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Lect 5. Partial wave method

In the center-force field, \vec{L} is conserved. We can separate the scattering amplitude in each particle wave channel. We choose (H, \vec{L}, l_z) as the complete set of conserved quantities. We want to solve

$$\left(-\frac{\hbar^2}{2m} \nabla^2 + V(r) \right) \psi = E \psi, \text{ under the scattering boundary condition of } \psi(r) \xrightarrow{r \rightarrow +\infty} e^{ikz} + f(\theta) \frac{e^{ikr}}{r}.$$

Partial wave means, we decompose this boundary condition into different channel of l . The incoming wave can be decomposed as

$$\begin{aligned} e^{ikz} &= e^{ikr \cos \theta} = \sum_{l=0}^{\infty} (2l+1) i^l j_l(kr) P_l(\cos \theta) \\ &= \sum_{l=0}^{\infty} \sqrt{4\pi(2l+1)} i^l j_l(kr) Y_{l0}(\theta) \\ &\xrightarrow{k r \rightarrow +\infty} \sum_{l=0}^{\infty} \sqrt{4\pi(2l+1)} i^l \frac{1}{2ikr} [e^{i(kr - \frac{l\pi}{2})} - e^{-i(kr - \frac{l\pi}{2})}] Y_{l0} \end{aligned}$$

$j_l(kr)$ is l -th spheric Bessel function, which is the solution of the Laplace equation in the spherical coordinate system:

radial part of

$$\frac{d^2 R(p)}{dp^2} + \frac{2}{p} \frac{dR(p)}{dp} + \left[1 - \frac{l(l+1)}{p^2} \right] R(p) = 0$$

where $p = kr$.

The scattering wave can be decomposed $f(\theta) = \sum_l f_l$, thus

in the l -th partial wave channel, the boundary condition \rightarrow

$$\left[\sqrt{4\pi(2l+1)} i^l j_l(kr) + \frac{f_l e^{ikr}}{r} \right] Y_{l0}(\theta)$$

On the other hand, we can solve the radial equation

$$\left[-\frac{\hbar}{2m} \nabla^2 + V(r) \right] \psi(r) = E \psi \quad \leftarrow \psi = \sum_{l=0}^{\infty} R_l(kr) Y_{l0}(\theta)$$

$$\Rightarrow \left[\frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{d}{dr} \right) + k^2 - \frac{l(l+1)}{r^2} - U(r) \right] R_l = 0, \text{ where } E = \frac{\hbar^2 k^2}{2m}$$

$$U(r) = \frac{2mV(r)}{\hbar^2}$$

at $kr \rightarrow +\infty$, $U(r) \rightarrow 0$. Thus in general R_l can be

Expanded as a linear superposition of incoming wave + scattering outgoing wave.

* In the spherical coordination, incoming wave is $j_l(kr)$,
out going wave is $h_l(kr)$.

Remember that there are two linearly independent solutions to

$$\left[\frac{1}{r^2} \frac{d}{dr} \left(r^2 \frac{d}{dr} \right) + k^2 - \frac{l(l+1)}{r^2} \right] R_l = 0.$$

$$\left. \begin{aligned} j_l(p) &= (-)^l p^l \left(\frac{1}{p} \frac{d}{dp} \right)^l \frac{\sin p}{p} \\ n_l(p) &= (-)^{l+1} p^l \left(\frac{1}{p} \frac{d}{dp} \right)^l \frac{\cos p}{p} \end{aligned} \right\} \quad p = kr$$

Examples: $j_0(kr) = \frac{\sin kr}{kr}$ $n_0(kr) = -\frac{\cos kr}{kr}$

$$j_1(kr) = \frac{\sin kr}{(kr)^2} - \frac{\cos kr}{kr} \quad n_1(kr) = -\frac{\cos kr}{(kr)^2} - \frac{\sin kr}{kr}$$

$$\rightarrow \frac{(kr)^l}{(2\ell+1)!!} \quad \rightarrow -(2\ell-1)!!/(kr)^{l+1} \quad (3)$$

$j_\ell(kr)$ is regular at $kr \rightarrow 0$, $n_\ell(kr)$ diverges at $kr \rightarrow 0$.

On the other hand as $kr \rightarrow +\infty$, they behave as

$$j_\ell(kr) \xrightarrow{r \rightarrow +\infty} \frac{1}{kr} \sin(kr - \frac{\ell\pi}{2}), \quad n_\ell(kr) \xrightarrow{r \rightarrow +\infty} \frac{-1}{kr} \cos(kr - \frac{\ell\pi}{2}).$$

From j_ℓ and n_ℓ , we can combine into propagating waves as

$$h_\ell(kr) = j_\ell(kr) + i n_\ell(kr) \xrightarrow{kr \rightarrow +\infty} \frac{1}{ikr} e^{i(kr - \ell\pi/2)}$$

* * * * *

So let us express the solution from solving the radial equation as

$$R_\ell(kr) \xrightarrow{kr \rightarrow +\infty} \sqrt{4\pi(2\ell+1)} i^\ell \left[j_\ell(kr) + \frac{a_\ell}{2} h_\ell(kr) \right] \quad \text{← coefficient}$$

$$= \sqrt{4\pi(2\ell+1)} i^\ell \frac{1}{2ikr} \left[\underbrace{(1+a_\ell) e^{i(kr - \ell\pi/2)}}_{\substack{\uparrow \\ \downarrow \\ \leftarrow \\ \rightarrow}} - \underbrace{e^{-[i(kr - \ell\pi/2)]}}_{\substack{\downarrow \\ \rightarrow \\ \leftarrow \\ \uparrow}} \right]$$

$$\Rightarrow |1+a_\ell| = 1 \quad \Rightarrow \text{let us parameterize } 1+a_\ell = e^{2i\delta_\ell} \\ \Rightarrow a_\ell = e^{i\delta_\ell} (e^{i\delta_\ell} - e^{-i\delta_\ell})$$

$$\Rightarrow \boxed{R_\ell(kr) \xrightarrow{kr \rightarrow +\infty} \sqrt{4\pi(2\ell+1)} i^\ell e^{i\delta_\ell} \frac{1}{kr} \sin(kr - \frac{\ell\pi}{2} + \delta_\ell)} = e^{i\delta_\ell} 2i \sin \delta_\ell$$

What we get from solving the radial equation. The net effect of scattering potential is the phase shift δ_ℓ . For free space, $\delta_\ell = 0$.

Compare $\rightarrow R_{\ell}(kr) \xrightarrow{kr \rightarrow +\infty} \sqrt{4\pi(2\ell+1)} \frac{i^{\ell} e^{i\delta_{\ell}}}{kr} \sin(kr - \frac{\ell\pi}{2} + \delta_{\ell})$ *

from solving radial Eq

and

$$\rightarrow \sqrt{4\pi(2\ell+1)} i^{\ell} j_{\ell}(kr) + \frac{f_{\ell}}{r} e^{ikr}$$

from : general argument

* → goes back to $R_{\ell}(kr) \xrightarrow{kr \rightarrow +\infty} \sqrt{4\pi(2\ell+1)} i^{\ell} [j_{\ell}(kr) + i e^{i\delta_{\ell}} h_{\ell}(kr)]$

$$h_{\ell}(kr) \xrightarrow{kr \rightarrow +\infty} \frac{1}{ikr} e^{ikr - \ell\pi/2}$$

$$\Rightarrow \sqrt{4\pi(2\ell+1)} i^{\ell} \frac{1}{ikr} e^{ikr - \ell\pi/2} = \frac{f_{\ell}}{r} e^{ikr}$$

$$\Rightarrow f_{\ell} = \frac{1}{k} e^{i\delta_{\ell}} \sin \delta_{\ell} \sqrt{4\pi(2\ell+1)}$$

$$\sigma(\theta) = |f(\theta)|^2 = \frac{4\pi}{k^2} \left| \sum_{\ell=0}^{\infty} \sqrt{2\ell+1} e^{i\delta_{\ell}} \sin \delta_{\ell} Y_{\ell 0}(\theta) \right|^2$$

$$O_t = \int \sigma(\theta) d\Omega = \frac{4\pi}{k^2} \sum_{\ell=0}^{\infty} (2\ell+1) \sin^2 \delta_{\ell}$$

* As long as we can get the phase shift δ_{ℓ} , we know the cross section. Scattering problem is reduced into solving radial equation with the proper boundary condition of $R_{\ell}(kr) \xrightarrow{r \rightarrow +\infty} \frac{i^{\ell} e^{i\delta_{\ell}}}{kr} \sin(kr - \frac{\ell\pi}{2} + \delta_{\ell})$

Discussion:

* Optical theorem:

$$f(\theta) = \sum_l f_l Y_{l0}(\theta) = \sum_l \frac{e^{i\delta_l}}{k} \sin \delta_l \sqrt{4\pi(2l+1)} Y_{l0}(\theta)$$

$$\text{Im } f(\theta=0) = \frac{1}{k} \sum_{l=0}^{\infty} \sin^2 \delta_l (2l+1) = \frac{k}{4\pi} \sigma_t \Rightarrow \boxed{\sigma_t = \frac{4\pi}{k} \text{Im } f(0)}$$

* The sign of the phase shift δ_l . Consider the radial Eq.

$$\frac{1}{r^2} \left(\frac{d}{dr} r^2 \frac{d}{dr} R_l \right) + \left[k^2 - \frac{l(l+1)}{r^2} - U(r) \right] R_l = 0$$

$$R_l \xrightarrow{kr \rightarrow +\infty} \frac{1}{kr} \sin(kr - \frac{l\pi}{2} + \delta_l) \quad \text{where } k = \sqrt{\frac{2mE}{\hbar^2}}$$

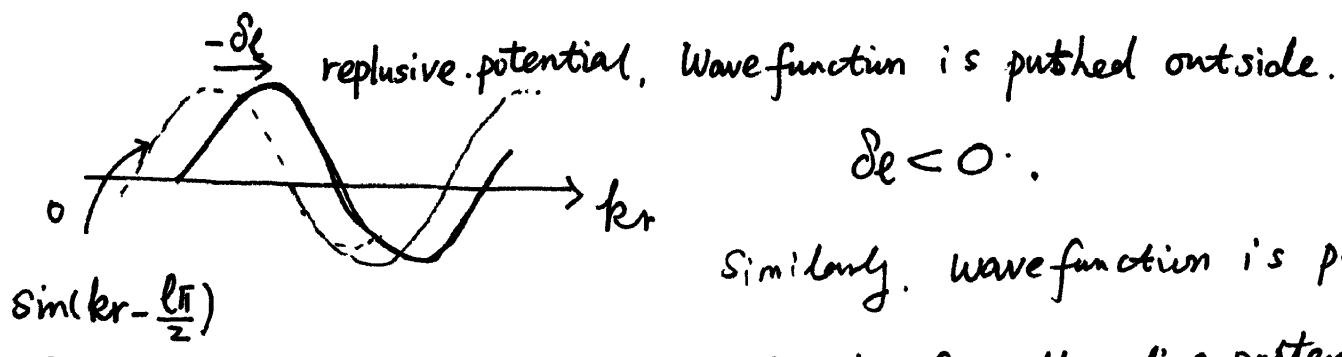
at $U(r) = 0$, we have $\delta_l = 0$.

there are

If $U(r) > 0$, it means that in the region of $U(r)$, less oscillating
more

occur $\Rightarrow \delta_l < 0$. (repulsive potential)

$\delta_l > 0$ (attractive potential).

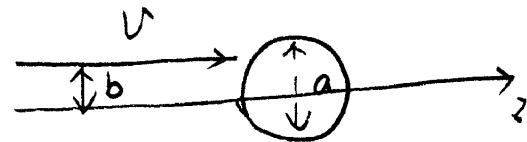


Similarly, wavefunction is pushed inside for attractive potential,
 $\delta_l > 0$.

* how many partial waves are needed?

Say, let us assume the range of force is a . Only where the distance $b \leq a$, the scattering effect, i.e. δ_e , is important.

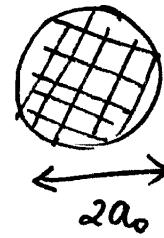
Thus $l = \ell\hbar \sim mvb \sim mva$



$l \sim \frac{mv}{\hbar} a = \frac{a}{\lambda}$, where λ is the de Broglie wave length of the incoming particle.

Example: hard sphere scattering.

$$V(r) = \begin{cases} \infty & r < a_0 \\ 0 & r > a_0 \end{cases}$$



Solving the Radial equation:

$$\left(\frac{1}{r^2} \left(\frac{d}{dr} r^2 \frac{d}{dr} R_e \right) + \left(k^2 - \frac{l(l+1)}{r^2} - \frac{2m}{\hbar^2} V(r) \right) R_e \right) = 0$$

$$R_e(kr) = \begin{cases} 0 & r < a \\ \cos \delta_e j_e(kr) - \sin \delta_e n_e(kr) & r > a \end{cases}$$

this form is the same as $j_e(kr) + i e^{i \delta_e} \sin \delta_e n_e(kr)$ up to

a phase factor. check! $\overleftarrow{j_e + i e^{i \delta_e} \sin \delta_e n_e} (j_e + i n_e)$

$$= (1 + i e^{i \delta_e} \sin \delta_e) j_e - e^{i \delta_e} \sin \delta_e n_e = e^{i \delta_e} [\cos \delta_e j_e - \sin \delta_e n_e]$$

Continuity at $r=a \Rightarrow \chi_0 = kr = ka$

$$R(ka) = 0 \Rightarrow \cos \delta_e(k) j_e(ka) = \sin \delta_e(k) n_e(ka)$$

$$\Rightarrow \tan \delta_e(k) = \frac{j_e(ka)}{n_e(ka)}$$

Low energy limit: $ka \rightarrow 0$:

$$j_e(ka) \xrightarrow{ka \rightarrow 0} \frac{(ka)^{\ell}}{(2\ell+1)!!} \quad n_e(ka) \xrightarrow{ka \rightarrow 0} -\frac{(2\ell-1)!!}{x^{\ell+1}}$$

$$\tan \delta_e(k) \xrightarrow{ka \rightarrow 0} -\frac{(ka)^{2\ell+1}}{[(2\ell-1)!!]^2 (2\ell+1)}$$

only the S-wave is important, we have

$$\boxed{\delta_0(k) \sim -\frac{(ka)}{x} < 0}$$

$\Omega_t \approx \frac{4\pi}{k^2} \sin^2 \delta_0 \approx 4\pi a^2$, which is 4 times larger

than the cross section.
classical