

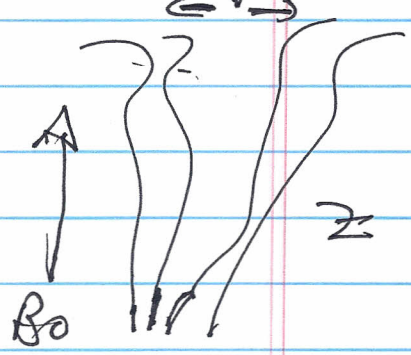
Transport - Stochastic Fields III

Misc

→ can describe field line separation in spectrum of magnetic perturbations

→ approach follows that of Richardson

now $l \leftrightarrow$ perp distance of lines



$$\frac{dl}{dz} \approx \frac{dl}{dz} \approx b_r$$

but $b_r \sim b_r(l)$ for Alfvénic turbulence, so

$$b_r(l) \sim \frac{E}{V_{A0}} l^{1/3}$$

$$\Rightarrow \frac{dl}{dz} \sim \frac{E}{V_{A0}} l^{1/3}$$

$$\frac{dl}{l^{1/3}} \sim \frac{E}{V_{A0}} dz$$

∴

$$\Rightarrow l^{2/3} \sim \frac{E^{1/3}}{v_{A0}} z$$

$$\therefore l_c^3 \sim \frac{E}{v_{A0}^3} z^3$$

and can define correlation length

$$\langle k_{\perp}^2 l_c^2 \rangle \sim 1 \quad \Rightarrow \quad \frac{E}{v_{A0}^3} k_{\perp}^2 l_c^3 \sim 1$$

$$\Rightarrow l_c \sim \left(\frac{E}{v_{A0}^3} \right)^{-1/3}$$

can compare to:

$$l_c = \left(\frac{k_{\perp}^2 D_m}{L_s^2} \right)^{-1/3}$$

Channels

- Particles $\left\{ \begin{array}{l} \text{why speak of heat transport} \\ \text{test particle diffusivity} \\ \text{why not particle transport?} \end{array} \right.$

Note: The 'point' of stochastic fields is $v_L \sim v_{the} \Delta M$
 \hookrightarrow electron speed.

Useful to:

\rightarrow consider kinetic eqn.

\rightarrow distinguish between:

a) self-consistent case

\tilde{B} produced by \tilde{J}_u in plasma

(i.e. EM instability - Ampere's Law)

b) \tilde{B} produced by external means (coil)

Mag

$$\frac{\partial f}{\partial t} + v_{||} \sigma_{||} f + v_{||} \frac{\partial B}{\partial B_0} \cdot \nabla_{\perp} f + \dots = 0$$

then integrating,

$$\frac{\partial \langle N_e \rangle}{\partial t} + \nabla_{\perp} \cdot \left\langle \frac{d^2 B_0}{B_0} \frac{\tilde{J}_{\perp}}{(-\text{Nobel})} \right\rangle + \dots = 0$$

but

$$\begin{aligned} \nabla_{\perp}^2 A_{\perp} &= -\frac{4\pi}{c} \tilde{J}_{\perp} \\ &= -\frac{4\pi}{c} (\tilde{J}_{\perp e} + \tilde{J}_{\perp i}) \end{aligned}$$

$$\tilde{J}_{\perp e} = \frac{-c}{4\pi} \nabla_{\perp}^2 A - \tilde{J}_{\perp i}$$

$$= \frac{-c}{4\pi} \nabla_{\perp}^2 A_{\perp} - \text{Nobel } \hat{V}_{\perp i}$$

↓
den flow
=

so

$$\frac{\partial \langle N_e \rangle}{\partial t} + \nabla_{\perp} \cdot \left\langle \frac{d^2 B_0}{B_0} \left(\frac{-c}{4\pi} \nabla_{\perp}^2 \hat{A}_{\perp} - \text{Nobel } \hat{V}_{\perp i} \right) \right\rangle + \dots = 0$$

$$d^2 B_0 = \nabla_{\perp}^2 \hat{A}_{\perp}$$

$$\Rightarrow \frac{d\langle Ne \rangle}{dt} + \frac{\partial}{\partial r} \left\langle \nabla_0 A_{11} \left(-\frac{e}{4\pi} \nabla_{\perp}^2 \hat{A}_{11} - \text{odd } V_{11} \right) \right\rangle = 0$$

$$\frac{d\langle Ne \rangle}{dt} = \frac{e}{4\pi} \frac{\partial}{\partial r} \left\langle \overset{\textcircled{1}}{(\nabla_0 \hat{A}_{11}) (\nabla_{\perp}^2 \hat{A}_{11})} \right\rangle + \text{odd} \left\langle \overset{\textcircled{2}}{\partial R_n \hat{V}_{11}} \right\rangle$$

①: Taylor identity calculation ↓

$$\begin{aligned} & \left\langle (\nabla_0 \hat{A}_{11}) (\partial_r^2 \hat{A}_{11} + \nabla_{\perp}^2 \hat{A}_{11}) \right\rangle \\ & \quad \text{odd} \\ & = \left\langle (\nabla_0 \hat{A}_{11}) (\partial_r^2 \hat{A}_{11}) \right\rangle \\ & = \left\langle \partial_r (\nabla_0 \hat{A}_{11}) (\partial_r \hat{A}_{11}) \right\rangle - \left\langle \nabla_0 (\partial_r \hat{A}_{11}) (\partial_r \hat{A}_{11}) \right\rangle \\ & \quad \text{odd} \\ & = \partial_r \left\langle (\nabla_0 \hat{A}_{11}) (\partial_r \hat{A}_{11}) \right\rangle \\ & = -\partial_r \left\langle (\partial_r \hat{B}_0) (\partial_r \hat{B}_0) \right\rangle \end{aligned}$$

Magnetic stress.

$$\frac{\partial \langle n_e \rangle}{\partial t} = - \frac{c}{4\pi} \frac{\partial n}{\partial t} \langle (d^2 B_r)(d^2 B_\theta) \rangle$$

↳ magnetic stress

$$+ n_0 k_B \langle d^2 B_r \hat{v}_{||} \rangle$$

↳ "magnetic flutter" of parallel ion flow.

Points:

- no explicit dependence on small electron inertia; i.e. $1/m_e$
- relation of magnetic stress to electron particle transport due to stochastic fields
- i.e. suggests effect on zonal flows as zonal flows are fundamentally charge transport effect.
- fluctuating parallel ion flows + magnetic tilt \Rightarrow change electron density.

N.B. For external field, key is to calculate

- δB_r in plasma
- \tilde{v}_{\perp} in plasma.

→ Re: zonal flows

Point: Magnetic wandering impacts zonal flows via charge

Result: - Fluctuations → DW model
i.e. H-M jets

- zonal mode: charge separation, radially

$$\nabla \cdot \underline{J} = 0$$

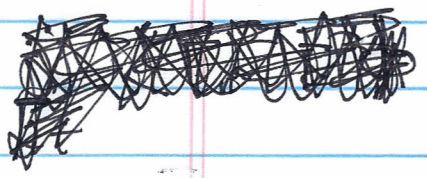
usual e.s; $\nabla_r \sigma_{pol} = 0$

$$\Rightarrow \frac{\partial}{\partial t} \nabla_{\perp}^2 \phi + \underline{v} \cdot \nabla \sigma^2 \phi = 0$$

linear pol. drift

NL pol. drift

σ_z , für $\langle \rangle \rightarrow$ zonal symmetry



$$\frac{\partial}{\partial t} \langle \nabla_r^2 \phi \rangle + \text{div} \langle \tilde{v}_r \nabla_r^2 \phi \rangle + \dots = 0$$

\downarrow
polarization
charge Q_{pol}

\downarrow
- Flux of
polarization
charge
(i.e. net difference cond
electrons)

- related to Reynolds
stress by Taylor
identity

\rightarrow compute via
modulation

$\frac{\partial}{\partial t}$, can view:

$$\frac{\partial}{\partial t} \langle Q_{pol} \rangle + \text{div} \langle \tilde{v}_r Q_{pol} \rangle = 0$$

$\int \rightarrow$

$$\Gamma_{charge} = \left(\langle \tilde{v}_r \tilde{n}_e \rangle + \langle \tilde{v}_r \tilde{Q}_{pol} \rangle \right) |_{r_0} - \left(\langle \tilde{v}_r \tilde{n}_e \rangle \right) |_{r_0}$$

$$\hat{n}_c = \hat{n}_e \Rightarrow \nabla \cdot \vec{J}_{charge} = 0 \quad \checkmark$$

But magnetic perturbation ~~induces~~ induce new means charge transport

$$\nabla \cdot \vec{J} = 0$$

$$\nabla_{\perp} \cdot \vec{J}_{\perp} + \nabla_{\parallel} \vec{J}_{\parallel} = 0$$

$$\nabla_{\parallel} = \nabla_{\parallel}^{(0)} + \frac{d\vec{B}}{B_0} \cdot \nabla_{\perp}$$

∴

$$\frac{\partial}{\partial t} \langle \nabla^2 \phi \rangle + \langle \vec{v} \cdot \nabla \nabla^2 \phi \rangle = \langle \vec{B} \cdot \nabla \vec{J}_{\parallel} \rangle$$

observed from 2D MHD

↓

non pot. charge transport
⇒ advection

↑

current flow along tilted lines

(n.b. current: electrons - ions)

|||

$$\Delta_{\perp} \langle \nabla_{\perp}^2 \psi \rangle = -\Delta_{\perp} \left\{ \langle \tilde{v}_r \tilde{v}_{\perp}^2 \rangle - \langle \tilde{R}_r \tilde{v}_{\perp} \rangle \right\}$$

so ~~current~~ current flow along tilted lines acts to affect charge balance

⇒ enters effect on Z.F.

Notes:

- sign not clear a priori
- for ω Alfvénic fluctuations \Rightarrow damps
i.e. acts on e_{\perp}
- obviously related to electron particle transport.

~> What Happens for $Ku \gg 1$?

- Recall:

~ patch fields

$$Ku \sim \frac{l_{ac}}{\Delta_{\perp}} \delta B / B_0 \sim \frac{l_{ac} \delta B / B_0}{\Delta |k_{\perp}|}$$

$$\sim l_{ac} / l_{NL} \sim l_{ac} / l_{NL}$$

ratio of autocorrelation to NL mixing length

Large $Ku \rightarrow$ NL scatt. processes control time/sp. scale.

~ flow

$$Ku \sim \frac{l_{ac} \tilde{v}}{\Delta} \sim \frac{l_{ac} \tilde{v}}{\Delta} / \frac{l_{ac} \tilde{v}}{\Delta} \sim \frac{l_{ac} \tilde{v}}{\Delta}$$

ion

de collisional DW's:

$$\tau_{ac} \sim (\Delta |k_{\perp} k_{\parallel}|)^{-1}$$

$$Ku \sim \frac{l_{ac}}{\Delta} (\Delta |k_{\perp} k_{\parallel}|)$$

2D GC Plasma - Simple / Compelling Example.

so Pooh:

$$-D_{\perp} \sim \int_0^{\infty} dt \langle \tilde{V}(t) \tilde{V}(0) \rangle$$

Different coeff of integral of correlation (i.e. time history \rightarrow resp)

$$\sim \int_0^{\infty} dt \sum_{\mathbf{n}} \tilde{V}_{\mathbf{n}}^2 R(t)$$

Memory fctn

Diffusion coefficient as integral of correlation function

$$R(t) = e^{-c(\omega - km v_{th})t} e^{-T/T_{\perp}}$$

From up

$km \gg 1$ limit corresponds to:

$$\rightarrow km \rightarrow \infty$$

$$\rightarrow \omega \rightarrow \infty$$

} \Rightarrow

time integral controlled by nonlinear scattering, not wave packet dispers

\Rightarrow 2D GC Plasma / Fluid:

e.e.
$$\frac{\partial \phi}{\partial t} + \nabla \phi \times z \cdot \nabla \phi = \nabla_{\perp}^2 \phi$$

$$\nabla_{\perp}^2 \phi = -4\pi \rho$$

(Taylor + McNamee)

\rightarrow 2D Fluid

\rightarrow GC Plasma.

Then generally,

$$D_{\perp} \approx \int_0^{\infty} d\tau \sum_{\mathbf{n}} |\tilde{W}_{\mathbf{n}}|^2 e^{-i\mathbf{k} \cdot \mathbf{r}_0} e^{+i\mathbf{k} \cdot \mathbf{r}(\tau)}$$

$$\text{but } \mathbf{r}(\tau) = \mathbf{r}_0 + \mathbf{dr}(\tau)$$

only evolution of \mathbf{r}_0

$d\mathbf{r} \rightarrow$ stochastic

\rightarrow need ensemble average

$$D_{\perp} \approx \int_0^{\infty} d\tau \sum_{\mathbf{n}} |\tilde{W}_{\mathbf{n}}|^2 \langle e^{i\mathbf{k} \cdot \mathbf{dr}(\tau)} \rangle d\tau$$

$$\approx \int_0^{\infty} d\tau \sum_{\mathbf{n}} |\tilde{W}_{\mathbf{n}}|^2 e^{-k_{\perp}^2 D_{\perp} \tau} d\tau$$

turbulence/scat itself controls correl. time.

e.g.

$$\langle e^{i\mathbf{k} \cdot \mathbf{dr}(\tau)} \rangle \approx \left\langle \left(1 + \frac{i\mathbf{k} \cdot \mathbf{dr}(\tau)}{1} - \frac{(\mathbf{k} \cdot \mathbf{dr}(\tau))^2}{2} + \dots \right) \right\rangle$$

$$\approx \left\langle \left(1 - \frac{k_{\perp}^2 D_{\perp} \tau}{2} \right) \right\rangle$$

$$\approx \left\langle \left(1 - k_{\perp}^2 D_{\perp} \tau \right) \right\rangle$$

$$\approx e^{-k_{\perp}^2 D_{\perp} \tau}$$

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$$D_{\perp} = \int_0^{\infty} dt \sum_{\underline{n}} |\tilde{W}_{\underline{n}}|^2 e^{-k_{\perp}^2 D_{\perp} t} dt$$

$$= \sum_{\underline{n}} |\tilde{W}_{\underline{n}}|^2 \frac{1}{k_{\perp}^2 D_{\perp}}$$

↑

can immediately note:
~~...~~
 - D_{\perp} controls γ_c
 - conservation of ρ
 $\int \rho d^3x \rightarrow \rho_{\text{avg}} \gamma_c$
 at large scale

N.B. \rightarrow note recursive structure of diffusivity!

$\rightarrow \gamma_c$ in integral set by scattering $\sim 1/k_{\perp}^2 D$, k_{\perp}^2 from cons.

$\Rightarrow \nu \sim \frac{c}{R} \underline{E} \times \underline{z} \rightarrow D_{\perp} \sim 1/B_0 \Rightarrow$ Bohm \propto

$(D_{\perp})^2 \approx \sum_{\underline{n}} |\tilde{W}_{\underline{n}}|^2 / k_{\perp}^2$

\rightarrow recursive definition

But 2D, assuming symmetric spectrum,

$$D_{\perp}^2 \approx \int dk_{\perp} \frac{1}{k_{\perp}} |\tilde{W}_{\underline{n}}|^2 / k_{\perp}^2$$

⇒

$$D_L \sim \frac{c^3}{B^2} \int d\omega_L \frac{|E_{\omega_L}|^2}{\omega_L}$$

$$D_L \sim \frac{c}{B_0} \left(\int d\omega_L \frac{|E_{\omega_L}|^2}{\omega_L} \right)^{1/2}$$

\downarrow
 $\sim 1/B_0$

⇒ Diffusivity

Now, can explain different spectra:

~ thermal equilibrium: (obs 7PM)
 ~ thermal

$$|E_{\omega}|^2 = \frac{4\pi}{l} \frac{k_B T}{(1 + k_D^2 \lambda_D^2)}$$

$e/l \rightarrow$ charge/length. \rightarrow Debye screening

$$= \frac{4\pi (e)}{e} \frac{k_B T}{(1 + k_D^2 \lambda_D^2)}$$

$$D_L \sim \frac{c}{B_0} \left(\int_{1/L}^{1/\lambda_D} d\omega_L \left[\frac{4\pi (e)}{e} \frac{k_B T}{(1 + k_D^2 \lambda_D^2)} \right] \omega_L \right)^{1/2}$$

$$\sim \frac{c k_B T}{e B} \left[(m^2)^{-1} \ln(L_0 \lambda) \right]^{1/2}$$

110

$$D_{\perp} \sim D_B \left[(v_A^2)^{-1} \ln(L_0/\lambda) \right]^{1/2}$$

- recovers basic Bohm scaling, even from thermal fluctuations
- scales (weakly) with $L_0 \Rightarrow$ not local, or intensive
- \Rightarrow simple example of "non-locality"
- "non-locality" appears from "slow mode"
 - i.e. $1/\nu \rightarrow \infty \sim k_{\perp}^2 D_{\perp}$ as $k_{\perp}^2 \rightarrow 0$
- if shear flow:

$$RCH = \int_0^{\infty} dt \exp\left[\nu(\omega - k_0 v_0 t) - t/\tau_c\right]$$

Interesting to note:

- can consider diffusion due random array charges ("rods")

For spectrum:

$$\underline{\underline{\underline{\nabla \cdot \underline{E}}}} = 4\pi \rho$$

$$= \frac{4\pi}{l} \sum_j \rho_j \delta(\underline{x} - \underline{x}_j)$$

$$i\underline{k} \cdot \underline{E}_k = \left(\frac{4\pi}{l} \right) \sum_j \rho_j e^{-i\underline{k} \cdot \underline{x}_j}$$

symmetric distribution \Rightarrow $\left. \begin{array}{l} \text{- random array} \\ \text{discrete charges} \end{array} \right\}$

$$|\underline{E}_k|^2 = \frac{1}{k_{\perp}^2} \left(\frac{4\pi}{l} \right)^2 \left\langle \sum_{j,j'} \rho_j \rho_{j'} e^{i\underline{k} \cdot (\underline{x}_j - \underline{x}_{j'})} \right\rangle$$

$$= \frac{16\pi^2 n \rho^2}{k_{\perp}^2 l^2}$$

$$\hookrightarrow \frac{1}{k_{\perp}^2}$$

$$D_{\perp}^2 \sim \frac{c^2}{B_0^2} \int dk_{\perp} \frac{16\pi^2 n \rho^2}{k_{\perp}^2 l^2} k_{\perp}$$

$$\sim \frac{c}{B_0} \int_{k_{\min}}^{k_{\max}} \frac{16\pi^2 n \rho^2}{k_{\perp}^3 l^2} dk_{\perp} \sim \frac{c^2}{B_0^2} \left(\frac{16\pi^2 n \rho^2}{l^2} \right) \frac{1}{k_{\perp, \min}^2}$$

$$k_{L \text{ min}} \sim 1 / L_0$$

↓
system size

$$D_k \sim \frac{c}{B_0} 4\pi \left(\frac{1}{f} \right)^2 L_0$$

↑
strong dependence on system size.