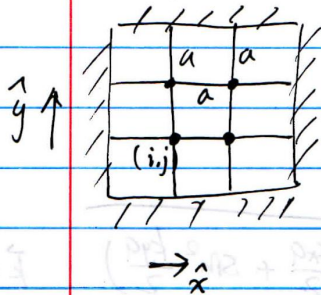


Yinming Shao

1. (FW 4.16)



(a) Let η_{ij} denote the small transverse displacements of the masses.

$$KE = \frac{1}{2} m \sum_{i=1}^N \sum_{j=1}^N \dot{\eta}_{ij}^2$$

$$PE = \frac{1}{2} k [(\Delta^x \eta)^2 + (\Delta^y \eta)^2] \quad k \rightarrow \frac{\tau}{a}$$

$$\Delta^x \eta = \sum_{i=1}^N \sum_{j=2}^N (\eta_{ij+1} - \eta_{ij})^2$$

$$\Delta^y \eta = \sum_{j=1}^N \sum_{i=2}^N (\eta_{i+j} - \eta_{ij})^2$$

$$L = T - V$$

$$= \frac{1}{2} m \sum_{i=1}^N \sum_{j=1}^N \dot{\eta}_{ij}^2 - \frac{\tau}{2a} \left[\sum_{i=1}^N \sum_{j=2}^N (\eta_{ij+1} - \eta_{ij})^2 + \sum_{j=1}^N \sum_{i=2}^N (\eta_{i+j} - \eta_{ij})^2 \right]$$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\eta}_{ij}} = m \ddot{\eta}_{ij}$$

$$\frac{\partial L}{\partial \eta_{ij}} = -\frac{\tau}{a} [-(\eta_{ij+1} - \eta_{ij}) + (\eta_{ij} - \eta_{ij-1}) - (\eta_{i+j} - \eta_{ij}) + (\eta_{ij} - \eta_{i-1+j})]$$

$$\text{EOM: } m \ddot{\eta}_{ij} + \frac{4\tau}{a} \eta_{ij} - \frac{\tau}{a} (\eta_{ij+1} + \eta_{ij-1} + \eta_{i+j} + \eta_{i-1+j}) = 0 \quad (*)$$

Let $x_i = ia$, $y_j = ja$

trial solution: $\eta_{ij} = \eta(x_i, y_j, t) = A e^{i(k_x x_i + k_y y_j - \omega t)}$

plug into EOM (*):

$$-m\omega^2 \eta + \frac{4\tau}{a} \eta - \frac{\tau}{a} (e^{ik_x a} + e^{-ik_x a} + e^{ik_y a} + e^{-ik_y a}) \eta = 0$$

$$\text{So: } \omega^2 = \frac{4\tau}{ma} - \frac{\tau}{ma} (2\cos k_x a + 2\cos k_y a)$$

$$= \frac{2\tau}{ma} [2 - \cos k_x a - \cos k_y a]$$

$$= \frac{4\tau}{ma} \left(\sin^2 \frac{k_x a}{2} + \sin^2 \frac{k_y a}{2} \right)$$

$$\text{Dispersion relation: } \omega(\vec{k}) = \sqrt{\frac{4\tau}{ma} \left(\sin^2 \frac{k_x a}{2} + \sin^2 \frac{k_y a}{2} \right)} \quad \vec{k} = k_x \hat{x} + k_y \hat{y}$$

(b) Continuum mass distribution.

$$\eta(x_i, y_j, t) \rightarrow u(x, y, t), \quad a \rightarrow 0$$

$$\text{mass density of string} = \frac{m}{a} = \sigma = \text{const.}$$

$$\text{EOM: } \ddot{\eta}_{ij} = \frac{\tau}{\sigma} \left[\frac{1}{a} \left(\frac{\eta_{i+1,j} - \eta_{i,j}}{a} - \frac{\eta_{i,j} - \eta_{i,j-1}}{a} \right) + \frac{1}{a} \left(\frac{\eta_{i+1,j} - \eta_{i,j}}{a} - \frac{\eta_{i,j} - \eta_{i,j-1}}{a} \right) \right]$$

$$\frac{\eta_{i+1,j} - \eta_{i,j}}{a} = \frac{\eta_{i+1,j} - \eta_{i,j}}{x_{i+1} - x_i} \xrightarrow{a \rightarrow 0} \frac{\partial u}{\partial x}$$

$$\frac{\eta_{i+1,j} - \eta_{i,j}}{a} = \frac{\eta_{i+1,j} - \eta_{i,j}}{x_{i+1} - x_i} \xrightarrow{a \rightarrow 0} \frac{\partial u}{\partial x} \quad \ddot{\eta}_{ij} = \frac{\partial^2 u}{\partial t^2}$$

$$\Rightarrow \boxed{\frac{1}{c^2} \frac{\partial^2 u}{\partial t^2} = \frac{\partial^2 u}{\partial x^2} + \frac{\partial^2 u}{\partial y^2}} \rightarrow \text{2D-wave Eq.}$$

$$c^2 = \tau/\sigma$$

Dispersion relation:

$$\omega^2 = \frac{4\tau}{ma} \cdot \frac{1}{4} (k_x^2 a^2 + k_y^2 a^2)$$

$$= \frac{4\tau a}{m} (k_x^2 + k_y^2) = c^2 k^2$$

$$\boxed{\omega = c |\vec{k}|} \quad \text{travelling waves.}$$

FW 4.16

(c) $\omega = \sqrt{\frac{c}{\rho}} |\vec{k}| = c |\vec{k}| \rightarrow$ isotropic
 ω only depends on the length of the wave vector ($\vec{k} = k_x \hat{x} + k_y \hat{y}$)

$\omega = \sqrt{\frac{4c}{ma}} \sqrt{\sin^2\left(\frac{k_x a}{2}\right) + \sin^2\left(\frac{k_y a}{2}\right)} \rightarrow$ anisotropic
 ω depends on direction of \vec{k} in the xy plane.

(d) Finite system, find exact normal-mode frequencies.

Let $\eta(x_i, y_j, t) = B e^{-i\omega t} \left[e^{ik_x x_i + ik_y y_j} - e^{-ik_x x_i + ik_y y_j} - e^{-ik_x x_i - ik_y y_j} + e^{ik_x x_i - ik_y y_j} \right]$

So that $\eta(0, y_j) = 0$ & $\eta(x_i, 0) = 0$ are automatically satisfied.

another two b.c. $\eta((N+1)a, y_j) = 0$ ①

$\eta(x_i, (N+1)a) = 0$ ②

$$\text{①} \rightarrow \left(e^{ik_x(N+1)a} - e^{-ik_x(N+1)a} \right) \left(e^{ik_y y_j} - e^{-ik_y y_j} \right) = 0$$

valid for any y_j

$$\text{So: } \sin k_x (N+1)a = 0$$

$$k_x = \frac{n\pi}{(N+1)a} \quad (n=1, 2, \dots, N)$$

② \rightarrow

$$k_y = \frac{m\pi}{(N+1)a} \quad (m=1, 2, \dots, N)$$

$$\text{Normal mode freq. } \omega_{nm} = \sqrt{\frac{4c}{ma}} \sqrt{\sin^2 \frac{n\pi}{2(N+1)} + \sin^2 \frac{m\pi}{2(N+1)}}$$

Continuum limit:

$$\omega_{nm}^2 \approx \frac{4\tau}{ma} \left[\left(\frac{n\pi}{2(N+1)} \right)^2 + \left(\frac{m\pi}{2(N+1)} \right)^2 \right]$$

$$= \frac{\tau}{ma} a^2 \left[\left(\frac{n\pi}{Lx} \right)^2 + \left(\frac{m\pi}{Ly} \right)^2 \right] \quad Lx = (N+1)a = Ly$$

$$c^2 = \frac{\tau}{\rho}$$

$$= c^2 (k_x^2 + k_y^2)$$

(e) Discrete Lagrangian

$$L = \frac{m}{2a^2} \sum_{i=1}^N \sum_{j=1}^N a^2 \dot{\eta}_{ij}^2 - \frac{\tau}{2} \left[\sum_{i=1}^N \sum_{j=0}^N a^2 \left(\frac{\eta_{i,j+1} - \eta_{ij}}{a} \right)^2 + \sum_{j=1}^N \sum_{i=0}^N a^2 \left(\frac{\eta_{i+1,j} - \eta_{ij}}{a} \right)^2 \right]$$

Continuum limit: $N \rightarrow \infty$, $a \rightarrow 0$, $\eta_{ij} \rightarrow u(x, y, t)$

$\frac{m}{a^2} \rightarrow \mu$, mass density

$$\sum_{i=1}^N \sum_{j=1}^N a^2 \rightarrow \sum_i \Delta x_i \sum_j \Delta y_j = \int_0^L dx \int_0^L dy \quad L = (N+1)a$$

$$\text{So: } L = \frac{\mu}{2} \int_0^L \int_0^L \dot{u}^2 dx dy - \frac{\tau}{2} \int_0^L \int_0^L (u_x^2 + u_y^2) dx dy$$

$$\text{Lagrangian density: } \mathcal{L} = \frac{\mu}{2} \dot{u}^2 - \frac{\tau}{2} u_x^2 - \frac{\tau}{2} u_y^2$$

Plug into Lagrangian eq. $\frac{\partial \mathcal{L}}{\partial u} - \frac{\partial \mathcal{L}}{\partial t} - \frac{\partial \mathcal{L}}{\partial x} \frac{\partial \mathcal{L}}{\partial u_x} - \frac{\partial \mathcal{L}}{\partial y} \frac{\partial \mathcal{L}}{\partial u_y} = 0$

$$\Rightarrow \mu \ddot{u} - \tau u_{xx} - \tau u_{yy} = 0$$

$$\Rightarrow \frac{1}{c^2} \frac{\partial^2 u}{\partial t^2} = \frac{\partial^2 u}{\partial x^2} + \frac{\partial^2 u}{\partial y^2}$$

4.6 Normally for a central potential,

Derivada

$$L = \frac{1}{2} m (\dot{r}^2 + r^2 \dot{\phi}^2) - V(r)$$

$$r \rightarrow r_0 + \epsilon, \phi \rightarrow \Omega t + \delta \theta$$

$$L \approx \frac{1}{2} m (\dot{\epsilon}^2 + (r_0 + \epsilon)^2 (\Omega + \dot{\delta})^2) - V(r_0) - \frac{1}{2} V''(r_0) \epsilon^2$$

can change sign of potential and ignore

$$\frac{dL}{d\epsilon} - \frac{dL}{d\dot{\epsilon}} = 0 \Rightarrow m \ddot{\epsilon} - m(r_0 + \epsilon)(\Omega + \dot{\delta})^2 + V'(r_0) + V''(r_0) \epsilon = 0$$

$$\frac{dL}{d\delta} - \frac{dL}{d\dot{\delta}} = 0 \Rightarrow \frac{d}{dt} (m(r_0 + \epsilon)^2 (\Omega + \dot{\delta})) = 0$$

constant

Equations of Motion

At equilibrium, $\epsilon = \dot{\epsilon} = \ddot{\epsilon} = 0, \delta = \dot{\delta} = 0$ (no oscillations)

plugging in: $\frac{d}{dt} (m r_0^2 \Omega) = 0 \rightarrow m r_0^2 \Omega = \text{constant} \rightarrow m r^2 \dot{\phi} = \text{constant } \ell$
 $-m r_0 \Omega^2 + V'(r_0) = 0 \rightarrow \frac{dV}{dr}|_{r_0} = m r_0 \Omega^2 = m r \dot{\phi}^2$ for circular orbits

Rewriting equation of motion:

$$m \ddot{\epsilon} - \frac{\ell^2}{m r^3} + V'(r_0) + V''(r_0) \epsilon = 0$$

Expanding our second term:

$$m \ddot{\epsilon} - \frac{\ell^2}{m r_0^3} + \frac{3\ell^2}{m r_0^4} \epsilon + V'(r_0) + V''(r_0) \epsilon = 0 \quad \text{terms cancel}$$

$$m \ddot{\epsilon} + \left[V''(r_0) + \frac{3\ell^2}{m r_0^4} \right] \epsilon = 0$$

Condition for stability:

frequency $V''(r_0) + \frac{3\ell^2}{m r_0^4} > 0$

4.6 $E = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\phi}^2) + V(r)$ for a central potential

(b) Have $r \rightarrow r_0 + \delta r$, $\phi \rightarrow (\Omega + \delta\Omega)t$

$V_{\text{eff}}(r) = V(r) + \frac{l^2}{2mr^2}$, where $l = mr^2\dot{\phi} = \text{constant}$ for our case

So, $E = \frac{1}{2}m\delta\dot{r}^2 + V_{\text{eff}}(r_0 + \delta r)$, where $V_{\text{eff}}(r_0 + \delta r) \approx V_{\text{eff}}(r_0) + \frac{dV_{\text{eff}}(r_0)}{dr}\delta r + \frac{1}{2}\frac{d^2V_{\text{eff}}(r_0)}{dr^2}\delta r^2$
can give constant by changing zero of potential

$E \approx \frac{1}{2}m\delta\dot{r}^2 + \frac{1}{2}V_{\text{eff}}''\delta r^2$

$\hookrightarrow \omega^2 = \frac{d^2V_{\text{eff}}}{dr^2}\bigg|_{r_0} > 0$ for stability

$V_{\text{eff}}' = 0$ for a circular orbit

$V_{\text{eff}}' = V'(r) - \frac{l^2}{mr^3} = 0 \rightarrow \frac{dV}{dr} = \frac{l^2}{mr^3}$

$V_{\text{eff}}'' = V'' + \frac{3}{r^4}\frac{dV}{dr} = \frac{1}{r^3}\frac{d}{dr}\left(r^3\frac{dV}{dr}\right)\bigg|_{r_0} > 0$

(c) $V = -\lambda r^{-n}$
 $V' = n\lambda r^{-n-1}$

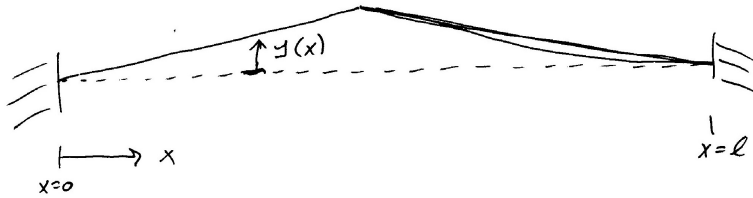
$\frac{1}{r^3}\frac{d}{dr}(r^3 n\lambda r^{-n-1}) = \frac{1}{r^3}\frac{d}{dr}(n\lambda r^{-n+2}) = n\lambda\frac{1}{r^3}(2-n)r^{-n+1} = n\lambda(2-n)r^{-n+1} > 0$

Given $\lambda > 0$, this is only true for $n < 2$ ✓

Tucker Ellefiot

7.1

$$a) y(x) = u(x, 0) = \begin{cases} \frac{2hx}{l} & 0 \leq x \leq \frac{1}{2}l \\ \frac{2h(l-x)}{l} & \frac{1}{2}l \leq x \leq l \end{cases}$$



$y(x)$ can be decomposed into normal modes

$$y(x) = \sum_{n=1}^{\infty} b_n \sin \frac{n\pi x}{l}$$

$$b_n = \frac{2}{l} \int_0^l y(x) \sin\left(\frac{n\pi x}{l}\right) dx$$

$$= \frac{2}{l} \int_0^{l/2} \frac{2hx}{l} \sin \frac{n\pi x}{l} dx + \frac{2}{l} \int_{l/2}^l \frac{2h(l-x)}{l} \sin \frac{n\pi x}{l} dx$$

$$\rightarrow b_n = \frac{4h}{l^2} \left(-\frac{lx}{n\pi} \cos \frac{n\pi x}{l} \Big|_0^{l/2} + \frac{l^2}{\pi^2 n^2} \sin \frac{n\pi x}{l} \Big|_0^{l/2} + \frac{lx}{n\pi} \cos \frac{n\pi x}{l} \Big|_{l/2}^l \right. \\ \left. - \frac{l^2}{n^2 \pi^2} \sin \frac{n\pi x}{l} \Big|_{l/2}^l - \frac{l^2}{n\pi} \cos \frac{n\pi x}{l} \Big|_{l/2}^l \right) \\ = \frac{8h}{n^2 \pi^2} \sin \frac{n\pi}{2}$$

$$= \begin{cases} \frac{8h}{n^2 \pi^2} (-1)^{\frac{n+1}{2}} & n \text{ odd} \\ 0 & n \text{ even} \end{cases}$$

$$dU_n = \frac{1}{2} T \left(\frac{dy_n}{dx} \right)^2 dx = \frac{T}{2} b_n^2 \left(\frac{\pi n}{l} \right)^2 \cos^2 \frac{n\pi x}{l} dx$$

$$U_n = \frac{T}{2} b_n^2 \left(\frac{\pi n}{l} \right)^2 \int_0^l \cos^2 \frac{n\pi x}{l} dx$$

$$= \frac{T}{2} b_n^2 \left(\frac{\pi n}{l} \right)^2 \left(\frac{x}{2} + \frac{\sin \left(2 \frac{n\pi x}{l} \right)}{4n\pi/l} \right) \Big|_0^l$$

$$= \begin{cases} \frac{16T \cdot h^2}{\pi^2 l \cdot n^2} & n \cdot \text{odd} \\ 0 & n \text{ even} \end{cases}$$

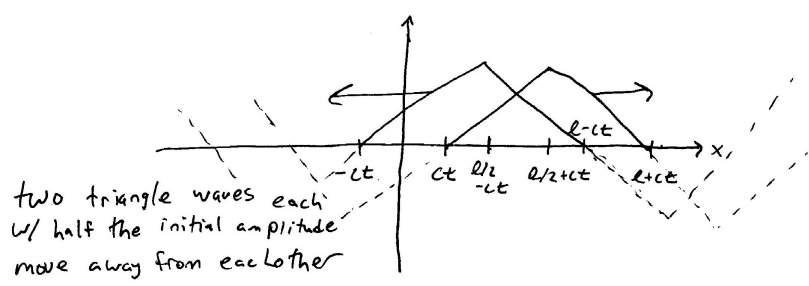
b) D'Alembert's solution:

$$u(x,t) = \frac{1}{2} [f(x-ct) + f(x+ct)] + \frac{1}{2c} \int_{x-ct}^{x+ct} g(\xi) d\xi$$

Initial conditions: $u(x,0) = f(x)$
 $\dot{u}(x,0) = g(x) = 0$

$$f(x-ct) = \begin{cases} \frac{2h(x-ct)}{l} & ct \leq x \leq l/2+ct \\ \frac{2h(l-x+ct)}{l} & ct + \frac{l}{2} \leq x \leq ct+l \end{cases}$$

$$f(x+ct) = \begin{cases} \frac{2h(x+ct)}{l} & -ct \leq x \leq l/2-ct \\ \frac{2h(l+ct-x)}{l} & \frac{l}{2}-ct \leq x \leq l-ct \end{cases}$$



BC's $u(0, t) = 0 \rightarrow f(-x) = -f(x)$

$u(l, t) = 0 \rightarrow f[l + (x-l)] = -f[l - (x-l)]$

Physics 200A Homework 6.4

Mark Derdzinski

December 2, 2013

Problem 4 Solution

Recall the Lagrangian density for a continuous string of length L with constant density μ and tension τ clamped at both ends:

$$\mathcal{L} = \frac{\mu}{2} y_t^2 - \frac{\tau}{2} y_x^2 \quad (1)$$

where y_t and y_x are the derivatives of the position $y(x, t)$ with respect to t or x . We want to express the Lagrangian, and eventually the Hamiltonian, in terms of fourier coefficients. We will expand $y(x, t)$ in the (complete) basis of spatial eigenfunctions

$$y(x, t) = \sum_{n=1}^{\infty} C_n \rho_n(x) \cos(\omega_n t + \phi_n) = \sum_{n=1}^{\infty} A_n(t) \rho_n(x) \quad (2)$$

Where the spatial eigenfunctions $\rho_n(x)$ are given by

$$\rho_n(x) = \sum_{n=1}^{\infty} \left(\frac{2}{L\mu}\right)^{\frac{1}{2}} \sin(k_n x) \quad (3)$$

Note our spatial eigenfunctions satisfy the orthonormality condition

$$\int_0^L \rho_n(x) \rho_m(x) dx = \delta_{nm} \quad (4)$$

We are now equipped to describe the system in terms of the time-dependent fourier coefficients. Suppressing x and t for clarity, the Lagrangian density expanded in fourier series becomes

$$\mathcal{L} = \frac{\mu}{2} \left[\sum_{n=1}^{\infty} \dot{A}_n \rho_n \right] \left[\sum_{m=1}^{\infty} \dot{A}_m \rho_m \right] - \frac{\tau}{2} \left[\sum_{n=1}^{\infty} A_n \frac{d\rho_n}{dx} \right] \left[\sum_{m=1}^{\infty} A_m \frac{d\rho_m}{dx} \right] \quad (5)$$

If we integrate over x to find the full Lagrangian, we can exploit the orthonormality of the ρ_n to simplify the product of sums:

$$L(A_n, \dot{A}_n, t) = \int_0^L \mathcal{L} dx = \frac{1}{2} \sum_{n=1}^{\infty} \left[\dot{A}_n^2 - \frac{\tau k_n^2}{\mu} A_n^2 \right] = \frac{1}{2} \sum_{n=1}^{\infty} \left[\dot{A}_n^2 - \omega_n^2 A_n^2 \right] \quad (6)$$

Where we have used the fact $\frac{\tau k_n^2}{\mu} = c^2 k_n^2 = \omega_n^2$. We are now equipped to find the Hamiltonian in terms of the fourier coefficients. The generalized momentum is

$$\pi_n = \frac{dL}{d\dot{A}_n} = \dot{A}_n \quad (7)$$

And so the Hamiltonian is given by

$$H(A_n, \dot{A}_n, t) = \sum_n^{\infty} \pi_n \dot{A}_n - L = \frac{1}{2} \sum_{n=1}^{\infty} [\pi_n^2 + \omega_n^2 A_n^2] \quad (8)$$

The Hamiltonian EOM for the fourier coefficients is now simply

$$\begin{aligned} \dot{\pi}_n &= -\frac{dH}{dA_n} \\ \implies \ddot{A}_n &= -\omega_n^2 A_n \end{aligned} \quad (9)$$

Note the Hamiltonian can be expressed in terms of

$$\begin{aligned} a_n^+ &= \pi_n + i\omega_n A_n \\ a_n^- &= \pi_n - i\omega_n A_n \\ \implies H &= \frac{1}{2} \sum_n^{\infty} a_n^+ a_n^- \end{aligned} \quad (10)$$

The a_n^+ and a_n^- are analogous to creation and annihilation operators for our purely classical system, where the zero-point energy is vanishing due to the classical commutator $[\pi_n, A_n] = 0$.

5) ~~the~~ $\frac{1}{2} \sigma \dot{f}^2 = T$ if $f(x,y)$ describes the surface

$$U = T \delta A = T (dS - dA), \quad dA = \hat{z} \cdot (\hat{n} ds)$$

where \hat{n} is given by $\nabla(\underline{z} - f(x,y)) = \frac{\hat{z} - \nabla f}{\sqrt{1 + (\nabla f)^2}}$ normal vector of the surface element

$$\Rightarrow dA = \frac{ds}{\sqrt{1 + (\nabla f)^2}} \Rightarrow \rho = T (\sqrt{1 + (\nabla f)^2} - 1)$$

density

$\rightarrow \mathcal{L} = \frac{1}{2} \sigma \dot{f}^2 - T (\sqrt{1 + (\nabla f)^2} - 1)$ and Lyrange eqs yield

$$\sigma f_{tt} = \nabla \cdot \left(\frac{T \nabla f}{\sqrt{1 + (\nabla f)^2}} \right) \quad \text{nonlinear wave eq.}$$

for small oscillations this reduces to

$$\underline{\sigma f_{tt} = T \nabla^2 f}$$

now consider $\mathcal{H} = \pi \dot{f} - \mathcal{L}$; $\pi = \frac{\partial \mathcal{L}}{\partial \dot{f}}$

$$\frac{d\mathcal{H}}{dt} = \pi \ddot{f} + \dot{\pi} \dot{f} - \frac{\partial \mathcal{L}}{\partial t} = \cancel{\pi \ddot{f}} + \dot{\pi} \dot{f} - \left(\frac{\partial \mathcal{L}}{\partial t} \dot{f} + \frac{\partial \mathcal{L}}{\partial f} \dot{f} + \frac{\partial \mathcal{L}}{\partial \nabla f} (\nabla \dot{f}) \right)$$

$$\rightarrow \dot{\pi} \dot{f} - \frac{\partial \mathcal{L}}{\partial f} \dot{f} - \frac{\partial \mathcal{L}}{\partial \nabla f} (\nabla \dot{f}) = \dot{\pi} \dot{f} - \cancel{\pi \ddot{f}} - \nabla \cdot \frac{\partial \mathcal{L}}{\partial \nabla f} \dot{f} - \frac{\partial \mathcal{L}}{\partial t} (\dot{f})$$

$\frac{d}{dt} \frac{\partial \mathcal{L}}{\partial \dot{f}} + \nabla \cdot \frac{\partial \mathcal{L}}{\partial \nabla f}$

$$\Rightarrow \frac{d\mathcal{H}}{dt} = -\nabla \cdot \frac{\partial \mathcal{L}}{\partial \dot{f}} \dot{f} - \frac{\partial \mathcal{L}}{\partial t} = \nabla \cdot \left(\frac{\partial \mathcal{L}}{\partial \dot{f}} \dot{f} \right)$$

$$\Rightarrow \frac{d\mathcal{H}}{dt} + \nabla \cdot \underbrace{\left(\frac{\partial \mathcal{L}}{\partial \dot{f}} \dot{f} \right)}_{\equiv S} = 0 \quad \text{Energy conservation}$$

$$\frac{d\mathcal{H}}{dt} + \nabla \cdot S = 0$$

for momentum conservation, consider $\frac{dS}{dt}$

$$\text{let } P_{\text{wave}} \equiv \frac{S}{v_{ph}^2} \quad , \quad v_{ph} \equiv \frac{\omega}{k} \quad . \quad \frac{dP}{dt} = \frac{1}{v_{ph}^2} \frac{dS}{dt}$$

$$= \frac{\omega}{T} \left[\partial_t \left(\frac{\partial \mathcal{L}}{\partial \dot{f}} \right) \dot{f} + \dot{f} \frac{\partial \mathcal{L}}{\partial \dot{f}} \right] ; \text{ use } \dot{f} = \frac{\omega}{k} \nabla^2 f \quad \& \quad \frac{\partial \mathcal{L}}{\partial \dot{f}} = -T \nabla f$$

$$\Rightarrow \frac{\omega}{T} \left[-T \partial_t (\nabla f) \dot{f} - T \nabla f \frac{\omega}{v_{ph}} \nabla^2 f \right]$$

$$= - \left[\nabla f \nabla^2 f + \sigma f \nabla f \right] = -\nabla \cdot \left[\frac{T(\nabla f)^2}{2} + \frac{\sigma(f)^2}{2} \right]$$

so we have $\underline{\underline{\epsilon}} \equiv \frac{T(\nabla f)^2}{2} + \frac{\sigma(f)^2}{2}$

$$\Rightarrow \frac{dP_{\text{wave}}}{dt} + \nabla \cdot \underline{\underline{\epsilon}} = 0 \quad \text{momentum conservation}$$

~~$$\underline{\underline{\epsilon}} = \frac{T}{2} \left(\frac{\partial f}{\partial x} \right)^2 + \frac{\sigma}{2} f^2$$~~

~~not~~ actually, have to be a little more careful
defining ϵ

$$S_x = f_t \frac{\partial \mathcal{L}}{\partial f_x} \quad \text{so} \quad \frac{d}{dt} S_x = f_{tt} \frac{d\mathcal{L}}{dt} + f_t \partial_t \frac{\partial \mathcal{L}}{\partial f_x}$$

$$\frac{d}{dt} S_x = \frac{T}{\sigma} (f_{xx} + f_{yy}) - T f_x + f_t (-T f_x)$$

$$= -\frac{T}{\sigma} (T(f_{xx} f_x + f_{yy} f_x) + \sigma f_t f_{xt})$$

$$= -\frac{T}{\sigma} \left(\partial_x \left(\frac{f_x^2 T}{2} + \frac{\sigma f_t^2}{2} \right) + T f_{yy} f_x \right)$$

$$= -\frac{T}{\sigma} \left[\frac{\partial}{\partial x} \left(\frac{T f_x^2}{2} + \frac{\sigma f_t^2}{2} - \frac{T f_y^2}{2} \right) + \frac{\partial}{\partial y} (T f_y f_x) \right]$$

$$\text{similarly, } \frac{dS_y}{dt} = -\frac{T}{\sigma} \left[\frac{\partial}{\partial y} \left(\frac{T f_y^2}{2} + \frac{\sigma f_t^2}{2} - \frac{T f_x^2}{2} \right) + \frac{\partial}{\partial x} (T f_x f_y) \right]$$

$$\rightarrow \underline{\epsilon} = \begin{pmatrix} T \left[\frac{f_y^2 - f_x^2}{2} \right] + \frac{\sigma f_t^2}{2} & T f_x f_y \\ T f_x f_y & \frac{T}{2} [f_y^2 - f_x^2] + \frac{\sigma f_t^2}{2} \end{pmatrix}$$

$$\& \text{ have } \partial_t P_{\text{wave}} + \nabla \cdot \underline{\epsilon} = 0$$

(2.10)

And if we want we can write

$$\partial_\mu T^{\mu\nu} = 0 \quad \text{where} \quad \partial_\mu = \left(\frac{1}{v_{ph}} \frac{\partial}{\partial t}, \frac{\partial}{\partial x}, \frac{\partial}{\partial y} \right)$$

$$T^{\mu\nu} = \begin{pmatrix} \mathcal{H} & S_x/v_{ph} & S_y/v_{ph} \\ S_x/v_{ph} & \epsilon_{xx} & \epsilon_{xy} \\ S_y/v_{ph} & \epsilon_{yx} & \epsilon_{yy} \end{pmatrix}$$