

**PHYSICS 140A : STATISTICAL PHYSICS
FINAL EXAMINATION SOLUTIONS**

Instructions: Do problem 4 (34 points) and any two of problems 1, 2, and 3 (33 points each)

(1) A noninteracting system consists of N dimers. Each dimer consists of two spins, S and σ , where $S \in \{-1, 0, +1\}$ and $\sigma \in \{-1, +1\}$. The Hamiltonian is

$$\hat{H} = -J \sum_{i=1}^N S_i \sigma_i - \mu_0 H \sum_{i=1}^N S_i.$$

Thus, the individual dimer Hamiltonian is $\hat{h} = -JS\sigma - \mu_0 HS$.

(a) Find the N -dimer free energy $F(T, N)$.

(b) Find the average $\langle S \rangle$ and the zero field susceptibility $\chi_S(T) = \left. \frac{\partial \langle S \rangle}{\partial H} \right|_{H=0}$.

(c) Find the average $\langle \sigma \rangle$ and the zero field susceptibility $\chi_\sigma(T) = \left. \frac{\partial \langle \sigma \rangle}{\partial H} \right|_{H=0}$.

(d) Examine the $J \rightarrow 0$ limits of $\chi_S(T)$ and $\chi_\sigma(T)$ and interpret your results physically.

Solution :

(a) There are six energy states for each dimer, listed in Tab. 1

S	σ	$\hat{h}(S, \sigma)$	S	σ	$\hat{h}(S, \sigma)$
+1	+1	$-J - \mu_0 H$	+1	-1	$J - \mu_0 H$
0	+1	0	0	-1	0
-1	+1	$+J + \mu_0 H$	-1	-1	$-J + \mu_0 H$

Table 1: Energy table for problem 1.

Thus, the single dimer partition function is

$$\begin{aligned} \zeta &= \text{Tr} e^{-\beta \hat{h}} = e^{\beta J} e^{\beta \mu_0 H} + 1 + e^{-\beta J} e^{-\beta \mu_0 H} + e^{-\beta J} e^{\beta \mu_0 H} + 1 + e^{\beta J} e^{-\beta \mu_0 H} \\ &= 2 + 4 \cosh(\beta \mu_0 J) \cosh(\beta \mu_0 H). \end{aligned}$$

For N noninteracting dimers, $Z = \zeta^N$ (the dimers are regarded as distinguishable). Thus,

$$F(T, N) = -N k_B T \ln \left(2 + 4 \cosh(J/k_B T) \cosh(\mu_0 H/k_B T) \right)$$

(b) We have

$$\langle S \rangle = \frac{\text{Tr} S e^{-\beta \hat{h}(S, \sigma)}}{\text{Tr} e^{-\beta \hat{h}(S, \sigma)}} = \frac{e^{\beta J} e^{\beta \mu_0 H} - e^{-\beta J} e^{-\beta \mu_0 H} + e^{-\beta J} e^{\beta \mu_0 H} - e^{\beta J} e^{-\beta \mu_0 H}}{2 + 4 \cosh(\beta J) \cosh(\beta H)},$$

so

$$\langle S \rangle = \frac{\cosh(J/k_B T) \sinh(\mu_0 H/k_B T)}{\cosh(J/k_B T) \cosh(\mu_0 H/k_B T) + \frac{1}{2}}$$

Expanding to linear order in H and taking the coefficient, we have

$$\chi_S(T) = \left. \frac{\partial \langle S \rangle}{\partial H} \right|_{H=0} = \frac{\cosh(J/k_B T)}{\cosh(J/k_B T) + \frac{1}{2}} \cdot \frac{\mu_0}{k_B T}$$

Note that usually we define $\chi = \frac{\partial M}{\partial T}$ with $M = \mu_0 \langle S \rangle$, so our result above differs by a factor of μ_0 .

(c) We have

$$\langle \sigma \rangle = \frac{\text{Tr } \sigma e^{-\beta \hat{h}(S, \sigma)}}{\text{Tr } e^{-\beta \hat{h}(S, \sigma)}} = \frac{e^{\beta J} e^{\beta \mu_0 H} + e^{-\beta J} e^{-\beta \mu_0 H} - e^{-\beta J} e^{\beta \mu_0 H} - e^{\beta J} e^{-\beta \mu_0 H}}{2 + 4 \cosh(\beta J) \cosh(\beta H)},$$

so

$$\langle \sigma \rangle = \frac{\sinh(J/k_B T) \sinh(\mu_0 H/k_B T)}{\cosh(J/k_B T) \cosh(\mu_0 H/k_B T) + \frac{1}{2}}$$

Expanding to linear order in H and taking the coefficient, we have

$$\chi_\sigma(T) = \left. \frac{\partial \langle \sigma \rangle}{\partial H} \right|_{H=0} = \frac{\sinh(J/k_B T)}{\cosh(J/k_B T) + \frac{1}{2}} \cdot \frac{\mu_0}{k_B T}$$

(d) As $J \rightarrow 0$ we have

$$\chi_S(T, J=0) = \frac{2\mu_0}{3k_B T}, \quad \chi_\sigma(T, J=0) = 0$$

The physical interpretation of these results is as follows. When $J = 0$, the individual dimer Hamiltonian is $\hat{h} = -\mu_0 H S$. The factor of $\frac{2}{3}$ in χ_S is due to the fact that $S = 0$ in $\frac{1}{3}$ of the states. The σ spins don't couple to the field at all in this limit, so $\chi_\sigma = 0$.

(2) Recall that a van der Waals gas obeys the equation of state

$$\left(p + \frac{a}{v^2}\right)(v - b) = RT,$$

where v is the molar volume. We showed that the energy per mole of such a gas is given by

$$\varepsilon(T, v) = \frac{1}{2} f RT - \frac{a}{v},$$

where T is temperature and f is the number of degrees of freedom per particle.

(a) For an ideal gas, the adiabatic equation of state is $vT^{f/2} = \text{const}$. Find the adiabatic equation of state (at fixed particle number) for the van der Waals gas.

(b) One mole of a van der Waals gas is used as the working substance in a Carnot engine (see Fig. 1). Find the molar volume at v_C in terms of v_B, T_1, T_2 , and constants.

(c) Find the heat Q_{AB} absorbed by the gas from the upper reservoir.

(d) Find the work done per cycle, W_{cyc} . *Hint: you only need to know Q_{AB} and the cycle efficiency η .*

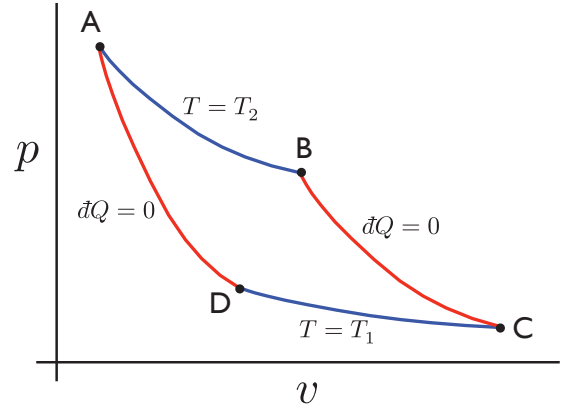


Figure 1: The Carnot cycle.

Solution :

(a) We have

$$\begin{aligned} 0 &= T ds = d\varepsilon + p dv \\ &= \frac{1}{2}fR dT + \left(p + \frac{a}{v^2}\right) dv \\ &= \frac{1}{2}fR dT + \frac{RT dv}{v-b} = \frac{1}{2}fRT d\ln[(v-b)T^{f/2}] , \end{aligned}$$

where $s = N_A S/N$ is the molar entropy. Thus, the adiabatic equation of state for the van der Waals gas is

$$ds = 0 \Rightarrow (v-b)T^{f/2} = \text{const.}$$

Setting $b = 0$, we recover the ideal gas result.

(b) Since BC is an adiabat, we have

$$(v_B - b)T_2^{f/2} = (v_C - b)T_1^{f/2} \Rightarrow v_C = b + (v_B - b) \left(\frac{T_2}{T_1}\right)^{f/2}$$

(c) We have, from the First Law,

$$\begin{aligned} Q_{AB} &= E_B - E_A + W_{AB} \\ &= \nu \left(\frac{a}{v_A} - \frac{a}{v_B}\right) + \nu \int_{v_A}^{v_B} dv p \\ &= \nu \left(\frac{a}{v_A} - \frac{a}{v_B}\right) + \nu \int_{v_A}^{v_B} dv \left[\frac{RT_2}{v-b} - \frac{a}{v^2}\right] , \end{aligned}$$

hence

$$Q_{AB} = \nu R T_2 \ln \left(\frac{v_B - b}{v_A - b} \right)$$

with $\nu = 1$.

(d) Since the cycle is reversible, we must have

$$\eta = \frac{W_{\text{cyc}}}{Q_{AB}} \Rightarrow W_{\text{cyc}} = \nu R (T_2 - T_1) \ln \left(\frac{v_B - b}{v_A - b} \right)$$

(3) In homework assignment #9, you showed that the grand partition function for a gas of q -state parafermions is

$$\Xi(T, V, \mu) = \prod_{\alpha} \left(\frac{1 - e^{(q+1)(\mu - \varepsilon_{\alpha})/k_B T}}{1 - e^{(\mu - \varepsilon_{\alpha})/k_B T}} \right),$$

where the product is over all single particle states. Consider now the case where the number of parafermions is not conserved, hence $\mu = 0$. We call such particles *paraphotons*.

(a) What is the occupancy $n(\varepsilon, T)$ of q -state paraphotons of energy ε ?

(b) Suppose the dispersion is the usual $\varepsilon(\mathbf{k}) = \hbar c k$. Assuming $g = 1$, find the single particle density of states $g(\varepsilon)$ in three space dimensions.

(c) Find the pressure $p(T)$. You may find the following useful:

$$\int_0^{\infty} dt \frac{t^{r-1}}{e^t - 1} = \Gamma(r) \zeta(r) \quad , \quad \int_0^{\infty} dt t^{r-1} \ln \left(\frac{1}{1 - e^{-t}} \right) = \Gamma(r) \zeta(r + 1).$$

(d) Show that $p = C_q n k_B T$, where n is the number density, and C_q is a dimensionless constant which depends only on q .

Solution :

(a) For $\mu \neq 0$, for a single parafermion state, we have

$$\begin{aligned} n &= -\frac{\partial \Omega}{\partial \mu} = \frac{1}{\beta} \frac{\partial \ln \Xi}{\partial \mu} \\ &= \frac{1}{e^{\beta(\varepsilon - \mu)} - 1} - \frac{q + 1}{e^{(q+1)\beta(\varepsilon - \mu)} - 1}. \end{aligned}$$

Setting $\mu = 0$, we find

$$n(\varepsilon, T) = \frac{1}{e^{\varepsilon/k_B T} - 1} - \frac{q + 1}{e^{(q+1)\varepsilon/k_B T} - 1}$$

(b) With $g = 1$, we have

$$g(\varepsilon) d\varepsilon = \frac{d^3k}{(2\pi)^3} = \frac{k^2 dk}{2\pi^2} \Rightarrow \boxed{g(\varepsilon) = \frac{k^2}{2\pi^2} \frac{dk}{d\varepsilon} = \frac{\varepsilon^2}{2\pi^2(\hbar c)^3}}$$

(c) The pressure is

$$\begin{aligned} p &= -\frac{\Omega}{V} = k_B T \int_0^\infty d\varepsilon g(\varepsilon) \left\{ \ln(1 - e^{-(q+1)\varepsilon/k_B T}) - \ln(1 - e^{-\varepsilon/k_B T}) \right\} \\ &= \frac{k_B T}{2\pi^2(\hbar c)^3} \int_0^\infty d\varepsilon \varepsilon^2 \left\{ \ln(1 - e^{-(q+1)\varepsilon/k_B T}) - \ln(1 - e^{-\varepsilon/k_B T}) \right\} \\ &= \frac{\zeta(4) (k_B T)^4}{\pi^2(\hbar c)^3} \cdot \left(1 - \frac{1}{(q+1)^3} \right). \end{aligned}$$

Thus,

$$\boxed{p(T) = \left(1 - (q+1)^{-3} \right) \cdot \frac{\zeta(4) (k_B T)^4}{\pi^2(\hbar c)^3}}$$

(d) We need to evaluate

$$\begin{aligned} n &= \int_0^\infty d\varepsilon g(\varepsilon) \left\{ \frac{1}{e^{\varepsilon/k_B T} - 1} - \frac{q+1}{e^{(q+1)\varepsilon/k_B T} - 1} \right\} \\ &= \frac{1}{2\pi^2(\hbar c)^3} \int_0^\infty d\varepsilon \varepsilon^2 \left\{ \frac{1}{e^{\varepsilon/k_B T} - 1} - \frac{q+1}{e^{(q+1)\varepsilon/k_B T} - 1} \right\} \\ &= \left(1 - (q+1)^{-2} \right) \cdot \frac{\zeta(3) (k_B T)^3}{\pi^2(\hbar c)^3} \end{aligned}$$

From this we derive

$$\boxed{C_q = \frac{p}{nk_B T} = \frac{\zeta(4)}{\zeta(3)} \cdot \frac{q^2 + 3q + 3}{q^2 + 3q + 2}}$$

(4) Provide brief but substantial answers to the following:

(a) A particle in $d = 3$ dimensions has the dispersion $\varepsilon(\mathbf{k}) = \varepsilon_0 \exp(ka)$. Find the density of states per unit volume $g(\varepsilon)$. Sketch your result.

(b) Find the information entropy in the distribution $p_n = C e^{-\lambda n}$, where $n \in \{0, 1, 2, \dots\}$. Choose C so as to normalize the distribution.

(c) An ideal gas at temperature $T = 300$ K undergoes an adiabatic free expansion which results in a doubling of its volume. What is the final temperature?

(d) For an N -particle noninteracting system, sketch the contributions ΔC_V to the heat capacity *versus* temperature for (i) a vibrational mode at energy $\hbar\omega_0$, and (ii) a two-level (Schottky) defect with energy splitting $\Delta = \varepsilon_1 - \varepsilon_0$. Take care to identify any relevant characteristic temperatures, as well as the limiting values of ΔC_V .

Solution :

(a) Inverting the dispersion relation, we obtain the expression $k(\varepsilon) = a^{-1} \ln(\varepsilon/\varepsilon_0) \Theta(\varepsilon - \varepsilon_0)$. We then have

$$g(\varepsilon) = \frac{k^2}{2\pi} \frac{dk}{d\varepsilon} = \frac{k^2}{2\pi} \cdot \frac{1}{a\varepsilon_0 e^{ak}}.$$

Thus,

$$g(\varepsilon) = \frac{1}{2\pi^2 a^3} \frac{1}{\varepsilon} \ln^2\left(\frac{\varepsilon}{\varepsilon_0}\right) \Theta(\varepsilon - \varepsilon_0)$$

The result is plotted in Fig. 2.

(b) Normalizing the distribution,

$$1 = \sum_{n=0}^{\infty} p_n = C \sum_{n=0}^{\infty} e^{-n\lambda} = \frac{C}{1 - e^{-\lambda}},$$

hence $C = 1 - e^{-\lambda}$. The information entropy is

$$S = - \sum_{n=0}^{\infty} p_n \ln p_n = - \ln(1 - e^{-\lambda}) + C\lambda \sum_{n=0}^{\infty} n e^{-\lambda n}.$$

Now

$$f(\lambda) = \sum_{n=0}^{\infty} e^{-n\lambda} = \frac{1}{1 - e^{-\lambda}} \Rightarrow \sum_{n=0}^{\infty} n e^{-n\lambda} = -\frac{df}{d\lambda} = \frac{1}{(e^\lambda - 1)(1 - e^{-\lambda})}.$$

Thus, the information entropy is

$$S(\lambda) = \frac{\lambda}{e^\lambda - 1} - \ln(1 - e^{-\lambda}).$$

Note that $S(\lambda \rightarrow 0) \sim 1 - \ln \lambda$ which diverges logarithmically with $1/\lambda$. This is approaching the uniform distribution. For $\lambda \rightarrow \infty$, we have $p_n = \delta_{n,0}$, and $S(\lambda \rightarrow \infty) = 0$.

(c) Under an adiabatic free expansion, $\Delta E = Q = W = 0$ with N conserved. Since $E = \frac{1}{2} f N k_B T$ is independent of volume for the ideal gas, there is no change in temperature, *i.e.*

$$T_{\text{final}} = T_{\text{initial}} = 100 \text{ K}$$

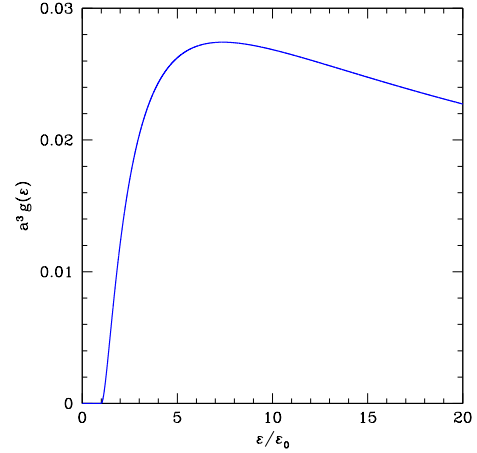


Figure 2: DOS for problem 4.a.

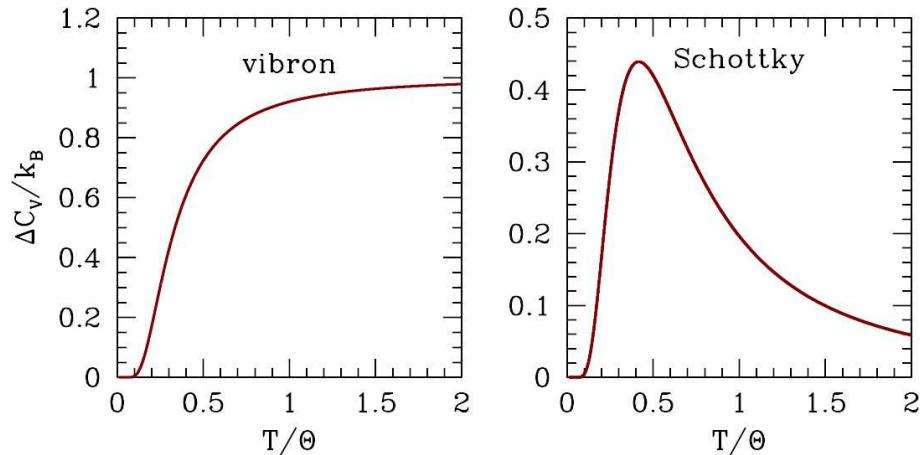


Figure 3: Heat capacities for a N identical vibrational modes (left) and Schottky defects (right).

(d) The characteristic temperatures for the vibrational mode (vibron) and Schottky defect are given by $\theta = \hbar\omega_0/k_B$ and $\theta = \Delta/k_B$, respectively. A detailed derivation of the heat capacity for these systems is provided in §§ 4.10.5-6 of the Lecture Notes. One finds

$$\Delta C_V = Nk_B \left(\frac{\theta}{T} \right)^2 \frac{e^{\theta/T}}{(e^{\theta/T} \mp 1)^2},$$

where the top sign is for the vibron and the bottom sign for the Schottky defect. All you were asked to do, however, was to provide a sketch (see Fig. 3). The $T \rightarrow \infty$ limit of the vibron result is given by the Dulong-Petit value of k_B per oscillator mode. For the Schottky defect, ΔC_V vanishes in both the $T \rightarrow 0$ and $T \rightarrow \infty$ limits.

(5) Write a well-defined expression for the greatest possible number expressible using only five symbols. Examples: $1 + 2 + 3$, 10^{100} , $\Gamma(99)$. [50 quatlous extra credit]

Solution :

Using conventional notation, my best shot would be

$$9^{9^{9^9}}$$

This is a very big number indeed: $9^9 \approx 3.7 \times 10^8$, so $9^{9^9} \sim 10^{3.5 \times 10^8}$, and $9^{9^{9^9}} \sim 10^{10^{3.7 \times 10^8}}$. But in the world of big numbers, this is still tiny. For a fun diversion, use teh google to learn about the Ackermann sequence and Knuth's up-arrow notation. Using Knuth's notation (see http://en.wikipedia.org/wiki/Knuth's_up-arrow_notation), one could write $9 \uparrow^{99} 9$, which is vastly larger. But even *these* numbers are modest compared with something called the "Busy Beaver sequence", which is a concept from computer science and Turing machines. For a very engaging essay on large numbers, see <http://www.scottaaronson.com/writings/bignumbers.html>.