

## A Brief Look at Other Directions

→ Primary focus on transport,  
secondary on self-organization  
(no profits!!)

→ Principal Application → Magnetic  
Confinement.

Other directions:

- Kinematic Waves, Traffic Flow  
→ Whitham

→ phenomenological macroscopic

→ car-following (CA → fluids)

Application: Time-Delay induced  
jams. (TA?)

(cf: Kosyga, P. D., Gurcan  
PRL 73).

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- Flocking / "Boids" (ML for Flocking Rule)

- Patterns in swarm of self-propelled agents.

c.f. { Toner and Tu PRE 198  
Marchetti, et al. RMP '13

- Phenomenological (compressible) Hydro.  
 $\rho, v$

Broken  $\sigma$ -I  $\rightarrow$  Flock in dissipative media.

$\rightarrow$  dissipative structure

{ Very similar to car-following,  
Modes, structures.

c.e. Flocking, etc. is big area in biophysics.

Flock vs Avalanche?

→ MFE

- more on Avalanche, etc:  
see { Holm, P.D. '18  
    { QV talk

- Zonal Flow → staircase

→ Dif-P 1/0 PRE,  
→ Ashpurner, P.D. '17

and

→ Ortschel-McIntyre '08

ZF → outer scale

Open issue: Scale selection struggle.

→ Kinetics → EPM, etc.

- Champ emission → avalanche  
    ↓  
    Cerenkov.

→ Phase space avalanche ↓ ↓

Kinetics → Mean Field Theory

→ Thoughts on Ausbacher

Mean Field Theory - DW → Transport

Consider DKE → simplicity

$$\frac{\partial f}{\partial t} + v_z \frac{\partial f}{\partial z} - \frac{c}{B_0} \nabla \phi \times \hat{z} \cdot \nabla f - \frac{1}{m_0} E_z \frac{\partial f}{\partial v_z} = 0$$

(c.c. electrons) → B<sub>0</sub> = B<sub>0</sub> z stat.

then

$$\frac{\partial \langle f \rangle}{\partial t} + \frac{\partial \langle \tilde{v}_r \rangle}{\partial r} + \frac{\partial \langle \tilde{a}_z \rangle}{\partial v_z} = 0$$

↑  
transport
↑  
heating

n.b.  $\frac{\partial f}{\partial t} = \frac{1}{T} \phi \langle f \rangle + \tilde{h}$

in linear theory, with Maxwellian:

$$\frac{\partial}{\partial t} \left( \frac{1}{T} \phi \langle f \rangle + \tilde{h} \right) + v_z \frac{\partial}{\partial z} \left( \frac{1}{T} \phi \langle f \rangle + \tilde{h} \right) - \frac{c}{B} \nabla \phi \times \hat{z} \cdot \nabla \left( \frac{1}{T} \phi \langle f \rangle + \tilde{h} \right) - \frac{1}{m_0} E_z \frac{\partial}{\partial v_z} \left( \frac{1}{T} \phi \langle f \rangle + \tilde{h} \right) = 0$$

(generally  $\hbar \omega < \hbar \omega_0$ ) ;

$$\begin{aligned} \partial_t \tilde{h} + v_z \partial_z \tilde{h} - \frac{c}{B_0} \nabla \phi \times \vec{e}_z \cdot \nabla \tilde{h} - \frac{\hbar c}{m \omega} E_z \partial_{v_z} \tilde{h} \\ = - \partial_t \frac{\hbar c \phi}{\omega - \omega_0} - \frac{c}{B_0} \nabla \phi \times \vec{e}_z \cdot \nabla \langle f_0 \rangle \end{aligned}$$

Letting

$$\mathcal{L} f_0 = \frac{\phi_0}{\omega - \hbar \omega_0} L_0 \langle f \rangle$$

$$L_0 = - \frac{c \hbar_0}{B_0} \frac{\partial}{\partial r} + \frac{\hbar c \hbar_0}{m_e} \frac{\partial}{\partial v_z}$$

Scattering operator

$$\begin{aligned} \partial_t \langle f \rangle = & \overset{\textcircled{1}}{\partial_r} D_{r,r} \overset{\textcircled{2}}{\partial_r} \langle f \rangle + \overset{\textcircled{2}}{\partial_r} D_{r,v} \overset{\textcircled{4}}{\partial_{v_z}} \langle f \rangle \\ & + \overset{\textcircled{3}}{\partial_{v_z}} D_{v_z,r} \overset{\textcircled{2}}{\partial_r} \langle f \rangle + \overset{\textcircled{4}}{\partial_{v_z}} D_{v_z,v} \overset{\textcircled{4}}{\partial_{v_z}} \langle f \rangle \end{aligned}$$

$$D_{D,r} = n \sum_{\mathbf{k}} \frac{e^2}{B_0} k_0^2 |\Phi_{\mathbf{k}}|^2 \frac{0}{\omega - k_z v_z}$$

symmetry  $\rightarrow$  only  $\downarrow$

$$D_{D,v} = D_{v,r} = n \sum_{\mathbf{k}} \frac{e^2}{B_0} \frac{k_z}{m_0} k_0^2 |\Phi_{\mathbf{k}}|^2 \frac{i}{\omega - k_z v_z}$$

$$D_{v,v} = n \sum_{\mathbf{k}} \frac{e^2}{m_0^2} k_z^2 |\Phi_{\mathbf{k}}|^2 \frac{1}{\omega - k_z v_z}$$

① + ②  $\rightarrow$  radial transport

$$\text{c.f.} \sim \partial_r \Gamma_{v,r}$$

④  $\rightarrow$  parallel heating

Illuminating to examine momentum flux:

$$\begin{aligned} \partial_t \int \langle P \rangle v_z &= \partial_r \left[ \int dv_z v_z D_{D,r} \partial_r \langle P \rangle \right] \\ &+ \left[ \int dv_z v_z D_{v,v} \partial_{v_z} \langle P \rangle \right] \\ &- \int dv_z D_{v,r} \partial_r \langle P \rangle \\ &- \int dv_z D_{v,v} \partial_{v_z} \langle P \rangle \end{aligned}$$

① →  $\chi_\phi$ , etc.;  $\Pi_{resid}$   
 $\partial_r \langle F \rangle \rightarrow \partial_r \langle T \rangle$ , etc.

② →  $\Pi_{resid}$ , also.

i.e.  $\Pi_{resid} = \int dV_2 \quad v_2 \sum_n \frac{c_n |k| \omega_n v_2 \langle \phi_n^2 \rangle}{4 B m \omega - \omega_n} \partial_r \langle F \rangle$

③, ④ → Turbulent acceleration  
 (aka! L.Wang, P.D.)

③ =  $-\int dV_2 \quad D_{2,r} \partial_r \langle F \rangle$

=  $-\int dV_2 \sum_n \frac{c_n |k| \omega_n v_2 \langle \phi_n^2 \rangle}{4 B m \omega - \omega_n} \partial_r \langle F \rangle$   
 symbolically

and not  $\sim \nabla \cdot \Pi$ ,  $\partial_r \langle F \rangle \rightarrow$  driver flow

④ → another piece of turbulent source

→ Correlation Times

→ Useful here to explore correlation times, for different cases, and compare to 1D



# PKU Lecture III

→ "QL" Theory For Simple Drift Wave Turbulence, cont'd

- correlation times, Kubo #  $\leftrightarrow$  relation to structure,
- relaxation - transport  $\leftrightarrow$  heating constraint
- avalanche ?!

→ PV Conservation, PV Dynamics

→ Taylor Identity, and Charney-Drazin (Momentum) Theorem,

→ Relation Modulational Theory & ZF Formation

Recall:

$$\frac{\partial F}{\partial t} + v_z \frac{\partial F}{\partial z} - \frac{c}{B_0} \nabla \varphi \times \mathbf{z} \cdot \nabla F - \frac{k_z E_z}{m_0} \frac{\partial F}{\partial v_z} = 0$$

$$\frac{k_z E_z}{m_0} \rightarrow \text{loss}$$

$$\delta F_{\perp}^G = \frac{\varphi_{\perp}}{\omega - k_z v_z} \underline{L}_{\perp} \langle F \rangle$$

$$\underline{L}_{\perp} = -\frac{c}{B_0} k_{\perp} \frac{\partial}{\partial r} + \frac{k_z E_z}{m_0} \frac{\partial}{\partial v_z}$$

$\delta$   
scattering operator.

$$\frac{\partial \langle F \rangle}{\partial t} = \underbrace{\partial_r D_{rr}}_{\textcircled{1}} \partial_r \langle F \rangle + \underbrace{\partial_n D_{nn}}_{\textcircled{2}} \partial_n \langle F \rangle + \underbrace{\partial_v D_{vv}}_{\textcircled{3}} \partial_v \langle F \rangle + \underbrace{\partial_v D_{vv}}_{\textcircled{4}} \partial_v \langle F \rangle$$

①, ② → Momentum Flux, including  
turbid,  $\nabla$

③, ④ → turbulent acceleration.

N.B. Show  $\Gamma_e^{\perp} = \Gamma_i^{\perp}$ , in QLT.

Now, seek examine  $\gamma_{\text{eff}}$  for  
 drift waves,

Spectral structure  $\Leftrightarrow$  Geometry  $\Leftrightarrow$  Kubo #  
 Connection

$$D_{\perp, \pm} = \sum_n \frac{c^2}{B_0^2} |\phi_n|^2 k_0^2 \frac{c}{\omega - k_z v_z}$$

assume trial spectrum structure:

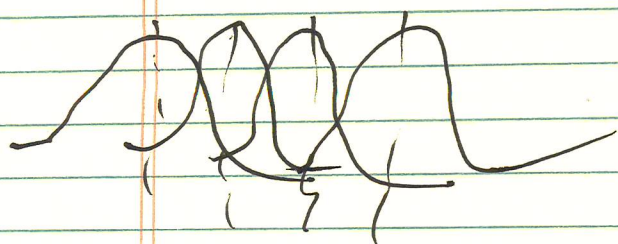
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$$|\phi_n|^2 = |\phi_0|^2 \left( \frac{\Delta k_0}{(k_0 - \bar{k}_0)^2 + (\Delta k_0)^2} \right) \left( \frac{\Delta k_z}{k_z^2 + (\Delta k_z)^2} \right)$$

$\uparrow$  centroid  $\bar{k}_0$        $\downarrow$  spread  $\Delta k_0$        $\downarrow$  spread  $\Delta k_z$

$\rightarrow$  ad-hoc but useful to extract  
 time scales.

For sheared system:



$$z' > 0$$

$$m = nq$$

$$\begin{aligned} dn &= \frac{M}{z^2} dx \\ &= \frac{k_0 v_z'}{z^2} dx \end{aligned}$$

$$|\Phi_{\perp}|^2 = |\Phi_0|^2 \delta'(k_0) F\left(\frac{\nu - \nu_0}{\Delta\nu}\right)$$

↳ spectral width, SFLC.

$$\sum_m \sum_n \rightarrow \int dk_0 \int \frac{h \text{arg}^2 dx}{z^2}$$

if overlaid

Sum over modes sweeps spectral width

→  $\Delta k_z$  corresponds to spectral width integration

→ can extend to ballooning space, extent mode slow line ↔ couplings.

Now,

$$D_{1,2} \approx \frac{\hat{c}}{\omega_{\perp} - \tilde{\omega}_z \nu_z + i/\tau_{20}}$$

$$1/\tau_{20} = \left| \frac{\Delta k_{\perp} \frac{d\omega}{dk_{\perp}}}{\omega} - \nu_z \Delta k_z \right|$$

~> have assumed  $d\omega/dk \rightarrow$

Result 1D:

$$1/T_{ev} = \left| \frac{d\omega}{dk} - v \right| \Delta k$$

resonant

$$= \left| \frac{d\omega}{dk} - \frac{\omega}{k} \right| \Delta k$$

i.e.

Wave-particle auto-correlation time depends on particle speed.

Can see:

- electrons (resonant) cons near Mach  $\omega/k_{||} \rightarrow \omega/k_{||} \leq v_{th}$  — ITC

$$1/T_{ev} = \left| \Delta k_{||} \frac{d\omega}{dk_{||}} - \frac{\omega}{k_{||}} \Delta k_{||} \right|$$

$$\approx \left| \Delta k_{||} \frac{d\omega}{dk_{||}} - v_{th} \Delta k_{||} \right|$$

$$1/\tau_{\text{EDW}} \sim (V_{\text{th}} \Delta k_{\text{EDW}}) \rightarrow \text{edw.}$$

$$1/\tau_{\text{EDW}} \sim (V_{\text{th}} \Delta k_{\text{EDW}}) \rightarrow \text{ITG, near} \\ \text{margin} \\ \text{(resonant case).}$$

$$1/\tau_{\text{EDW}} \sim (\Delta k_{\text{EDW}} dW/dk_{\text{EDW}}) \rightarrow \text{ITG above} \\ \text{margin}.$$

N.B. Ion diffusion on EDW is non-resonant.

Also:

→ Geometry / spectral structure reduces sensitivity (to dispersion) of  $\tau_{\text{ec}}$ .

$$\rightarrow k_{\text{u}} \sim \tau_{\text{ec}} \frac{\tilde{v}}{\Delta} < 1 \Rightarrow \text{validates} \\ \text{stochastic} \\ \text{transport} \\ \text{modelling.}$$

if:  $\tau_{\text{ec}} \omega_{*} < 1$

$$\rightarrow k_{\text{u}} < 1 \rightarrow \text{EDW} \\ \rightarrow \sim \text{mod, ITG}$$

$$k_{\text{u}} \rightarrow 1 \text{ for fluid ITG.}$$

# Precession Resonance: An Interesting Twist

$$D_{\pm, \pm} = \frac{C^2}{B_0^2} \sum_N \frac{N^2 \tau^2}{N^2} |\Phi_N|^2 \frac{i}{\omega - \omega_0 \epsilon}$$

↕  
broaden average.

as before;

$$\frac{i}{\omega - \omega_0 \epsilon} \Rightarrow \frac{1}{\frac{d\omega}{d\epsilon} \Delta\epsilon - \omega_0 \Delta\epsilon}$$

$\epsilon = \omega / \omega_0$  for resonant particles

$$\sim \frac{1}{|\Delta\epsilon| \left[ \frac{d\omega}{d\epsilon} - \frac{\omega}{\epsilon} \right]}$$

$$\frac{1}{\tau_{20}} \equiv |\Delta\epsilon| \left[ \frac{d\omega}{d\epsilon} - \frac{\omega}{\epsilon} \right]$$

$\rightarrow$   $\Delta\epsilon \neq 0 \rightarrow$  dispersion sensitive

$\rightarrow$  precession resonance mode  
(CTEM, EPM)  $\tau_{20} \neq 0$

Other cases

→ precession:

$$D = \frac{c^2}{B_0} \sum_n \frac{n^2 \langle \vec{\phi}_n \rangle^2}{r^2} \frac{c}{\omega - \omega_0}$$

as before

$$\frac{c}{\omega - \omega_0} \sim \frac{1}{\left| \frac{d\omega}{dn} \right| \Delta n - \omega_0 \Delta n} \quad \epsilon = \omega / \omega_0$$

$$\epsilon \sim \frac{\omega}{\omega_0} \sim \frac{1}{\left| \frac{d\omega}{dn} \right| \Delta n - \omega_0 \Delta n} \frac{\omega}{n_0}$$

$$\tau_{ac} \sim \frac{1}{|\Delta n| \left[ \left| \frac{d\omega}{dn} \right| - \frac{\omega}{n_0} \right]}$$

→  $|\Delta n| \propto 10 \rightarrow$  dispersion sensitive

→ CTM tend to run low  $\uparrow$ , as

$\tau_{ac} \uparrow$

N.B.: Precession resonance drives turbulence  
has long wave-particle correlation times.



large  $k_{\parallel}$ .

→ ? How treat?

- granulations <sup>strongly</sup> ⇒ correlated clusters?

~~\*\*\*~~  $y.k_{\parallel} = \text{Ph.D.}$ , et seq.

- but still statistical??

→ Useful to consider single, few coherent CTEM.

BGK solution?

⇒ HW: Effect  $E \times B$  shear on  $\gamma_{\text{EB}}$

- EDW

- precession

What is Q.L. stationary for sample drift wave?

7/3

→ Relaxation

$$\frac{\hbar k}{m} v_z \rightarrow \frac{\hbar k}{m_0} \frac{\omega}{v_z}$$

from  
vetch

$$L_n \langle f \rangle = \frac{\hbar k}{\Sigma_0} \frac{\partial \langle f \rangle}{\partial r} + \frac{\omega}{v_z} \frac{\partial \langle f \rangle}{\partial v_z}$$

and

$$\partial_t \langle f \rangle = \sum_n [L_n |\phi_n|^2 \pi \delta(\omega - \omega_n) L_n \langle f \rangle]$$

relaxn: what is Q.L. stationary state?

$$L_n \rightarrow 0$$

$$-\frac{\hbar k}{\Sigma_0} \frac{\partial \langle f \rangle}{\partial r} + \frac{\omega}{v_z} \frac{\partial \langle f \rangle}{\partial v_z} \rightarrow 0$$

~~scribble~~ \*

$$\left( -\frac{\hbar k}{\Sigma_0} \frac{1}{A r} + \frac{\omega / v_z}{A v_z} \right) \langle f \rangle \rightarrow 0$$

$$A r = \frac{A v_z^2}{2} \frac{\hbar k}{\Sigma_0 \omega v_z} = 0$$

$$A \left( r = \frac{v_z^2}{2} \frac{\hbar k}{\Sigma_0 \omega v_z} \right) = 0$$

constraint  
for  
relaxed state

Formally:

$$\partial_t \langle F \rangle = \left[ \sum_{\mu} L_{\mu} |\Phi_{\mu}|^2 \pi c (\omega - \omega_{\mu}) L_{\mu} \right] \langle F \rangle$$

$$\partial_t \int dr \int dv_3 \frac{\langle F \rangle^2}{2} = - \int dr \int dv_3 \sum_{\mu} |\Phi_{\mu}|^2 \pi c (\omega - \omega_{\mu}) + (L_{\mu} \langle F \rangle)^2$$

$$L_{\mu} \langle F \rangle = 0 \Rightarrow \text{DW "plateau"}$$

N.B. - Note this is not simply  
flattening of density

-  $L_{\mu} \rightarrow 0$  before  $\frac{d\langle F \rangle}{dt} \rightarrow 0$ , as  
heating  $\Delta v_3^2$  occurs.

$$\Delta r = \frac{\Delta v_z^2}{2} \frac{v_{th}}{\Omega_e \omega_{pe}}$$

$$r = \frac{v_z^2}{2} \frac{v_{th}}{\Omega_e \omega_{pe}} = \text{const.}$$

→ i.e. any  $\Delta r > 0$  displacement  $\Rightarrow$  heating ( $\Delta v_z^2 > 0$ )

how QL saturate without  $\Delta r \rightarrow 0$ ?

→ DW extends energy on heating via Landau damping  $\Rightarrow$  route for non-trivial QL stat.

W on W-wave.

$$\Delta v_z^2 \sim \Delta r \frac{\partial v_{th}^2}{\partial n} \frac{1}{L_n} \frac{\partial v_{th}^2}{\partial n}$$

$$\frac{\Delta v_z^2}{v_{th}^2} \sim \frac{\Delta r}{L_n} \frac{v_{th}^2}{v_{th}^2}$$

→ heating - transport relation

Lesson

- DW's do distort distribution function

- relevant to ITG near-machine

- heating often overlooked

Comment:

→ How include source  $\Rightarrow$  headed toward sink

Flux driven

$$\partial_t \langle F \rangle = L D L \langle F \rangle + \int_0^n (r, t)$$

dynamic coupled } transport  
heating

→ What of avalanche?

- avalanche in phase space

- must evolve in  $r$  (as usual)

and  $V_z$  d.e. heating penalty for step  $\Delta r$ .

"bi-variate Burgers" with

constraint on radial,  $v_{||}$  scattering?

and Energetics:

- Now DKE /  $Q_L$ ?

$$\partial_t \langle E_{mi} \rangle + \frac{\partial}{\partial r} Q_e - \langle E_2 J_2 \rangle = 0$$

↑  
radial transport

and Thm:

$$\partial_t W_{ow} + \partial_r S_r + \langle E_2 J_2 \rangle_R = 0$$

↑  
radial wave energy density flux

$$\Rightarrow \langle E_2 J_2 \rangle_R = \partial_t \langle E_{mi} \rangle_R + \partial_r Q_{e,R}$$

$$\partial_t \langle E_{mi} \rangle_R + \partial_t W_{ow} + \partial_r (Q_{e,R} + S_r) = 0$$

↓ RP + wave as before
↓ transport
↓ inhomog wave transport

n.b.  $S_r \leftrightarrow Q_{e,R}$

$S_r \sim \nabla_r \sum \omega$  → RF generation

- Dominant Balance:

$$\partial_t \langle E_{kin} \rangle_R + \nabla_n Q_{e,R} = 0$$

N.B. show:  $\Gamma_e = \Gamma_b$   
 $\int_{res}$   $\int_{NR}$

- at edge:

$$W_{ow} \rightarrow \Sigma_{kin}^{pts}$$

$$\nabla_n \rightarrow Q_{e,R}$$

Do wave radiative losses significant,  
 $\Rightarrow ZF$

~~W~~

c.s.  $W = \frac{k_0 V_A}{1 + k_1^2 \lambda^2}$

$$\nabla_r = - \frac{2 \boxed{k_r k_0} \lambda^2 V_A}{(1 + k_1^2 \lambda^2)^2}$$

$\nabla_r \Sigma \rightarrow$   
wave energy density flux  
 $\downarrow$   
Reynolds stress

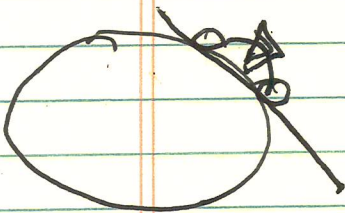
→ PV Mixing + G-D Thms — QLT for Fluid DW

Closely related to Vlasov Dynamics  
is  $\beta$ -Plane / QG

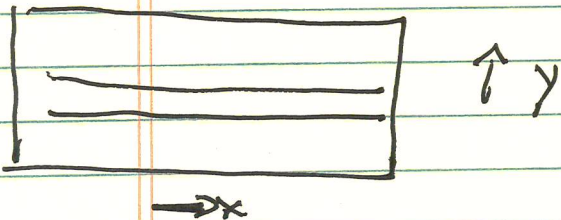
$$\frac{\partial}{\partial t} (\nabla^2 \phi + \beta y) + \langle \underline{v} \rangle \cdot \nabla (\nabla^2 \phi + \beta y) + \nabla \cdot \nabla (\nabla^2 \phi + \beta y) - \nu \nabla^2 (\nabla^2 \phi + \beta y) = 0$$

$$\partial_t \nabla^2 \phi + \langle \underline{v} \rangle_x \partial_x \nabla^2 \phi + \underline{\tilde{v}} \cdot \nabla \nabla^2 \phi - \nu \nabla^2 \nabla^2 \phi = - \tilde{\nabla}_y (\beta + \langle \nabla^2 \phi \rangle)$$

Mean vorticity



$$\Sigma = \langle \varepsilon \rangle + \tilde{\Sigma}$$



keys:

$$\partial_t \Sigma + \underline{v} \cdot \nabla \Sigma = \nu \nabla^2 \Sigma$$

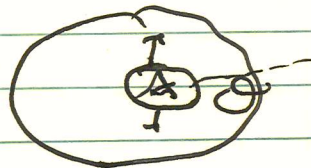


Kelvin's Thm.  $\int \underline{v} \cdot d\underline{l} = \int \underline{\omega} \cdot d\underline{a} = \text{const.}$

$$\underline{\omega} \rightarrow \underline{\omega} + 2\underline{\Omega}$$

$$C = \int d\underline{a} \cdot (\underline{\omega} + 2\underline{\Omega})$$

$$\dot{C} = 0 \Rightarrow$$



$$A \frac{d\underline{\omega}}{dt} = -2\underline{\Omega} \sin \theta_0 \frac{dA}{dt}$$

$$\frac{d\underline{\omega}}{dt} = \frac{-2\underline{\Omega} \sin \theta_0}{A} \frac{dA}{dt}$$

$$= -2\underline{\Omega} \frac{d\theta}{dt} \sin \theta_0$$

→ x. by

$$= -\frac{2\underline{\Omega}}{R} (R \frac{d\theta}{dt}) \sin \theta_0$$

$$= -\underline{\Omega} \underline{v}_y$$

$$\underline{\omega} = \nabla^2 \phi$$

$$\underline{v} = \frac{-\nabla \phi}{2\underline{\Omega}}$$