^{-1}} =$$

$$= 2.152 \text{ \AA}^{-1} \Rightarrow \boxed{\delta = 0.465 \text{ \AA}}$$

so the effective length of the well is

$$L_{\text{eff}} = L + \delta = 4.465 \text{ \AA}$$

so a better estimate for the ground state energy is

$$E_1' = \frac{\hbar^2 \pi^2}{2m_e L_{\text{eff}}^2} = 2.35 \times \left(\frac{4}{4.465}\right)^2 eV \Rightarrow$$

$$\boxed{E_1' = 1.89 eV}$$

(c) Transmission coefficient

$$T = e^{-2 \sqrt{\frac{2m}{\hbar^2} (U - E_1')} \cdot 1 \text{ \AA}} = e^{-4.36} = 0.013$$

$$\Rightarrow \boxed{\text{probability of escape in 1 attempt is } 1.3\%}$$



## Problem 6

$$E_{n_1, n_2} = \frac{\hbar^2 \pi^2}{2m_e L^2} (n_1^2 + n_2^2) \equiv E_0 (n_1^2 + n_2^2)$$

There can be 2 electrons in each state

Largest wavelength photon is

absorbed for transition with smallest  $\Delta E$

That is the transition  $(2,2) \rightarrow (1,3)$  or  $(3,1)$

$$\Delta E = E_0 (10 - 8) = 2E_0 = \frac{hc}{\lambda}$$

$$E_0 = \frac{\hbar^2 \pi^2}{2m_e L^2} = \frac{3.81 \pi^2}{25} \text{ eV} = 1.509 \text{ eV}$$

$$\Rightarrow \lambda = \frac{12,400 \text{ \AA}}{3.008} = 4122 \text{ \AA} \quad (a)$$

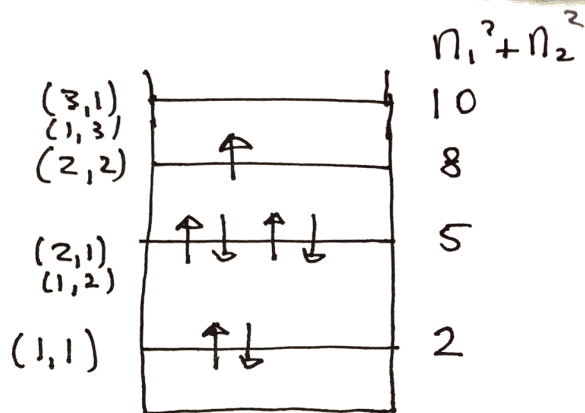
(b) Next largest:  $(2,1) \rightarrow (2,2)$

$$\Delta E = E_0 (8 - 5) = 3E_0 = \frac{hc}{\lambda}$$

$$\Rightarrow \lambda = 2748 \text{ \AA} \quad (b)$$

(c)  $E_{\text{first state}} = E_0 (2 \times 2 + 4 \times 5 + 1 \times 8) = 32 E_0$

$$\Rightarrow E_{\text{first}} = 48.13 \text{ eV}$$



## Problem 7

$L = 4\hbar$  acc. to Bohr  $\Rightarrow n = 4$

(a) de Broglie wavelength is  $\lambda = \frac{h}{p} = \frac{h r}{p r}$

since  $L = p r$  is the angular momentum, and  $\hbar = h/2\pi$

$$\lambda = \frac{h r}{4\hbar} = \frac{2\pi r}{4} = \frac{\pi r}{2}. \quad \text{Since } r = r_n = a_0 n^2 \text{ with}$$

$$a_0 = 0.528 \text{ \AA} \text{ and } n = 4 \Rightarrow \lambda = \frac{\pi}{2} a_0 \cdot 16 = 8\pi a_0 = 2\pi a_0 \cdot 4$$

$$\Rightarrow \boxed{\lambda = 13.30 \text{ \AA}} \quad (a)$$

(b) Fr Schrödinger,  $L = \sqrt{l(l+1)}\hbar$ . Since  $n = 4$ , maximum allowed  $l$  is  $l = 3$ , so

$$\boxed{L = \sqrt{12}\hbar = 3.46\hbar} \quad (b)$$

Schrödinger knows that  $n = 4$  since he measured the electron's energy

$$\boxed{E_4 = -\frac{E_0}{4^2} = -0.85 \text{ eV}} \quad \text{with } E_0 = 13.6 \text{ eV} \quad (b)$$

(c) Bohr says the most likely value for the distance is  $r_n = a_0 n^2$ ,

$$\text{so } \boxed{r_4 = 8.46 \text{ \AA}}$$

Schrödinger says: If  $l = 3$ , which gives  $L$  closest to Bohr's value and

largest  $r \Rightarrow R(r) \propto r^3 e^{-r/4a_0} \Rightarrow P(r) \propto r^8 e^{-r/2a_0}$

Most likely  $r$  is for  $P'(r) = 0 \Rightarrow$

$$8r^7 - \frac{r^8}{2a_0} = 0 \Rightarrow \boxed{r = 16a_0 = 8.46 \text{ \AA}} \quad \text{agrees with Bohr's.}$$

## Problem 8

$$n=3, l=2, m_l=-1$$

For  $n=2$ ,  $l=1$  or  $l=0$ . Selection rule is  $\Delta l = \pm 1$ ,  
and  $\Delta m_l = 0, \pm 1$ . So  $l=1$  for final state, and  
 $m_l = -1$  or  $m_l = 0$  for final state

$$\Delta E = \frac{hc}{\lambda} = E_0 \left( \frac{1}{4} - \frac{1}{9} \right) = \frac{5}{36} E_0 \Rightarrow$$

$$\lambda = \frac{36}{5} \times \frac{12,400}{13.6} \text{ \AA} \Rightarrow \boxed{\lambda = 6565 \text{ \AA}} \quad (a)$$

Possible final states:

$$\boxed{\begin{array}{l} n=2, l=1, m_l=-1 \\ n=2, l=1, m_l=0 \end{array}} \quad (a)$$

(b) In a magnetic field, energy shifts by

$$\Delta E_B = +\mu_B B m_l$$

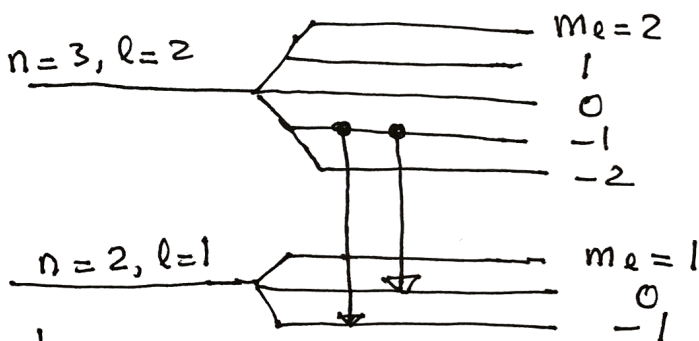
(i) For  $m_l = -1 \rightarrow m_l = -1$ ,  
Same as before:  $\lambda = 6565 \text{ \AA}$

(ii) For  $m_l = -1 \rightarrow m_l = 0$

$$\Delta E = E_0 \left( \frac{1}{4} - \frac{1}{9} \right) - \mu_B B = \frac{hc}{\lambda'}$$

$$= \frac{5}{36} E_0 - 5.79 \times 10^{-5} \times 50 \text{ eV} = 1.888 \text{ eV} - 2.895 \times 10^{-3} \text{ eV} = 1.886 \text{ eV}$$

$$\Rightarrow \boxed{\lambda' = 6575 \text{ \AA}} \quad (b)$$



(c) (i) There are 2 states with  $m_l = -1$ :  $l=1, m_l = -1$  and  $l=2, m_l = -1$   
There is 1 state with  $m_l = -2$ . So  $\boxed{\text{prob}(m_l = -1) / \text{prob}(m_l = -2) = 2}$

With a magnetic field, energy of  $m_l = -2$  state is lower  $\Rightarrow$  becomes more probable

$$\boxed{\frac{\text{prob}(m_l = -1)}{\text{prob}(m_l = -2)} = 2 e^{-\Delta E_B / k_B T} = 2 e^{-0.112} = 1.79}$$



### Problem 9

$$\Psi(x) = \sqrt{\frac{2}{L}} \sin\left(\frac{\pi x}{L}\right)$$

$$\langle Q \rangle = \int dx \Psi^*(x) [Q] \Psi(x)$$

$$\Delta Q = \sqrt{\langle Q^2 \rangle - \langle Q \rangle^2}$$

$$[P] = \frac{\hbar}{i} \frac{d}{dx} \quad ; \quad [P^2] = -\hbar^2 \frac{d^2}{dx^2}$$

$$(a) \quad [P] \Psi(x) = \frac{\pi}{L} \sqrt{\frac{2}{L}} \cos\left(\frac{\pi x}{L}\right) \neq q \Psi(x)$$

$\Rightarrow$   $[P]$  is not sharp, it is fuzzy  $\Rightarrow$  no eigenvalue

$$[P^2] \Psi(x) = \frac{\hbar^2 \pi^2}{L^2} \Psi(x) \Rightarrow$$

$[P^2]$  is sharp. Its eigenvalue is  $\frac{\hbar^2 \pi^2}{L^2}$

(b) Since  $[Q] = [P^2]$  is a sharp observable,

$$\langle P^2 \rangle = \frac{\hbar^2 \pi^2}{L^2}, \quad \Delta(P^2) = 0$$

For  $[Q] = [P]$  ,

$\langle P \rangle = 0$  by symmetry

$$\Delta P = \sqrt{\langle P^2 \rangle - \langle P \rangle^2} = \frac{\hbar \pi}{L}$$