

- 2-2 (a) Scalar equations can be considered in this case because relativistic and classical velocities are in the same direction.

$$p = \gamma mv = 1.90mv = \frac{mv}{\left[1 - (v/c)^2\right]^{1/2}} \Rightarrow \frac{1}{\left[1 - (v/c)^2\right]^{1/2}} = 1.90 \Rightarrow v = \left[1 - \left(\frac{1}{1.90}\right)^2\right]^{1/2} c$$

$$= 0.85c$$

- (b) No change, because the masses cancel each other.

- 2-9 (a) When $K = (\gamma - 1)mc^2 = 5mc^2$, $\gamma = 6$ and $E = \gamma mc^2 = 6(0.5110 \text{ MeV}) = 3.07 \text{ MeV}$.

(b) $\frac{1}{\gamma} = \left[1 - \left(\frac{v}{c}\right)^2\right]^{1/2}$ and $v = c \left[1 - \left(\frac{1}{\gamma}\right)^2\right]^{1/2} = c \left[1 - \left(\frac{1}{6}\right)^2\right]^{1/2} = 0.986c$

- 2-12 (a) When $K_e = K_p$, $m_e c^2 (\gamma_e - 1) = m_p c^2 (\gamma_p - 1)$. In this case $m_e c^2 = 0.5110 \text{ MeV}$ and $m_p c^2 = 938 \text{ MeV}$, $\gamma_e = \left[1 - (0.75)^2\right]^{1/2} = 1.5119$. Substituting these values into the first equation, we find $\gamma_p = 1 + \frac{m_e c^2 (\gamma_e - 1)}{m_p c^2} = 1 + \frac{(0.5110)(1.5119 - 1)}{939} = 1.000279$.

But $\gamma_p = \frac{1}{\left[1 - (u_p/c)^2\right]^{1/2}}$; therefore $u_p = c(1 - \gamma_p^{-2})^{1/2} = 0.0236c$.

- (b) When $p_e = p_p$, $\gamma_p m_p u_p = \gamma_e m_e u_e$ or $u_p = \left(\frac{\gamma_e}{\gamma_p}\right) \left(\frac{m_e}{m_p}\right) u_e$,

$$u_p = \left(\frac{1.5119}{1.000279}\right) \left[\frac{0.5110/c^2}{939/c^2}\right] (0.75c) = 6.17 \times 10^{-4} c.$$

- 2-13 (a) $E = 400mc^2 = \gamma mc^2$

$$\gamma = \left(1 - \frac{v^2}{c^2}\right)^{-1/2} = 400$$

$$\left(1 - \frac{v^2}{c^2}\right) = \left(\frac{1}{400}\right)^2$$

$$\frac{v}{c} = \left[1 - \frac{1}{400^2}\right]^{1/2}$$

$$v = 0.999997c$$

- (b) $K = E - mc^2 = (400 - 1)mc^2 = 399mc^2 = (399)(938.3 \text{ MeV}) = 3.744 \times 10^5 \text{ MeV}$

- 2-15 (a) $K = \gamma mc^2 - mc^2 = Vq$ and so, $\gamma^2 = \left(1 + \frac{Vq}{mc^2}\right)^2$ and $\frac{v}{c} = \left\{1 - \left(1 + \frac{Vq}{mc^2}\right)^{-2}\right\}^{1/2}$

$$\frac{v}{c} = \left\{ 1 - \frac{1}{1 + (5.0 \times 10^4 \text{ eV}/0.511 \text{ MeV})^2} \right\}^{1/2} = 0.4127$$

or $v = 0.413c$.

$$(b) \quad K = \frac{1}{2}mv^2 = Vq$$

$$v = \left(\frac{2Vq}{m} \right)^{1/2} = \left\{ \frac{2(5.0 \times 10^4 \text{ eV})}{0.511 \text{ MeV}/c^2} \right\}^{1/2} = 0.442c$$

(c) The error in using the classical expression is approximately $\frac{3}{40} \times 100\%$ or about 7.5%

$$2-19 \quad \Delta m = 6m_p + 6m_n - m_C = [6(1.007276) + 6(1.008665) - 12] \text{ u} = 0.095646 \text{ u},$$

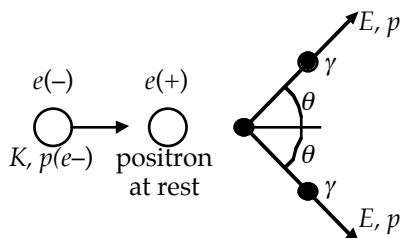
$$\Delta E = (931.49 \text{ MeV/u})(0.095646 \text{ u}) = 89.09 \text{ MeV}.$$

Therefore the energy per nucleon = $\frac{89.09 \text{ MeV}}{12} = 7.42 \text{ MeV}$.

$$2-20 \quad \Delta m = m - m_p - m_e = 1.008665 \text{ u} - 1.007276 \text{ u} - 0.0005485 \text{ u} = 8.404 \times 10^{-4} \text{ u}$$

$$E = c^2(8.404 \times 10^{-4} \text{ u}) = (8.404 \times 10^{-4} \text{ u})(931.5 \text{ MeV/u}) = 0.783 \text{ MeV}.$$

2-21



Conservation of mass-energy requires $K + 2mc^2 = 2E$ where K is the electron's kinetic energy, m is the electron's mass, and E is the gamma ray's energy.

$$E = \frac{K}{2} + mc^2 = (0.500 + 0.511) \text{ MeV} = 1.011 \text{ MeV}.$$

Conservation of momentum requires that $p_{e^-} = 2p \cos \theta$ where p_{e^-} is the initial momentum of the electron and p is the gamma ray's momentum, $\frac{E}{c} = 1.011 \text{ MeV}/c$. Using $E_{e^-}^2 = p_{e^-}^2 c^2 + (mc^2)^2$ where E_{e^-} is the electron's total energy, $E_{e^-} = K + mc^2$, yields

$$p_{e^-} = \frac{1}{c} \sqrt{K^2 + 2Kmc^2} = \frac{\sqrt{(1.00)^2 + 2(1.00)(0.511)} \text{ MeV}}{c} = 1.422 \text{ MeV}/c.$$

Finally, $\cos \theta = \frac{p_{e^-}}{2p} = 0.703$; $\theta = 45.3^\circ$.

2-23 In this problem, M is the mass of the initial particle, m_l is the mass of the lighter fragment, v_l is the speed of the lighter fragment, m_h is the mass of the heavier fragment, and v_h is the speed of the heavier fragment. Conservation of mass-energy leads to

$$Mc^2 = \frac{m_l c^2}{\sqrt{1 - v_l^2/c^2}} + \frac{m_h c^2}{\sqrt{1 - v_h^2/c^2}}$$

From the conservation of momentum one obtains

$$(m_l)(0.987c)(6.22) = (m_h)(0.868c)(2.01)$$

$$m_l = \frac{(m_h)(0.868c)(2.01)}{(0.987)(6.22)} = 0.284m_h$$

Substituting in this value and numerical quantities in the mass-energy conservation equation, one obtains $3.34 \times 10^{-27} \text{ kg} = 6.22m_l + 2.01m_h$ which in turn gives $3.34 \times 10^{-27} \text{ kg} = (6.22)(0.284)m_l + 2.01m_h$ or $m_h = \frac{3.34 \times 10^{-27} \text{ kg}}{3.78} = 8.84 \times 10^{-28} \text{ kg}$ and $m_l = (0.284)m_h = 2.51 \times 10^{-28} \text{ kg}$.

2-26 Energy conservation: $\frac{1}{\sqrt{1-0^2}} 1400 \text{ kg}c^2 + \frac{900 \text{ kg}c^2}{\sqrt{1-0.85^2}} = \frac{Mc^2}{\sqrt{1-v^2/c^2}}$;

$3108 \text{ kg} \sqrt{1 - \frac{v^2}{c^2}} = M$. Momentum conservation: $0 + \frac{900 \text{ kg}(0.85c)}{\sqrt{1-0.85^2}} = \frac{Mv}{\sqrt{1-v^2/c^2}}$;

$1452 \text{ kg} \sqrt{1 - \frac{v^2}{c^2}} = \frac{Mv}{c}$.

(a) Dividing gives $\frac{v}{c} = \frac{1452}{3108} = 0.467$ $v = 0.467c$.

(b) Now by substitution $3108 \text{ kg} \sqrt{1 - 0.467^2} = M = 2.75 \times 10^3 \text{ kg}$.

2-31 Conservation of momentum γmu :

$$\frac{mu}{\sqrt{1-u^2/c^2}} + \frac{m(-u)}{3\sqrt{1-u^2/c^2}} = \frac{Mv_f}{\sqrt{1-v_f^2/c^2}} = \frac{2mu}{3\sqrt{1-u^2/c^2}}.$$

Conservation of energy γmc^2 :

$$\frac{mc^2}{\sqrt{1-u^2/c^2}} + \frac{mc^2}{3\sqrt{1-u^2/c^2}} = \frac{Mc^2}{\sqrt{1-v_f^2/c^2}} = \frac{4mc^2}{3\sqrt{1-u^2/c^2}}.$$

To start solving we can divide: $v_f = \frac{2u}{4} = \frac{u}{2}$. Then

$$\begin{aligned} \frac{M}{\sqrt{1-u^2/4c^2}} &= \frac{4m}{3\sqrt{1-u^2/c^2}} = \frac{M}{(1/2)\sqrt{4-u^2/c^2}} \\ M &= \frac{2m\sqrt{4-u^2/c^2}}{3\sqrt{1-u^2/c^2}} \end{aligned}$$

Note that when $v \ll c$, this reduces to $M = \frac{4m}{3}$, in agreement with the classical result.

2-33 The energy that arrives in one year is

$$E = \mathcal{P} \Delta t = (1.79 \times 10^{17} \text{ J/s})(3.16 \times 10^7 \text{ s}) = 5.66 \times 10^{24} \text{ J}.$$

$$\text{Thus, } m = \frac{E}{c^2} = \frac{5.66 \times 10^{24} \text{ J}}{(3.00 \times 10^8 \text{ m/s})^2} = 6.28 \times 10^7 \text{ kg}.$$

2-8 (a) $E_R = mc^2 = (1.67 \times 10^{-27} \text{ kg})(3 \times 10^8 \text{ m/s})^2 = 1.503 \times 10^{-10} \text{ J} = 939.4 \text{ MeV}$
 (Numerical round off gives a slightly larger value for the proton mass)

(b) $E = \gamma mc^2 = \frac{1.503 \times 10^{-10} \text{ J}}{(1 - (0.95c/c)^2)^{1/2}} = 4.813 \times 10^{-10} \text{ J} \approx 3.01 \times 10^3 \text{ MeV}$

(c) $K = E - mc^2 = 4.813 \times 10^{-10} \text{ J} - 1.503 \times 10^{-10} \text{ J} = 3.31 \times 10^{-10} \text{ J} = 2.07 \times 10^3 \text{ MeV}$

2-10 $E = \gamma mc^2$; $1.5mc^2 = \gamma mc^2$; $\gamma = 1.5 \Rightarrow 1.5 = \frac{1}{[1 - (v^2/c^2)]^{1/2}}$; $v = c \left[1 - \left(\frac{1}{1.5} \right)^2 \right]^{1/2} = 0.75c$

2-11 (a) $K = 50 \times 10^9 \text{ eV}$; $mc^2 = 938.27 \text{ MeV}$;

$$E = K + mc^2 = (50 \times 10^9 \text{ eV}) + (938.27 \times 10^6 \text{ eV}) = 50\,938.3 \text{ MeV}$$

$$E^2 = p^2 c^2 + m^2 c^4 \Rightarrow p = \left[\frac{E^2 - m^2 c^4}{c^2} \right]^{1/2}$$

$$p = \frac{\left[(50\,938.3 \text{ MeV})^2 - (938.27 \text{ MeV})^2 \right]^{1/2}}{c} = 5.09 \times 10^{10} \text{ eV}/c$$

$$= \frac{5.09 \times 10^{10} \text{ eV}}{3 \times 10^8 \text{ m/s}} (1.6 \times 10^{-19} \text{ J/eV}) = 2.71 \times 10^{-17} \text{ kg} \cdot \text{m/s}$$

(b) $E = \gamma mc^2 = \frac{mc^2}{[1 - (v/c)^2]^{1/2}} \Rightarrow v = c \left[1 - \left(\frac{mc^2}{E} \right)^2 \right]^{1/2}$

$$= (3 \times 10^8 \text{ m/s}) \left[1 - \left(\frac{938.27 \text{ MeV}}{50\,938.3 \text{ MeV}} \right)^2 \right]^{1/2} = 2.9995 \times 10^8 \text{ m/s}$$

2-17 $\Delta m = m_{\text{Ra}} - m_{\text{Rn}} - m_{\text{He}}$ (an atomic unit of mass, the u, is one-twelfth the mass of the ^{12}C atom or $1.66054 \times 10^{-27} \text{ kg}$)

$$\Delta m = (226.0254 - 220.175 - 4.0026) \text{ u} = 0.0053 \text{ u}$$

$$E = (\Delta m)(931 \text{ MeV/u}) = (0.0053 \text{ u})(931 \text{ MeV/u}) = 4.9 \text{ MeV}$$