PHYSICS 210A: STATISTICAL PHYSICS HW ASSIGNMENT #2 SOLUTIONS

(1) Compute the density of states D(E,V,N) for a three-dimensional gas of particles with Hamiltonian $\hat{H} = \sum_{i=1}^{N} A |\mathbf{p}_i|^4$, where A is a constant. Find the entropy S(E,V,N), the Helmholtz free energy F(T,V,N), and the chemical potential $\mu(T,p)$.

Solution:

Let's solve the problem for a general dispersion $\varepsilon(p) = A|p|^{\alpha}$. The density of states is

$$D(E, V, N) = \frac{V^N}{N!} \int \frac{d^d p_1}{h^d} \cdots \int \frac{d^d p_N}{h^d} \, \delta(E - Ap_1^{\alpha} - \ldots - Ap_N^{\alpha}) \; .$$

The Laplace transform is

$$\begin{split} \widehat{D}(\beta, V, N) &= \frac{V^N}{N!} \bigg(\int \! \frac{d^d p}{h^d} \, e^{-\beta A p^\alpha} \bigg)^{\!N} \\ &= \frac{V^N}{N!} \bigg(\frac{\Omega_d}{h^d} \int \limits_0^\infty \! dp \, p^{d-1} \, e^{-\beta A p^\alpha} \bigg)^{\!N} \\ &= \frac{V^N}{N!} \bigg(\frac{\Omega_d \, \Gamma(d/\alpha)}{\alpha h^d A^{d/\alpha}} \bigg)^{\!N} \beta^{-Nd/\alpha} \; . \end{split}$$

Now we inverse transform, recalling

$$K(E) = \frac{E^{t-1}}{\Gamma(t)} \iff \widehat{K}(\beta) = \beta^{-t}$$
.

We then conclude

$$D(E,V,N) = \frac{V^N}{N!} \bigg(\frac{\Omega_d \, \Gamma(d/\alpha)}{\alpha h^d A^{d/\alpha}} \bigg)^N \frac{E^{\frac{Nd}{\alpha}-1}}{\Gamma(Nd/\alpha)}$$

and

$$\begin{split} S(E,V,N) &= k_{\mathrm{B}} \ln D(E,V,N) \\ &= N k_{\mathrm{B}} \ln \left(\frac{V}{N}\right) + \frac{d}{\alpha} \, N k_{\mathrm{B}} \, \ln \! \left(\frac{E}{N}\right) + N k_{\mathrm{B}} a_0 \; , \end{split}$$

where a_0 is a constant, and we take the thermodynamic limit $N \to \infty$ with V/N and E/N fixed. From this we obtain the differential relation

$$\begin{split} dS &= \frac{Nk_{\mathrm{B}}}{V}\,dV + \frac{d}{\alpha}\,\frac{Nk_{\mathrm{B}}}{E}\,dE + s_0\,dN \\ &= \frac{p}{T}\,dV + \frac{1}{T}\,dE - \frac{\mu}{T}\,dN\;, \end{split}$$

where s_0 is a constant. From the coefficients of dV and dE, we conclude

$$\begin{split} pV &= N k_{\mathrm{B}} T \\ E &= \frac{d}{\alpha} N k_{\mathrm{B}} T \; . \end{split}$$

Note that we have replaced $E = \frac{d}{\alpha} N k_{\rm B} T$ in order to express F in terms of its 'natural variables' T, V, and N.

The Helmholtz free energy is

$$\begin{split} F &= E - TS = E - Nk_{\mathrm{B}}T\ln\left(\frac{V}{N}\right) - \frac{d}{\alpha}Nk_{\mathrm{B}}T\ln\left(\frac{E}{N}\right) - Nk_{\mathrm{B}}Ta_{0} \\ &= \frac{d}{\alpha}Nk_{\mathrm{B}}T - \frac{d}{\alpha}Nk_{\mathrm{B}}T\ln\left(\frac{d}{\alpha}k_{\mathrm{B}}T\right) - Nk_{\mathrm{B}}T\ln\left(\frac{V}{N}\right) - Nk_{\mathrm{B}}Ta_{0} \;. \end{split}$$

The chemical potential is

$$\begin{split} \mu &= T \left(\frac{\partial F}{\partial N} \right)_{\!\! T,V} = -\frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T \ln \! \left(\frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T \right) + \frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T - k_{\rm\scriptscriptstyle B} T \ln \! \left(\frac{V}{N} \right) + \left(1 - a_0 \right) k_{\rm\scriptscriptstyle B} T \\ &= -\frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T \ln \! \left(\frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T \right) + \frac{d}{\alpha} \, k_{\rm\scriptscriptstyle B} T - k_{\rm\scriptscriptstyle B} T \ln \! \left(\frac{k_{\rm\scriptscriptstyle B} T}{p} \right) + \left(1 - a_0 \right) k_{\rm\scriptscriptstyle B} T \; . \end{split}$$

Suppose we wanted the heat capacities C_V and C_p . Setting dN = 0, we have

$$\begin{split} dQ &= dE + p \, dV \\ &= \frac{d}{\alpha} \, N k_{\mathrm{B}} \, dT + p \, dV \\ &= \frac{d}{\alpha} \, N k_{\mathrm{B}} \, dT + p \, d \bigg(\frac{N k_{\mathrm{B}} T}{p} \bigg) \; . \end{split}$$

Thus,

$$C_V = \frac{dQ}{dT}\Big|_V = \frac{d}{\alpha}Nk_{\mathrm{B}} \quad , \quad C_p = \frac{dQ}{dT}\Big|_p = \left(1 + \frac{d}{\alpha}\right)Nk_{\mathrm{B}} \; .$$

(2) Consider a gas of classical spin- $\frac{3}{2}$ particles, with Hamiltonian

$$\hat{H} = \sum_{i=1}^{N} \frac{p_i^2}{2m} - \mu_0 H \sum_i S_i^z ,$$

where $S_i^z \in \left\{-\frac{3}{2}, -\frac{1}{2}, +\frac{1}{2}, +\frac{3}{2}\right\}$ and H is the external magnetic field. Find the Helmholtz free energy F(T,V,H,N), the entropy S(T,V,H,N), and the magnetic susceptibility $\chi(T,H,n)$, where n=N/V is the number density.

Solution:

The partition function is

$$Z = {\rm Tr} \; e^{-\hat{H}/k_{\rm B}T} = \frac{1}{N!} \frac{V^N}{\lambda_T^{dN}} \Big(2 \cosh(\mu_0 H/2k_{\rm B}T) + 2 \cosh(3\mu_0 H/2k_{\rm B}T) \Big)^N \; , \label{eq:Z}$$

SO

$$F = -Nk_{\mathrm{B}}T\ln\biggl(\frac{V}{N\lambda_{T}^{d}}\biggr) - Nk_{\mathrm{B}}T - Nk_{\mathrm{B}}T\ln\biggl(2\cosh(\mu_{0}H/2k_{\mathrm{B}}T) + 2\cosh(3\mu_{0}H/2k_{\mathrm{B}}T)\biggr)\;,$$

where $\lambda_T = \sqrt{2\pi\hbar^2/mk_{\mathrm{B}}T}$ is the thermal wavelength. The entropy is

$$\begin{split} S = -\left(\frac{\partial F}{\partial T}\right)_{V,N,H} &= Nk_{\mathrm{B}} \ln \left(\frac{V}{N\lambda_{T}^{d}}\right) + (\frac{1}{2}d+1)Nk_{\mathrm{B}} + N \ln \left(2\cosh(\mu_{0}H/2k_{\mathrm{B}}T) + 2\cosh(3\mu_{0}H/2k_{\mathrm{B}}T)\right) \\ &- \frac{\mu_{0}H}{2T} \cdot \frac{\sinh(\mu_{0}H/2k_{\mathrm{B}}T) + 3\sinh(3\mu_{0}H/2k_{\mathrm{B}}T)}{\cosh(\mu_{0}H/2k_{\mathrm{B}}T) + \cosh(3\mu_{0}H/2k_{\mathrm{B}}T)} \;. \end{split}$$

The magnetization is

$$M = -\left(\frac{\partial F}{\partial H}\right)_{T,V,N} = \tfrac{1}{2}N\mu_0 \cdot \frac{\sinh(\mu_0 H/2k_{\mathrm{B}}T) + 3\sinh(3\mu_0 H/2k_{\mathrm{B}}T)}{\cosh(\mu_0 H/2k_{\mathrm{B}}T) + \cosh(3\mu_0 H/2k_{\mathrm{B}}T)} \;.$$

The magnetic susceptibility is

$$\chi(T,H,n) = \frac{1}{V} \left(\frac{\partial M}{\partial H} \right)_{T,V,N} = \frac{n\mu_0^2}{4k_{\rm B}T} f(\mu_0 H/2k_{\rm B}T)$$

where

$$f(x) = \frac{d}{dx} \left(\frac{\sinh x + 3\sinh(3x)}{\cosh x + \cosh(3x)} \right).$$

In the limit $H\to 0$, we have f(0)=5, so $\chi=4n\mu_0^2/4k_{\rm B}T$ at high temperatures. This is a version of Curie's law.

(3) Compute the RMS volume fluctuations in the T-p-N ensemble.

Solution:

Averages within the T - p - N ensemble are computed by

$$\langle A \rangle = \frac{\operatorname{Tr} A e^{-\beta(\hat{H}+pV)}}{\operatorname{Tr} e^{-\beta(\hat{H}+pV)}} .$$

Let $Y = \operatorname{Tr}^{-\beta(\hat{H}+pV)} = e^{-\beta G}$. Then

$$\begin{split} \langle V^2 \rangle &= \frac{1}{\beta^2 Y} \frac{\partial^2 Y}{\partial p^2} = \beta^{-2} \, e^{\beta G} \, \frac{\partial^2}{\partial p^2} \, e^{-\beta G} \\ &= -\frac{1}{\beta} \, \frac{\partial^2 G}{\partial p^2} + \left(\frac{\partial G}{\partial p}\right)^2 \,, \end{split}$$

and since $\frac{\partial G}{\partial p} = V$, we have

$$\langle V^2 \rangle - \langle V \rangle^2 = -k_{\rm B} T \, \frac{\partial^2 G}{\partial n^2} \, .$$

For the case of a nonrelativistic ideal gas, we have

$$\begin{split} \langle V^k \rangle &= \int\limits_0^\infty \!\! dV \, e^{-\beta p V} \, Z(T,V,N) \, V^k \, \Bigg/ \int\limits_0^\infty \!\! dV \, e^{-\beta p V} \, Z(T,V,N) \\ &= \int\limits_0^\infty \!\! dV \, e^{-\beta p V} \, V^{N+k} \, \Bigg/ \int\limits_0^\infty \!\! dV \, e^{-\beta p V} \, V^N = \frac{(N+k)!}{N!} \bigg(\frac{k_{\rm\scriptscriptstyle B} T}{p}\bigg)^k \,, \end{split}$$

since $Z(T, V, N) = \frac{1}{N!} (V/\lambda_T)^N$. Thus,

$$\langle V \rangle = (N+1) \frac{k_{\rm B}T}{p} \qquad , \qquad \langle V^2 \rangle = (N+1)(N+2) \left(\frac{k_{\rm B}T}{p}\right)^2$$

and therefore

$$V_{\rm rms}^2 = \langle V^2 \rangle - \langle V \rangle^2 = (N+1) \left(\frac{k_{\rm B}T}{p}\right)^2 \quad \Rightarrow \quad V_{\rm rms} = N^{1/2} \frac{k_{\rm B}T}{p} \,.$$

Thus $V_{\rm rms}/\langle V \rangle = N^{-1/2} \ll 1$. This is, once again, the Central Limit Theorem in action.

(4) For the system described in problem (1), compute the distribution of speeds $\bar{f}(v)$. Find the most probable speed, the mean speed, and the RMS speed.

Solution:

Again, we solve for the general case $\varepsilon(\mathbf{p}) = Ap^{\alpha}$. The momentum distribution is

$$g(\mathbf{p}) = C e^{-\beta A p^{\alpha}} ,$$

where C is a normalization constant, defined so that $\int d^d p \ g(\mathbf{p}) = 1$. Changing variables to $t \equiv \beta A p^{\alpha}$, we find

$$C = \frac{\alpha \left(\beta A\right)^{\frac{d}{\alpha}}}{\Omega_d \Gamma\left(\frac{d}{\alpha}\right)}.$$

The velocity v is given by

$$\boldsymbol{v} = \frac{\partial \varepsilon}{\partial \boldsymbol{p}} = \alpha A p^{\alpha - 1} \, \hat{\boldsymbol{p}} \,.$$

Thus, the speed distribution is given by

$$\bar{f}(v) = C \int d^d p \ e^{-\beta A p^{\alpha}} \, \delta(v - \alpha A p^{\alpha - 1}) \ .$$

Now

$$\delta(v - \alpha A p^{\alpha - 1}) = \frac{\delta(p - (v/\alpha A)^{1/(\alpha - 1)})}{\alpha(\alpha - 1)Ap^{\alpha - 2}}.$$

We therefore have

$$\bar{f}(v) = \frac{C}{\alpha(\alpha - 1)A} p^{d-\alpha+1} e^{-\beta Ap^{\alpha}} \bigg|_{p=(v/\alpha A)^{1/(\alpha-1)}}.$$

We can now calculate

$$\langle v^r \rangle = C \int \!\! d^d\! p \, e^{-\beta A p^\alpha} \big(\alpha A p^{\alpha-1} \big)^r \; , \label{eq:constraint}$$

and so

$$||v||_r = \langle v^r \rangle^{1/r} = \alpha A^{\alpha^{-1}} \left(k_{\mathrm{B}} T \right)^{1-\alpha^{-1}} \left(\frac{\Gamma\left(\frac{d-r}{\alpha} + r\right)}{\Gamma\left(\frac{d}{\alpha}\right)} \right)^{1/\alpha}.$$

To find the most probable speed, we extremize $\bar{f}(v)$. We obtain

$$\beta A p^{\alpha} = \frac{d - \alpha + 1}{\alpha}$$
,

which means

$$v = \alpha A \left(\frac{d - \alpha + 1}{\alpha \beta A} \right)^{1 - \alpha^{-1}} = (\alpha A)^{\alpha^{-1}} (d - \alpha + 1)^{1 - \alpha^{-1}} (k_{\rm B} T)^{1 - \alpha^{-1}} .$$