

Physics 235

Lecture II

Stochastic Fields:  
Kubo #, Collisionless and  
Collisional Heat Transport

Recall:

→ Chirikov criterion for onset  
Hamiltonian chaos

→  $\kappa_u = \sigma_r / \Delta_r \sim \rho_{au} \frac{dB}{B_0} / \Delta_r$

(→  $\nabla T_{au} / \Delta$ )

→  $\kappa_u \ll 1$

$$D_M = \sum_{\underline{u}} \left( \frac{\rho_{Bu} r}{B_0} \right)^2 \pi \delta(\kappa_u)$$
$$= \left( \sum_{\underline{u}} \left( \frac{\rho_{Bu} r}{B_0} \right)^2 \right) \rho_{au}$$

More on Kubo #

For radial excursion!

$$dr/dz = \tilde{B}_r/B_0$$

so  $dr \approx \int_0^l (\tilde{B}_r/B_0) dz$

Now, /no trajectory de-coheres from perturbation for  $l > l_{ac}$

What do the symbols mean?  $\rightarrow$  autocorrelation length

$l_{ac} \approx 1/|\Delta(\omega_r)|$  i.e. inverse spectral bandwidth

$\left\{ dr \approx l_{ac} \tilde{B}_r/B_0 \right\} \rightarrow \left\{ \begin{array}{l} \text{time excursion of} \\ \text{I } l_{ac} \end{array} \right.$

Can identify  $\Delta_r \equiv$  scatterer radial correlation length (i.e. spatial spectral width)

then.

$Ku \approx dr/\Delta_r \approx \frac{l_{ac}}{\Delta_r} \tilde{B}_r/B_0 \rightarrow$  Kubo #

and can then post:

$\rightarrow Ku < 1 \Rightarrow$  many kicks in coherence length  $\Rightarrow$  diffusion process

$k_{cu} \sim 1 \rightarrow$  B.B.K. "natural state" of EM turbulence  
 $k_{cu} \sim 1 \rightarrow$  critical balance.

$\rightarrow k_{cu} > 1 \rightarrow$  more than one  $\Delta_n$  in  $k_{cu} l$   
 $\rightarrow$  strong scattering  $\leftrightarrow$  decorrelation.

QLT

Here  $k_{cu} \leq 1$ , at first. So, proceed via Quasilinear theory.

$$\Gamma_n = \left\langle \frac{\tilde{B}_r}{B_0} \tilde{F} \right\rangle$$

$$= \sum_n \frac{\tilde{B}_{r-n}}{B_0} \tilde{F}_n$$

$$-i \left( k_z - k_0 \frac{B_0}{B_0} \right) \tilde{F}_n = -\tilde{B}_{r-n} \frac{\partial \langle f \rangle}{\partial v}$$

So

$$\Gamma_n = -D_n \frac{\partial \langle f \rangle}{\partial v}$$

$$D_n = \sum_n \left| \frac{\tilde{B}_{r-n}}{B_0} \right|^2 \pi \delta \left( k_z - k_0 \frac{B_0}{B_0} \right)$$

↓  
magnetic diffusivity

$$= \sum_n \left| \frac{\tilde{B}_{r-n}}{B_0} \right|^2 \pi \delta(k_{cu})$$

(RSTZ EG)

$$\approx \left\langle \left( \frac{\partial B_r}{B_0} \right)^2 \right\rangle_{lac}$$

What is lac?

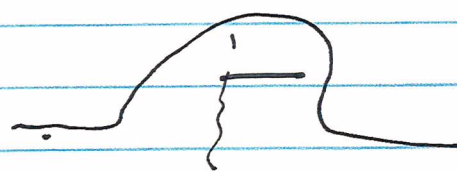
lec

N.B.:  $\sum_n = \sum_{m,n}$  spatial spread  
scatters  $\rightarrow$  lec

$$n = \frac{m}{\lambda}, \quad dn = \frac{dm}{\lambda} \approx \frac{1}{\lambda} dx$$

$\Rightarrow$  spatial scale of spectral width ( $\Delta \nu$ )  
sets  $|k_{\perp}| \sim \frac{|k_0 \Delta \nu|}{L_0}$

$\lambda_{av} \sim L_0 / |k_0 \Delta \nu|$



Lines then diffuse as:

$$\langle \Delta \nu^2 \rangle \sim D_M Z$$

Broaden Alternative  $\rightarrow$  Orbit averaging

N.B. Line Liouville eqn can be obtained  
by reducing / simplify in  $\Delta K E$

$$\frac{\partial f}{\partial t} + v_{||} \hat{n}_0 \cdot \nabla f + \frac{v_{\perp}}{B} \cdot \nabla f - \frac{c}{B} \nabla \phi \times \hat{z} \cdot \nabla f$$

$$+ v_{||} \frac{dB_{||}}{B_0} \cdot \nabla f - \frac{k_{||}}{m_e} E_{||} \frac{\partial f}{\partial v_{||}} = C(f)$$

$$\Rightarrow \Lambda_0 \cdot \nabla F + \frac{dB_0}{B_0} \cdot \nabla F = 0 \quad \checkmark$$

Scales

Now, scales:

$l_{sc}$   $\rightarrow$  (scatters)  
 $\rightarrow$  field line memory length  
 self-coherence of scattering field.

$l_0 \rightarrow$  into decorrelation length  
 (length over which line scattered from its upo)  
 i.e.  $r \frac{d\theta}{dz} = \frac{B_0(r)}{B_0}$

but  $r$  scattered,  $\Rightarrow$

$$\frac{dy}{dz} = B_0(r_0) + \frac{B_0'(r_0)}{B_0} dr$$

$$\frac{d}{dz} \frac{dy}{dz} \approx \frac{B_0''(r_0)}{B_0} dr$$

$$\langle dy^2 \rangle = \left\langle \left( \int \left( \frac{B_0''}{B_0} \right) dr dz \right)^2 \right\rangle$$

⇒

$$\langle dy^2 \rangle \sim \frac{B_0^{1/2}}{B_0^2} Z^2 \langle dr^2 \rangle$$

$$\sim \frac{B_0^{1/2}}{3B_0^2} D_M Z^3$$

akin

$$\langle dx^2 \rangle \sim D_M \sqrt{\frac{2}{3}} \quad \text{or } 10$$

For orbit deformation length:

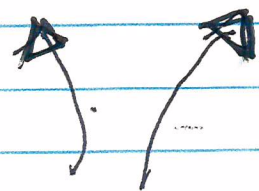
$$k_0^2 \langle dy^2 \rangle \sim k_0^2 \frac{B_0^{1/2}}{3B_0^2} D_M Z^3$$

⇒

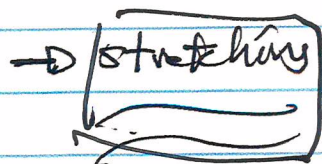
$$k_0 \sim \left( \frac{k_0^2 B_0^{1/2} D_M}{3B_0^2} \right)^{-1/3}$$

$$\sim \left( \frac{k_0^2 D_M}{L_s^2 \frac{1}{3}} \right)^{-1/3}$$

analogous  
to shear  
dispersion  
 $\left( \frac{k_0^2 v^2}{3} \right)^{1/3}$

Also: 

orbit exponentiation  
length  
(separation)



show via 2pt.  $\langle \sigma(t) \sigma(t') \rangle$

chaotic stretching

→ For QL regime validity:

$l_{ac} < l_c$

→  $l_{ac} < l_c$  (show!)  
 $l_{ac} < l_c \rightarrow ?$  (prop.)

and another (artificial) length:  $l_{mp}$ .

⇒  $l_{ac} < l_c < l_{mp}$  → so called "collisionless regime"  
 $l_{ac} < l_{mp} < l_c$  → collisional

which brings us to: something physical

### (Electron) Heat Transport

Theme: - waves  
- interaction processes

N.B. - nobody cares about "line" diffusion

why do this?

- people (i.e. experimentalists) do care about:

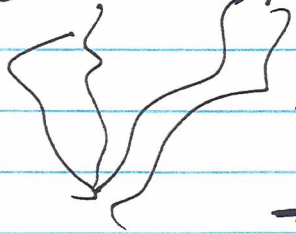
- heat
  - particle
  - momentum
- } transport.

Rechester & Rosenbluth  
PR L'78  
a MUST!

∴ lets begin with heat transport!

→ Consider  $l_{ac} < l_c < l_{imp}$  : heat diffusivity

- lines wander



What is  $v_{\perp}$ ?  
 ⇒ "of course etc"  
 $x_{\perp} \sim v_{\perp} t$   
 ⇒ but is it so simple?

Recall = thought exm.

- but, lets assume parallel collisions (only) happen. (Particle stays on line!).

so motion along line is diffusive

$$\Delta z^2 \sim D_{\parallel} t \sim \frac{v_{\parallel}^2}{v_{\perp}^2} t$$

↓  
parallel thermal diffusion

→ so: fun w/ heat:

$$\langle \Delta r^2 \rangle \sim D_{\perp} z \sim D_{\perp} (v_{\parallel} t)^{1/2}$$

so: radial scatter

$$x_{\perp} \equiv d \langle \Delta r^2 \rangle / dt \sim D_{\perp} (v_{\parallel})^{1/2} / t^{1/2}$$

for



Point: → line may wander

but

→ particle kicked back ~~off~~ along line

→ even though bc  $\ll$   $\lambda_{mp}$ ,

no net radial wander as particle kicked back.

Less on:

→ collisions control crossability

⊗ → need get kicked off field line

→ Need:

- Coarse graining →

- FLR →  $\rho_e$

-  $\lambda_L$

- drifts.

⊗ } minimum resolution scale

⇒ applied every  $\lambda_{mp}$

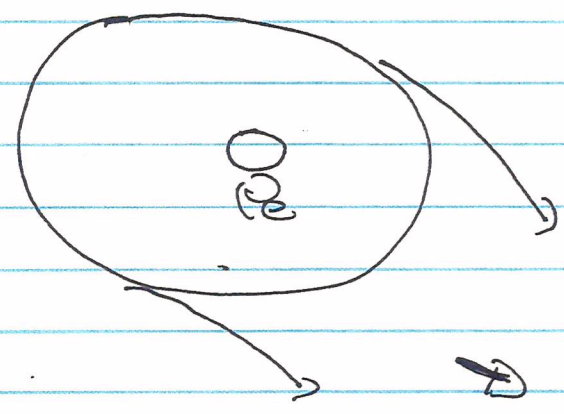
smear particle location over a resolution cell.

$\Rightarrow$  coarse graining reduces "active volume"

so

$\rightarrow$  consider the following argument:

① Consider disk of  $r \sim \rho_0$

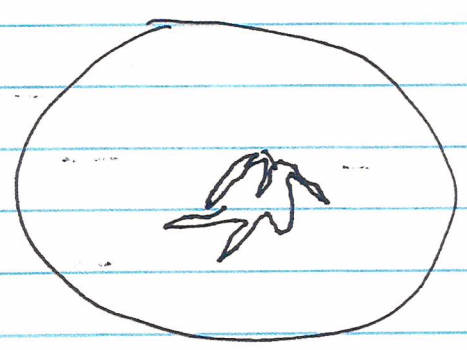


②

Map disk forward, noting that  $\underline{D \cdot B} = 0$   
 $\Rightarrow$  map is area preserving

after  $\sim$   $l_{msf}$

$\pm h_L > 0$   
 $\pm h_L < 0$



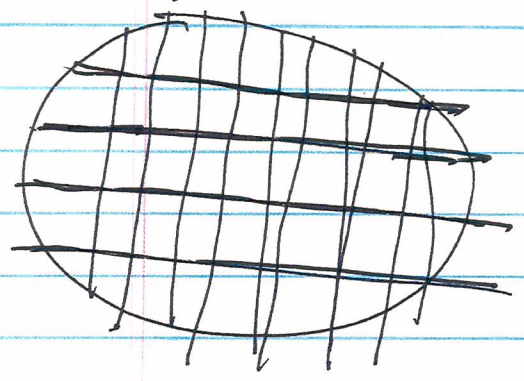
( $h \rightarrow$  Lyapunov Expt)

width

$$\sim w \sim \rho_0 e^{-l_{msf}/l_c}$$

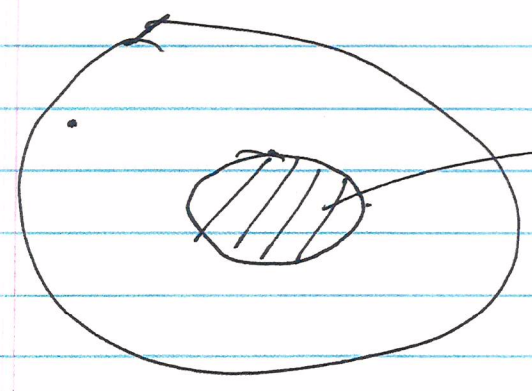
$$(\sim \rho_0 e^{+l_{msf}/l_c})$$

③ but coarse graining occurs at  $l_{mp}$



particle/contour  
"re-sets/smeared"  
to nearest grid  
site

~~50~~  
④



$$A_{co} \vec{F} = A_0 \vec{f}_0$$

( $\vec{F} < f_0$ )

coarse graining  
of structure  
from previous  
V.O.F.

and can continue...

⑤ Ludwig Boltzmann showed us NO  
memory between steps (1  $l_{mp}$  /  
collision time)

so initial spot expands, with  
random walk, as

$$\langle \sigma^2 \rangle \sim D \tau l_{mp} \quad / \quad \text{in } l_{mp}$$

v.e.

coarse graining interval stats  
 $\langle \delta v^2 \rangle$  step!

⇒

⑥ then, for  $\chi_{\perp}$  (Eq 1mp):  $1/\tau_c \sim \nu_c$

$$\chi_{\perp} \sim \langle \delta v^2 \rangle / \nu_c \sim D_M \frac{1mp}{\tau_c}$$

$\sim \text{von DM.}$

⇒

$$\chi_{\perp} \sim \text{von DM}$$

→ collisionless stochastic field next |  
 diffusivity

→ manifestly independent of collisionality

→ yet clearly dependent on  
 collisions and coarse graining

Lesson: Coarse graining essential  
 to irreversibility

~~Collisions → arrow of time~~

Collisions → arrow of time.



on

Coarse graining essential to high  
 particle ~~off~~ field line, on edge  
 collisions back-scatter  
 reversed wendler.

## Stoch Fields, cont'd

### Exercises (suggested) :

- i.) Derive the magnetic diffusivity with magnetic drifts. How do these modify  $D_{\perp}$ ? Explain why high energy particles (runaways) are confined longer than thermal.
- ii.) Formulate the theory of diffusion due stochastic fields in toroidal geometry using ballooning mode formalism for the fluctuations.
- iii.) What happens to net cross field transport in a standing spectrum of e.s. and magnetic perturbations. When might transport vanish? Why?

→ Collisional Regime — More challenging

Here:  $l_{co} < l_{mp} < l_c$   
(short mean free path)

Point: →  $l_{mp} < l_c$  ⇒ particle random  
walks parallel and undergoes  
many kicks in  $l_{co}$ . So parallel  
motion is diffusive.

→ perpendicular motion is continuous  
coarse graining/spreading, at  
 $D_{\perp} \sim \rho_e^2 v_{te} \sim \rho_e^2 \frac{v_{te}^2}{l_{mp}}$

So, can write:

$$\langle dr^2 \rangle \sim D_{\parallel} l_{co} \sigma$$

↓  
parallel correlation length  
(significant diffusive regime)

but also note that parallel motion is  
diffusive, so:

but time set by:

$$\chi_{||} / l_{cd}^2 \sim 1/t \quad \Delta$$

$$\Rightarrow \frac{\langle \sigma r^2 \rangle}{t} \sim \frac{\chi_{||}}{l_{cd}^2} D_M l_{cd}$$

$$\sim D_M \frac{\chi_{||}}{l_{cd}} \sim D_M \chi_{||} / l_{cd}$$

$$\chi_{\perp} = D_M \frac{\chi_{||}}{l_{cd}}$$

perpendicular heat conductivity in collisional regime.

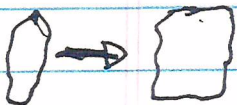
$$\chi_{\perp} = v_{th} D_M \frac{l_{mfp}}{l_{cd}}$$

Now what is  $l_{cd}$ ?

Notice  $l_{cd}$  is set by competition between 2 processes:

① ~~width  $\sigma$  increases due to diffusion~~

width  $\sigma$  increases due to diffusion (cross section)





so  $(d\sigma)^2 \sim (D_{\perp} dt)$   
 $d\sigma \sim (D_{\perp} dt)^{1/2}$

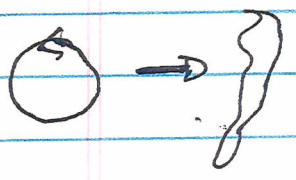
but  $\chi_{11} / (dL)^2 \sim 1/dt$

$\Rightarrow d\sigma \sim \left( \frac{D_{\perp} (dL)^2}{\chi_{11}} \right)^{1/2}$

$\frac{D_{\perp}}{\chi_{11}} \sim \frac{D_{\perp}}{L^2}$

$d\sigma \sim \left( \frac{D_{\perp}}{\chi_{11}} \right)^{1/2} dL$

② width shrinks, due stochastic instability and area conservation:



$d\sigma/dL = -\sigma/L_c$  (exponential decay)

then balance at:

$$d\sigma \sim \underbrace{\left( D_{\perp} / \chi_{11} \right)^{1/2}}_{\text{smoothly}} dL \sim \underbrace{\sigma}_{\text{local thinning}} dL$$

$$\sigma \sim l_c \left( D_{\perp} / \chi_{11} \right)^{1/2}$$

N.B.: Can select  $\sigma$  from:

$$\partial_t T - \chi_{11} \nabla_{11}^2 T - D_{\perp} \nabla_{\perp}^2 T = 0$$

$$\Rightarrow \frac{\chi_{11}}{l_c^2} \sim \frac{D_{\perp}}{\sigma^2}$$

$$\sigma \sim l_c \left( D_{\perp} / \chi_{11} \right)^{1/2}$$

Finally, need ~~correlation~~ length  $l_c$  for chunk size  $\sigma$ . Assume set by  $k_0$



$$k_0^{-1} \sim \sigma e^{z/l_c} \Big|_{l_c} \sim \sigma e^{l_c/l_c}$$

$$h_{cs} \sim h_c \ln(1/k_{cs})$$

$$h_{cs} \sim h_{cs} \left( \frac{\chi_{II}}{D_{\perp}} \right)^{1/2}$$

$$h_{cs} \sim h_c \ln \left( \left( \frac{\chi_{II}}{D_{\perp}} \right)^{1/2} / k_{cs} \right)$$

$$\Rightarrow \chi_{\perp} \sim D_{\perp} \chi_{II} / h_{cs}$$

Apart from a log factor:

$$\chi_{\perp} \sim v_{th} D_{\perp} \left( \frac{v_{th}}{h_c} \right)$$

$\ll 1$

$\Rightarrow$  reduced relative to collisionless values

- Lesson:
- collisions reduce (length scale) reduce  $\chi_{eff}$  relative to "Collisionless case"
  - interplay of perp and parallel diffusion
  - again, critical to knock particle off field line.

Now, the above calculation requires thought. Its much more convenient to crank ~~code~~ mindlessly.

⇒ Hydro approach: Kadomtsev and Pogutse (not mindless, but systematic)

Consider heat flux along wiggling fields  
d.e.

$$\underline{q} = -\chi_{||} \nabla_{||} T \hat{b} - \chi_{\perp} \underline{\nabla}_{\perp} T$$

↓ parallel conduction      ↓ perp. conduction

$$\chi_{||} \gg \chi_{\perp}$$

Strömgren codes

Here:  $\underline{b} = \underline{b}_0 + \underline{\tilde{b}}$   
 $\downarrow$  unperturbed  $\rightarrow$  Fluctuating

$$\nabla_{||} = \partial_z + \underline{\tilde{b}} \cdot \underline{\nabla}_{\perp}$$

↑  
piece along  
wavy line

8 seek mean radial heat flux

$$\langle q_{||} \rangle = -\kappa_{||} \left\{ \begin{array}{l} \textcircled{1} \langle b_{||}^2 \rangle \partial_r \langle T \rangle \\ \textcircled{2} \langle b_{||} \partial_z \tilde{T} \rangle \\ \textcircled{3} \langle b_{||} b_{\perp} \partial_r \tilde{T} \rangle \end{array} \right\} \left. \begin{array}{l} \text{usual} \\ \text{quadratic} \end{array} \right\}$$

$$= -\kappa_{\perp} \partial_r \langle T \rangle \rightarrow \text{cubic}$$

Now  $\underline{\textcircled{3}} \sim \frac{\kappa_{||} \tilde{b}_{||} \tilde{b}_{\perp} \tilde{T} / \Delta r}{\kappa_{||} \tilde{b}_{||} \tilde{T} / l_{ab}}$

$$\sim \frac{\tilde{b}_{\perp} l_{ab}}{\Delta r} \sim \kappa_{\perp}$$

∞ cubic nonlinearity dominates for  $Ku > 1$ .

$Ku < 1 \Rightarrow$  drop cubic.

To compute  $\langle \Sigma_r \rangle$ , need

- retain ① (usual), and ②

- iterate for  $\tilde{T}$  using

$$\underline{\nabla \cdot \tilde{q}} = 0 \quad \text{c.e. abs QLT.}$$

Thinking (geop!) first:

$$\begin{aligned} \langle \Sigma_r \rangle &\approx -\psi_{||} \left[ \langle k_{||}^2 \rangle \partial_r T + \langle \partial_{||} \partial_z \tilde{T} \rangle - \partial_z \partial_r \langle T \rangle \right] \\ &\approx -\psi_{||} \left[ \langle \tilde{b}_r \overbrace{b \cdot \nabla T} \rangle \right] - \psi_{\perp} \partial_z \langle T \rangle \\ &\quad \downarrow \text{linearization:} \\ &\quad \partial_{||} \partial_r \langle T \rangle + \partial_z \tilde{T} \end{aligned}$$

Point: - need non-zero  $\nabla \cdot \mathbf{T}$  fluctuations to drive heat flux

- c.e. temperature cant be constant along line, to drive parallel heat flux

-  $\nabla \cdot \mathbf{q} = 0 \Rightarrow$  result must imply  $v_{\perp}$  dependence!  
 $v_{\perp}$  to balance

⊗

$$\langle q_z \rangle = -\nu_{||} \left[ \langle \tilde{v}_r^2 \rangle \partial_r \langle T \rangle + \langle \tilde{v}_r \partial_z \tilde{T} \rangle \right] - \nu_{\perp} \nabla_{\perp} \langle T \rangle$$

$$\nabla \cdot \mathbf{q} = 0$$

$$\Rightarrow \nu_{||} \tilde{q}_{||} + \nabla_{\perp} \cdot \tilde{\mathbf{q}}_{\perp} = -\nu_{||} \partial_z \langle \tilde{v}_r \partial_r \langle T \rangle \rangle$$

c.e.



$$Q = \chi_{\parallel} \left[ (\partial_z + \tilde{b} \cdot \nabla) (\tau_0 + \tilde{\tau}) (\underline{b} + \tilde{b}) \right] - \chi_{\perp} \nabla_{\perp}^2 \tau$$

||

$$- \chi_{\parallel} \partial_z^2 \tau - \chi_{\perp} \nabla_{\perp}^2 \tau = - \chi_{\parallel} \partial_z \tilde{b} \frac{\partial \langle \tau \rangle}{\partial r}$$

↑

$$\frac{\partial \tau}{\partial r} = \frac{- \chi_{\parallel} k_z \tilde{b}_{\parallel} \langle \tau \rangle}{(\chi_{\parallel} k_z^2 + \chi_{\perp} k_{\perp}^2)}$$

||

$$\chi_{\parallel} \langle \tilde{b}^2 \rangle \frac{\partial \langle \tau \rangle}{\partial r} - \chi_{\parallel} \langle \tilde{b} \partial_z \tau \rangle$$

$$= - \chi_{\parallel} \sum_{\mathbf{n}} \left( - \frac{\chi_{\parallel} k_{\parallel}^2 |\tilde{b}_{\parallel}|^2}{\chi_{\parallel} k_z^2 + \chi_{\perp} k_{\perp}^2} + |\tilde{b}_{\perp}|^2 \right) \frac{\partial \langle \tau \rangle}{\partial r}$$

$$= - \chi_{\parallel} \frac{\partial \langle \tau \rangle}{\partial r} \sum_{\mathbf{n}} \left( \frac{- \chi_{\parallel} k_{\parallel}^2}{\chi_{\parallel} k_z^2 + \chi_{\perp} k_{\perp}^2} + \frac{\chi_{\parallel} k_{\parallel}^2 + \chi_{\perp} k_{\perp}^2}{\chi_{\parallel} k_z^2 + \chi_{\perp} k_{\perp}^2} \right)$$



80

$$\langle Q_r \rangle_{NL} = -\chi_{11} \frac{\partial \langle T \rangle}{\partial r} \sum_n \frac{\chi_{\perp} k_{\perp}^2 |b_n|^2}{\chi_{11} k_{11}^2 + \chi_{\perp} k_{\perp}^2}$$

Note explicit dependence on  $\chi_{\perp}$ !

80

$$\langle Q_r \rangle_{NL} \approx -\chi_{11} \frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \int d\underline{k}_{\parallel} \frac{\chi_{\perp} k_{\perp}^2 \langle \tilde{b}_n^2 \rangle}{\chi_{11} (k_{\parallel}^2 + \frac{\chi_{\perp} k_{\perp}^2}{\chi_{11}})}$$

$$= -\frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \int d\underline{k}_{\parallel} \frac{\chi_{\perp} k_{\perp}^2 \langle \tilde{b}_n^2 \rangle}{\left( \frac{k_{\parallel}^2}{(\chi_{\perp}/\chi_{11}) k_{\perp}^2} + 1 \right) \left( \frac{\chi_{\perp} k_{\perp}^2}{\chi_{11}} \right)}$$

$$= -\frac{\partial \langle T \rangle}{\partial r} \int d\underline{k}_{\perp} \frac{k_{\perp}^2 (\chi_{11} \chi_{\perp})^{1/2}}{\sqrt{k_{\perp}^2}} \langle \tilde{b}_n^2 \rangle_{\text{pac}}$$

~~auto correlation~~  
auto correlation

bandwidth  $k_{\perp}$

auto correlation  $\rho_{cc}$  enters via normalization

$\Rightarrow$

$$\langle q_r \rangle_{cc} \approx -\sqrt{\chi_{11} \chi_{\perp}} \langle \tilde{b}^2 \rangle_{cc} \sqrt{\chi_{\perp}^2} \frac{d\langle T \rangle}{dr}$$

Note: - need  $\nabla_{11} \hat{T} \neq -\tilde{b}_r d\langle T \rangle / dr$   
 ( $B \cdot \nabla T \neq 0$ ) for  $\perp$  heat flux

-  $\langle \tilde{b}^2 \rangle_{cc} \sim D_M$

$\sqrt{\chi_{\perp}^2} \sim 1 / \Delta_{\perp}$

so

$$\langle q_r \rangle \approx -\chi_{\perp} \chi_{\perp} \frac{d\langle T \rangle}{dr} - \chi_{\perp} \frac{d\langle T \rangle}{dr}$$

$$\chi_{\perp} \approx \frac{\sqrt{\chi_{11} \chi_{\perp}} D_M}{\Delta_{\perp}}$$

$$\left( \begin{array}{l} \chi_{11} \chi_{\perp} \sim \\ \frac{\chi_{11}^2}{\chi_{\perp}} \sim D_B \end{array} \right)$$

$$\chi_{\perp \text{eff}} \approx \frac{D_B D_M}{\Delta_{\perp}}$$

-  $\chi_{\perp \text{eff}}$  scales with  $Rehm$ , not Spitzer ( $\chi_{\parallel}$ )

- width of  $\omega p$  line important, again.

To compare R & R:

$$\chi_{\perp} \sim \sqrt{\chi_{\parallel} \chi_{\perp}} \frac{\langle \tilde{b}^2 \rangle^{1/2} l_{eb}}{\Delta_{\perp}}$$

what is  $\Delta_{\perp}$ ?

Now

$$\frac{\chi_{\parallel}}{l_0^2} \sim \frac{\chi_{\perp}}{\Delta_{\perp}^2}$$

diffusion set  
 $\perp$  scale.

$$\Delta_{\perp} \sim l_0 \sqrt{\chi_{\perp} / \chi_{\parallel}}$$

↑  
 enters spectrum,  
 (small layer)

⇒

$$\chi_I \sim \sqrt{\chi_{II} \chi_I} \frac{\langle \delta^2 \rangle_{lc}}{lc} \frac{1}{lc (\chi_I / \chi_{II})^{1/2}}$$

$$\chi_I \sim \frac{\chi_{II} \Delta M}{lc}$$

• so - modulo  $\chi_{II}, \Delta M$ ; agrees with  $R \times R$  to within log. factor

$$- \chi_I \sim \chi_{II} \Delta M \frac{\ln \Delta M}{lc}$$

⇒ covers diffusion in kernel of stochastic fields

⇒ Lesson: Take care re: irreversibility ↓