$L = \frac{\lambda^2}{4\pi^2 Cc^2} = \frac{c550 \times 10^{-9} \text{ m}}{4\pi^2 c17 \times 10^{-12} \text{ Fbc } 2.998 \times 10^8 \text{ m/sb}^2} = 5.00 \times 10^{-21} \text{ H}.$

same as the frequency of an electromagnetic wave.

 $f\lambda = c$. Thus,

This is exceedingly small.

LEARN The frequency is

 $\frac{\lambda}{2\pi\sqrt{IC}} = c.$

5. **THINK** The frequency of oscillation of the current in the LC circuit of the generator is

 $f = 1/2\pi\sqrt{LC}$, where C is the capacitance and L is the inductance. This frequency is the

EXPRESS If f is the frequency and λ is the wavelength of an electromagnetic wave, then

$$\frac{\lambda}{2\pi\sqrt{LC}} = c.$$

ANALYZE The solution for L is

$$f = \frac{c}{\lambda} = \frac{3.0 \times 10^8 \text{ m/s}}{550 \times 10^{-9} \text{ m}} = 5.45 \times 10^{14} \text{ Hz}.$$

The EM wave is in the visible spectrum.

6. The emitted wavelength is

$$\lambda = \frac{c}{f} = 2\pi c \sqrt{LC} = 2\pi \left(2.998 \times 10^8 \,\text{m/s} \right) \sqrt{\left(0.253 \times 10^{-6} \,\text{H} \right) \left(25.0 \times 10^{-12} \,\text{F} \right)} = 4.74 \,\text{m}.$$

8. The intensity of the signal at Proxima Centauri is

 $I = \frac{P}{4\pi r^2} = \frac{1.0 \times 10^6 \text{ W}}{4\pi \text{ b4.3 lygc9.46} \times 10^{15} \text{ m/lyh}^2} = 4.8 \times 10^{-29} \text{ W/m}^2.$

11. (a) The amplitude of the magnetic field is

$$B_m = \frac{E_m}{c} = \frac{2.0 \text{V/m}}{2.008 \times 10^8 \text{ m/s}} = 6.67 \times 10^{-9} \text{ T} \approx 6.7 \times 10^{-9} \text{ T}.$$

(b) Since the \vec{E} -wave oscillates in the z direction and travels in the x direction, we have $B_x = B_z = 0$. So, the oscillation of the magnetic field is parallel to the y axis.

(c) The direction (+x) of the electromagnetic wave propagation is determined by $\vec{E} \times \vec{B}$. If the electric field points in +z, then the magnetic field must point in the -y direction.

With SI units understood, we may write

$$B_{y} = B_{m} \cos \left[\pi \times 10^{15} \left(t - \frac{x}{c} \right) \right] = \frac{2.0 \cos \left[10^{15} \pi \left(t - x/c \right) \right]}{3.0 \times 10^{8}}$$
$$= \left(6.7 \times 10^{-9} \right) \cos \left[10^{15} \pi \left(t - \frac{x}{c} \right) \right]$$

13. (a) We use
$$I = E_m^2/2\mu_0 c$$
 to calculate E_m :
$$E_m = \sqrt{2\mu_0 I_c} = \sqrt{2c4\pi \times 10^{-7} \text{ T} \cdot \text{m/A} / c} = \sqrt{2c4\pi \times 10^{-7} \text{$$

 $=1.03\times10^3\,\mathrm{V}\,/\,\mathrm{m}.$

(b) The magnetic field amplitude is therefore

 $B_m = \frac{E_m}{c} = \frac{1.03 \times 10^4 \text{ V/m}}{2.998 \times 10^8 \text{ m/s}} = 3.43 \times 10^{-6} \text{ T}.$

 $I = \frac{E_m^2}{2\mu_0 c} = \frac{b2.0 \,\text{V/mg}^2}{2c4\pi \times 10^{-7} \,\text{T} \cdot \text{m/A} / c2.998 \times 10^8 \,\text{m/sh}} = 5.3 \times 10^{-3} \,\text{W/m}^2.$

 $P = 4\pi r^2 I_{\text{avg}} = 4\pi (10 \,\text{m})^2 (5.3 \times 10^{-3} \,\text{W/m}^2) = 6.7 \,\text{W}.$

(b) The intensity is

(c) The power of the source is

19. **THINK** The plasma completely reflects all the energy incident on it, so the radiation pressure is given by $p_r = 2I/c$, where I is the intensity.

EXPRESS The intensity is I = P/A, where P is the power and A is the area intercepted by the radiation.

ANALYZE Thus, the radiation pressure is

$$p_r = \frac{2I}{c} = \frac{2P}{Ac} = \frac{2(1.5 \times 10^9 \text{ W})}{(1.00 \times 10^{-6} \text{m}^2)(2.998 \times 10^8 \text{ m/s})} = 1.0 \times 10^7 \text{ Pa}.$$

LEARN In the case of total absorption, the radiation pressure would be $p_r = I/c$, a factor of 2 smaller than the case of total reflection.

20. (a) The radiation pressure produces a force equal to

$$F_r = p_r \left(\pi R_e^2 \right) = \left(\frac{I}{c} \right) \left(\pi R_e^2 \right) = \frac{\pi \left(1.4 \times 10^3 \,\mathrm{W/m^2} \right) \left(6.37 \times 10^6 \,\mathrm{m} \right)^2}{2.998 \times 10^8 \,\mathrm{m/s}} = 6.0 \times 10^8 \,\mathrm{N}.$$

(b) The gravitational pull of the Sun on the Earth is

$$F_{\text{grav}} = \frac{GM_sM_e}{d_{es}^2} = \frac{\left(6.67 \times 10^{-11} \,\mathrm{N \cdot m^2 / kg^2}\right) \left(2.0 \times 10^{30} \,\mathrm{kg}\right) \left(5.98 \times 10^{24} \,\mathrm{kg}\right)}{\left(1.5 \times 10^{11} \,\mathrm{m}\right)^2}$$
$$= 3.6 \times 10^{22} \,\mathrm{N},$$

22. The radiation pressure is

 $p_r = \frac{I}{c} = \frac{10 \,\mathrm{W/m^2}}{2.998 \times 10^8 \,\mathrm{m/s}} = 3.3 \times 10^{-8} \,\mathrm{Pa}.$

26. The mass of the cylinder is $m = \rho(\pi D^2/4)H$, where D is the diameter of the cylinder. Since it is in equilibrium

$$F_{\text{net}} = mg - F_r = \frac{\pi H D^2 g \rho}{4} - \left(\frac{\pi D^2}{4}\right) \left(\frac{2I}{c}\right) = 0.$$

We solve for H:

 $H = \frac{2I}{gc\rho} = \left(\frac{2P}{\pi D^2/4}\right) \frac{1}{gc\rho}$ 2(4.60W)

 $[\pi(2.60\times10^{-3} \text{ m})^2/4](9.8 \text{ m/s}^2)(3.0\times10^8 \text{ m/s})(1.20\times10^3 \text{ kg/m}^3)$

$$H = \frac{2}{gc}$$

 $=4.91\times10^{-7}$ m.

EXPRESS If the beam carries energy U away from the spaceship, then it also carries momentum p = U/c away. By momentum conservation, this is the magnitude of the momentum acquired by the spaceship. If P is the power of the laser, then the energy carried away in time t is U = Pt.

29. **THINK** The laser beam carries both energy and momentum. The total momentum of

the spaceship and light is conserved.

ANALYZE We note that there are 86400 seconds in a day. Thus, p = Pt/c and, if m is mass of the spaceship, its speed is

$$v = \frac{p}{m} = \frac{Pt}{mc} = \frac{(10 \times 10^3 \text{ W})(86400 \text{ s})}{(1.5 \times 10^3 \text{ kg})(2.998 \times 10^8 \text{ m/s})} = 1.9 \times 10^{-3} \text{ m/s}.$$

LEARN As expected, the speed of the spaceship is proportional to the power of the laser beam.

1/2. After passing through the second one it is further reduced by a factor of $\cos^2(\pi - \theta_1 - \theta_2) = \cos^2(\theta_1 + \theta_2)$. Finally, after passing through the third one it is again reduced by a factor of $\cos^2(\pi - \theta_2 - \theta_3) = \cos^2(\theta_2 + \theta_3)$. Therefore,

32. After passing through the first polarizer the initial intensity I_0 reduces by a factor of

$$=4.5\times10^{-4}$$
.
Thus, 0.045% of the light's initial intensity is transmitted.

 $\frac{I_f}{I_0} = \frac{1}{2}\cos^2(\theta_1 + \theta_2)\cos^2(\theta_2 + \theta_3) = \frac{1}{2}\cos^2(50^\circ + 50^\circ)\cos^2(50^\circ + 50^\circ)$

34. In this case, we replace $I_0 \cos^2 70^\circ$ by $\frac{1}{2}I_0$ as the intensity of the light after passing through the first polarizer. Therefore,

$$I_f = \frac{1}{2} I_0 \cos^2(90^\circ - 70^\circ) = \frac{1}{2} (43 \text{ W/m}^2)(\cos^2 20^\circ) = 19 \text{ W/m}^2.$$

- 36. (a) The fraction of light that is transmitted by the glasses is
 - - $\frac{I_f}{I_0} = \frac{E_f^2}{E_0^2} = \frac{E_v^2}{E_v^2 + E_h^2} = \frac{E_v^2}{E_v^2 + (2.3E_v)^2} = 0.16.$

(b) Since now the horizontal component of \vec{E} will pass through the glasses,

 $\frac{I_f}{I_0} = \frac{E_h^2}{E_v^2 + E_h^2} = \frac{(2.3E_v)^2}{E_v^2 + (2.3E_v)^2} = 0.84.$

37. **THINK** A polarizing sheet can change the direction of polarization of the incident beam since it allows only the component that is parallel to its polarization direction to pass.

EXPRESS The 90° rotation of the polarization direction cannot be done with a single sheet. If a sheet is placed with its polarizing direction at an angle of 90° to the direction of polarization of the incident radiation, no radiation is transmitted.

We place the first sheet with its polarizing direction at some angle θ , between 0 and 90°, to the direction of polarization of the incident radiation. Place the second sheet with its polarizing direction at 90° to the polarization direction of the incident radiation. The transmitted radiation is then polarized at 90° to the incident polarization direction. The intensity is

ANALYZE (a) The 90° rotation of the polarization direction can be done with two sheets

$$I = I_0 \cos^2 \theta \cos^2 (90^\circ - \theta) = I_0 \cos^2 \theta \sin^2 \theta$$

where I_0 is the incident radiation. If θ is not 0 or 90°, the transmitted intensity is not zero

(b) Consider n sheets, with the polarizing direction of the first sheet making an angle of θ = 90°/n relative to the direction of polarization of the incident radiation. The polarizing direction of each successive sheet is rotated 90°/n in the same sense from the polarizing direction of the previous sheet. The transmitted radiation is polarized, with its direction of polarization making an angle of 90° with the direction of polarization of the incident radiation. The intensity is

$$I = I_0 \cos^{2n}(90^{\circ}/n).$$

We want the smallest integer value of n for which this is greater than $0.60I_0$. We start with n = 2 and calculate $\cos^{2n}(90^{\circ}/n)$. If the result is greater than 0.60, we have obtained the solution. If it is less, increase n by 1 and try again. We repeat this process, increasing n by 1 each time, until we have a value for which $\cos^{2n}(90^{\circ}/n)$ is greater than 0.60. The first one will be n = 5.

LEARN The intensities associated with n = 1 to 5 are:

$$\begin{split} I_{n-1} &= I_0 \cos^2(90^\circ) = 0 \\ I_{n-2} &= I_0 \cos^4(45^\circ) = I_0 / 4 = 0.25 I_0 \\ I_{n-3} &= I_0 \cos^6(30^\circ) = 0.422 I_0 \\ I_{n-4} &= I_0 \cos^8(22.5^\circ) = 0.531 I_0 \\ I_{n-5} &= I_0 \cos^{10}(18^\circ) = 0.605 I_0 \end{split}$$

39. (a) Since the incident light is unpolarized, half the intensity is transmitted and half is absorbed. Thus the transmitted intensity is $I = 5.0 \text{ mW/m}^2$. The intensity and the electric field amplitude are related by $I = E_m^2 / 2\mu_0 c$, so

$$E_m = \sqrt{2\mu_0 cI} = \sqrt{2(4\pi \times 10^{-7} \text{ H/m})(3.00 \times 10^8 \text{ m/s})(5.0 \times 10^{-3} \text{ W/m}^2)}$$

= 1.9 V/m.

(b) The radiation pressure is $p_r = I_a/c$, where I_a is the absorbed intensity. Thus

$$p_r = \frac{5.0 \times 10^{-3} \text{ W/m}^2}{3.00 \times 10^8 \text{ m/s}} = 1.7 \times 10^{-11} \text{ Pa.}$$

41. As the polarized beam of intensity I_0 passes the first polarizer, its intensity is reduced to $I_0 \cos^2 \theta$. After passing through the second polarizer, which makes a 90° angle with the first filter, the intensity is $I = (I_0 \cos^2 \theta) \sin^2 \theta = I_0 / 10$

which implies
$$\sin^2\theta\cos^2\theta = 1/10$$
, or $\sin\theta\cos\theta = \sin2\theta/2 = 1/\sqrt{10}$. This leads to $\theta = 70^\circ$ or 20° .

45. Note that the normal to the refracting surface is vertical in the diagram. The angle of refraction is $\theta_2 = 90^\circ$ and the angle of incidence is given by $\tan \theta_1 = L/D$, where D is the height of the tank and L is its width. Thus

$$\theta_1 = \tan^{-1} \left(\frac{L}{D} \right) = \tan^{-1} \left(\frac{1.10 \text{ m}}{0.850 \text{ m}} \right) = 52.31^\circ.$$

The law of refraction yields

$$n_1 = n_2 \frac{\sin \theta_2}{\sin \theta} = (1.00) \left\{ \frac{\sin 90^\circ}{\sin 52.31^\circ} \right\} = 1.26,$$

where the index of refraction of air was taken to be unity.

47. The law of refraction states

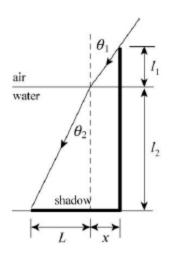
We take medium 1 to be the vacuum, with $n_1 = 1$ and $\theta_1 = 32.0^{\circ}$. Medium 2 is the glass, with $\theta_2 = 21.0^{\circ}$. We solve for n_2 :

 $n_1 \sin \theta_1 = n_2 \sin \theta_2$.

 $n_2 = n_1 \frac{\sin \theta_1}{\sin \theta_2} = (1.00) \left| \frac{\sin 32.0^{\circ}}{\sin 21.0^{\circ}} \right| = 1.48.$

55. THINK Light is refracted at the air-water interface. To calculate the length of the shadow of the pole, we first calculate the angle of refraction using the Snell's law.

EXPRESS Consider a ray that grazes the top of the pole, as shown in the diagram below.



Here $\theta_1 = 90^\circ - \theta = 90^\circ - 55^\circ = 35^\circ$, $\ell_1 = 0.50$ m, and $\ell_2 = 1.50$ m. The length of the shadow is d = x + L.

ANALYZE The distance x is given by

$$x = \ell_1 \tan \theta_1 = (0.50 \text{ m}) \tan 35^\circ = 0.35 \text{ m}.$$

According to the law of refraction, $n_2 \sin \theta_2 = n_1 \sin \theta_1$. We take $n_1 = 1$ and $n_2 = 1.33$ (from Table 33-1). Then,

$$\theta_2 = \sin^{-1} \left(\frac{\sin \theta_1}{n_2} \right) = \sin^{-1} \left(\frac{\sin 35.0^{\circ}}{1.33} \right) = 25.55^{\circ}.$$

L is given by

$$L = \ell_2 \tan \theta_2 = (1.50 \text{ m}) \tan 25.55^\circ = 0.72 \text{ m}.$$

Thus, the length of the shadow is d = 0.35 m + 0.72 m = 1.07 m.

LEARN If the pole were empty with no water, then $\theta_1 = \theta_2$ and the length of the shadow would be

$$d' = \ell_1 \tan \theta_1 + \ell_2 \tan \theta_1 = (\ell_1 + \ell_2) \tan \theta_1$$

by simple geometric consideration.

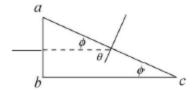
59. **THINK** Total internal reflection happens when the angle of incidence exceeds a critical angle such that Snell's law gives $\sin \theta_2 > 1$.

EXPRESS When light reaches the interfaces between two materials with indices of refraction n_1 and n_2 , if $n_1 > n_2$, and the incident angle exceeds a critical value given by

$$\theta_c = \sin^{-1}\left(\frac{n_2}{n_1}\right),\,$$

then total internal reflection will occur.

In our case, the incident light ray is perpendicular to the face ab. Thus, no refraction occurs at the surface ab, so the angle of incidence at surface ac is $\theta = 90^{\circ} - \phi$, as shown in the figure below.



ANALYZE (a) For total internal reflection at the second surface, $n_g \sin (90^\circ - \phi)$ must be greater than n_a . Here n_g is the index of refraction for the glass and n_a is the index of refraction for air. Since $\sin (90^\circ - \phi) = \cos \phi$, we want the largest value of ϕ for which $n_g \cos \phi \ge n_a$. Recall that $\cos \phi$ decreases as ϕ increases from zero. When ϕ has the largest value for which total internal reflection occurs, then $n_g \cos \phi = n_a$, or

$$\phi = \cos^{-1} \left[\frac{n_a}{n_o} \right] = \cos^{-1} \left[\frac{1}{1.52} \right] = 48.9^{\circ}.$$

The index of refraction for air is taken to be unity.

(b) We now replace the air with water. If $n_w = 1.33$ is the index of refraction for water, then the largest value of ϕ for which total internal reflection occurs is

$$\phi = \cos^{-1} \left[\frac{n_w}{n_g} \right] = \cos^{-1} \left[\frac{1.33}{1.52} \right] = 29.0^{\circ}.$$

LEARN Total internal reflection cannot occur if the incident light is in the medium with lower index of refraction. With $\theta_c = \sin^{-1}(n_2/n_1)$, we see that the larger the ratio n_2/n_1 , the larger the value of θ_c .

62. (a) Reference to Fig. 33-24 may help in the visualization of why there appears to be a "circle of light" (consider revolving that picture about a vertical axis). The depth and the radius of that circle (which is from point a to point f in that figure) is related to the tangent of the angle of incidence. The diameter of the circle in question is given by $d = 2h \tan \theta_c$. For water n = 1.33, so Eq. 33-47 gives $\sin \theta_c = 1/1.33$, or $\theta_c = 48.75^\circ$. Thus,

$$d = 2h \tan \theta_c = 2(2.00 \text{ m})(\tan 48.75^\circ) = 4.56 \text{ m}.$$

(b) The diameter d of the circle will increase if the fish descends (increasing h).

- 68. (a) We use Eq. 33-49: $\theta_R = \tan^{-1} n_w = \tan^{-1} (1.33) = 53.1^{\circ}$.
- (b) Yes, since n_w depends on the wavelength of the light.