

5-12 Using $p = \frac{h}{\lambda} = mv$, we find that $v = \frac{h}{m\lambda} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{(9.11 \times 10^{-31} \text{ kg})(1 \times 10^{-10} \text{ m})} = 7.27 \times 10^6 \text{ m/s}$.

From the principle of conservation of energy, we get

$$eV = \frac{mv^2}{2} = \frac{(9.11 \times 10^{-31} \text{ kg})(7.27 \times 10^6 \text{ m/s})^2}{2} = 2.41 \times 10^{-17} \text{ J} = 151 \text{ eV}.$$

Therefore $V = 151 \text{ V}$.

5-15 For a free, non-relativistic electron $E = \frac{m_e v_0^2}{2} = \frac{p^2}{2m_e}$. As the wavenumber and angular frequency of the electron's de Broglie wave are given by $p = \hbar k$ and $E = \hbar \omega$, substituting these results gives the dispersion relation $\omega = \frac{\hbar k^2}{2m_e}$. So $v_g = \frac{d\omega}{dk} = \frac{\hbar k}{m_e} = \frac{p}{m_e} = v_0$.

5-17 $E^2 = p^2 c^2 + (m_e c^2)^2$
 $E = \left[p^2 c^2 + (m_e c^2)^2 \right]^{1/2}$. As $E = \hbar \omega$ and $p = \hbar k$

$$\hbar \omega = \left[\hbar^2 k^2 c^2 + (m_e c^2)^2 \right]^{1/2} \text{ or}$$

$$\omega(k) = \left[k^2 c^2 + \frac{(m_e c^2)^2}{\hbar^2} \right]^{1/2}$$

$$v_p = \frac{\omega}{k} = \frac{\left[k^2 c^2 + (m_e c^2 / \hbar)^2 \right]^{1/2}}{k} = \left[c^2 + \left(\frac{m_e c^2}{\hbar k} \right)^2 \right]^{1/2}$$

$$v_g = \left. \frac{d\omega}{dk} \right|_{k_0} = \frac{1}{2} \left[k^2 c^2 + \left(\frac{m_e c^2}{\hbar} \right)^2 \right]^{-1/2} 2kc^2 = \frac{kc^2}{\left[k^2 c^2 + (m_e c^2 / \hbar)^2 \right]^{1/2}}$$

$$v_p v_g = \left\{ \frac{\left[k^2 c^2 + (m_e c^2 / \hbar)^2 \right]^{1/2}}{k} \right\} \left\{ \left[k^2 c^2 + (m_e c^2 / \hbar)^2 \right]^{-1/2} \right\} = c^2$$

Therefore, $v_g < c$ if $v_p > c$.

5-18 $\Delta x \Delta p \geq \frac{\hbar}{2}$ where $\Delta p = m \Delta v = (0.05 \text{ kg})(10^{-3} \times 30 \text{ m/s}) = 1.5 \times 10^{-3} \text{ kg} \cdot \text{m/s}$. Therefore,

$$\Delta x = \frac{\hbar}{2\Delta p} = \frac{6.626 \times 10^{-34} \text{ J} \cdot \text{s}}{4\pi(1.5 \times 10^{-3} \text{ kg} \cdot \text{m/s})} = 3.51 \times 10^{-32} \text{ m}.$$

$$5-19 \quad K = \frac{mv^2}{2} = \frac{p^2}{2m}:$$

$$(1 \times 10^6 \text{ eV})(1.6 \times 10^{-19} \text{ J/eV}) = \frac{p^2}{2(1.67 \times 10^{-27} \text{ kg})} \Rightarrow p = 2.312 \times 10^{-20} \text{ kg} \cdot \text{m/s},$$

$$\Delta p = 0.05p = 1.160 \times 10^{-21} \text{ kg} \cdot \text{m/s} \text{ and } \Delta x \Delta p = \frac{h}{4\pi}. \text{ Thus}$$

$$\Delta x = \frac{6.63 \times 10^{-34} \text{ J} \cdot \text{s}}{(1.16 \times 10^{-21} \text{ kg} \cdot \text{m/s})(4\pi)} = 4.56 \times 10^{-14} \text{ m}.$$

Note that non-relativistic treatment has been used, which is justified because the kinetic energy is only $\frac{(1.6 \times 10^{-13}) \times 100\%}{1.50 \times 10^{-10}} = 0.11\%$

of the rest energy.

$$5-23 \quad (a) \quad \Delta p \Delta x = m \Delta v \Delta x \geq \frac{\hbar}{2}$$

$$\Delta v \geq \frac{h}{4\pi m \Delta x} = \frac{2\pi \text{ J} \cdot \text{s}}{4\pi(2 \text{ kg})(1 \text{ m})} = 0.25 \text{ m/s}$$

(b) The duck might move by $(0.25 \text{ m/s})(5 \text{ s}) = 1.25 \text{ m}$. With original position uncertainty of 1m, we can think of Δx growing to $1 \text{ m} + 1.25 \text{ m} = 2.25 \text{ m}$.

$$5-24 \quad (a) \quad \Delta x \Delta p = \hbar \text{ so if } \Delta x = r, \Delta p \approx \frac{\hbar}{r}$$

$$(b) \quad K = \frac{p^2}{2m_e} \approx \frac{(\Delta p)^2}{2m_e} = \frac{\hbar^2}{2m_e r^2}$$

$$U = -\frac{ke^2}{r}$$

$$E = \frac{\hbar^2}{2m_e r^2} - \frac{ke^2}{r}$$

(c) To minimize E take $\frac{dE}{dr} = -\frac{\hbar^2}{m_e r^3} + \frac{ke^2}{r^2} = 0 \Rightarrow r = \frac{\hbar^2}{m_e ke^2} = \text{Bohr radius} = a_0$. Then

$$E = \left(\frac{\hbar}{2m_e}\right) \left(\frac{m_e ke^2}{\hbar^2}\right)^2 - ke^2 \left(\frac{m_e ke^2}{\hbar^2}\right) = \frac{m_e k^2 e^4}{2\hbar^2} = -13.6 \text{ eV}.$$

5-25 To find the energy width of the γ -ray use $\Delta E \Delta t \geq \frac{\hbar}{2}$ or

$$\Delta E \geq \frac{\hbar}{2\Delta t} \geq \frac{6.58 \times 10^{-16} \text{ eV} \cdot \text{s}}{(2)(0.10 \times 10^{-9} \text{ s})} \geq 3.29 \times 10^{-6} \text{ eV}.$$

As the intrinsic energy width of $\sim \pm 3 \times 10^{-6}$ eV is so much less than the experimental resolution of ± 5 eV, the intrinsic width can't be measured using this method.

5-26 The full width at half-maximum (FWHM) is 110 MeV. So $\Delta E = 55$ MeV and using

$$\Delta E_{\min} \Delta t_{\min} = \frac{\hbar}{2},$$

$$\Delta t_{\min} = \frac{\hbar}{2\Delta E} = \frac{6.58 \times 10^{-16} \text{ eV} \cdot \text{s}}{2(55 \times 10^6 \text{ eV})} \approx 6.0 \times 10^{-24} \text{ s}$$

$$\tau = \text{lifetime} \sim 2\Delta t_{\min} = 1.2 \times 10^{-23} \text{ s}$$

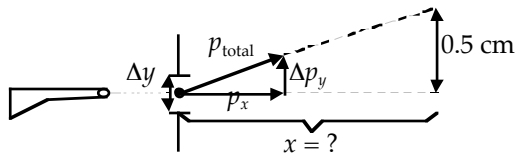
5-27 For a single slit with width a , minima are given by $\sin \theta = \frac{n\lambda}{a}$ where $n = 1, 2, 3, \dots$ and

$$\sin \theta \approx \tan \theta = \frac{x}{L}, \quad \frac{x_1}{L} = \frac{\lambda}{a} \quad \text{and} \quad \frac{x_2}{L} = \frac{2\lambda}{a} \Rightarrow \frac{x_2 - x_1}{L} = \frac{\lambda}{a} \quad \text{or}$$

$$\lambda = \frac{a\Delta x}{L} = \frac{5 \text{ \AA} \times 2.1 \text{ cm}}{20 \text{ cm}} = 0.525 \text{ \AA}$$

$$E = \frac{p^2}{2m} = \frac{h^2}{2m\lambda^2} = \frac{(hc)^2}{2mc^2\lambda^2} = \frac{(1.24 \times 10^4 \text{ eV} \cdot \text{\AA})^2}{2(5.11 \times 10^5 \text{ eV})(0.525 \text{ \AA})^2} = 546 \text{ eV}$$

5-31



$\Delta y \Delta p_y \sim \hbar$ $\Delta p_y = \frac{\hbar}{\Delta y}$. From the diagram, because the momentum triangle and space

triangle are similar, $\frac{\Delta p_y}{p_x} = \frac{0.5 \text{ cm}}{x}$;

$$x = \frac{(0.5 \text{ cm})p_x}{\Delta p_y} = \frac{(0.5 \text{ cm})p_x \Delta y}{\hbar} = \frac{(0.5 \times 10^{-2} \text{ m})(0.001 \text{ kg})(100 \text{ m/s})(2 \times 10^{-3} \text{ m})}{1.05 \times 10^{-34} \text{ J} \cdot \text{s}}$$

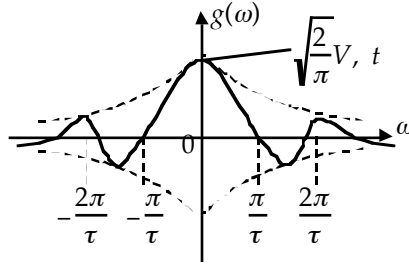
$$= 9.5 \times 10^{27} \text{ m}$$

Once again we see that the uncertainty relation has no observable consequences for macroscopic systems.

5-34 (a) $g(\omega) = (2\pi)^{-1/2} \int_{-\infty}^{\infty} V(t)(\cos \omega t - i \sin \omega t) dt$, $V(t) \sin \omega t$ is an odd function so this

integral vanishes leaving $g(\omega) = 2(2\pi)^{-1/2} \int_0^{\tau} V_0 \cos \omega t dt = \left(\frac{2}{\pi}\right)^{1/2} V_0 \frac{\sin \omega \tau}{\omega}$. A

sketch of $g(\omega)$ is given below.



- (b) As the major contribution to this pulse comes from ω 's between $-\frac{\pi}{\tau}$ and $\frac{\pi}{\tau}$, let

$$\Delta\omega \approx \frac{\pi}{\tau} \text{ and since } \Delta t = \tau.$$

$$\Delta\omega\Delta t = \left(\frac{\pi}{\tau}\right)\tau = \pi$$

- (c) Substituting $\Delta t = 0.5 \mu\text{s}$ in $\Delta\omega = \frac{\pi}{\Delta t}$ we find $\frac{\Delta 1}{2\Delta t} = \frac{1}{2(0.5 \times 10^{-6} \text{ s})} = 1 \times 10^6 \text{ Hz}$.

As the range is $2\Delta f$, the range is $2 \times 10^6 \text{ Hz}$. For $\Delta t = 0.5 \text{ ns}$, the range is $2\Delta f = 2 \times 10^9 \text{ Hz}$.

5-35 (a)

$$f(x) = \frac{1}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} a(k) e^{ikx} dk = \frac{A}{\sqrt{2\pi}} \int_{-\infty}^{+\infty} e^{-\alpha^2(k-k_0)^2} e^{ikx} dk = \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} \int_{-\infty}^{+\infty} e^{-\alpha^2(k^2 - (2k_0 + ix/\alpha^2)k)} dk$$

. Now complete the square in order to get the integral into the standard form

$$\int_{-\infty}^{+\infty} e^{-az^2} dz:$$

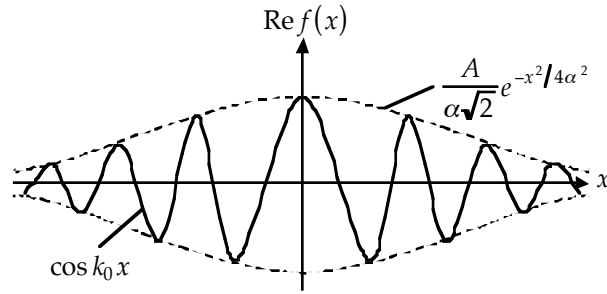
$$e^{-\alpha^2(k^2 - (2k_0 + ix/\alpha^2)k)} = e^{+\alpha^2(k_0 + ix/2\alpha^2)^2} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2}$$

$$\begin{aligned} f(x) &= \frac{A}{\sqrt{2\pi}} e^{-\alpha^2 k_0^2} e^{+\alpha^2(k_0 + ix/2\alpha^2)^2} \int_{k=-\infty}^{+\infty} e^{-\alpha^2(k - (k_0 + ix/2\alpha^2))^2} dk \\ &= \frac{A}{\sqrt{2\pi}} e^{-x^2/4\alpha^2} e^{ik_0 x} \int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz \end{aligned}$$

where $z = k - \left(k_0 + \frac{ix}{2\alpha^2}\right)$. Since $\int_{z=-\infty}^{+\infty} e^{-\alpha^2 z^2} dz = \frac{\pi^{1/2}}{\alpha}$, $f(x) = \frac{A}{\alpha\sqrt{2}} e^{-x^2/4\alpha^2} e^{ik_0 x}$. The

real part of $f(x)$, $\text{Re } f(x)$ is $\text{Re } f(x) = \frac{A}{\alpha\sqrt{2}} e^{-x^2/4\alpha^2} \cos k_0 x$ and is a gaussian

envelope multiplying a harmonic wave with wave number k_0 . A plot of $\text{Re } f(x)$ is shown below:



Comparing $\frac{A}{\alpha\sqrt{2}}e^{-x^2/4\alpha^2}$ to $Ae^{-(x/2\Delta x)^2}$ implies $\Delta x = \alpha$.

(c) By same reasoning because $\alpha^2 = \frac{1}{4\Delta k^2}$, $\Delta k = \frac{1}{2\alpha}$. Finally $\Delta x\Delta k = \alpha\left(\frac{1}{2\alpha}\right) = \frac{1}{2}$.

5-36 $E = K = \frac{1}{2}mu^2 = hf$ and $\lambda = \frac{h}{mu}$. $v_{\text{phase}} = f\lambda = \frac{mu^2}{2h} \frac{h}{mu} = \frac{u}{2} = v_{\text{phase}}$. This is different from the speed u at which the particle transports mass, energy, and momentum.

6-2 (a) Normalization requires

$$1 = \int_{-\infty}^{\infty} |\psi|^2 dx = A^2 \int_{-\frac{L}{4}}^{\frac{L}{4}} \cos^2\left(\frac{2\pi x}{L}\right) dx = \left(\frac{A^2}{2}\right) \int_{-\frac{L}{4}}^{\frac{L}{4}} \left(1 + \cos\left(\frac{4\pi x}{L}\right)\right) dx$$

$$\text{so } A = \frac{2}{\sqrt{L}}.$$

$$\begin{aligned} \text{(b)} \quad P &= \int_0^{\frac{L}{8}} |\psi|^2 dx = A^2 \int_0^{\frac{L}{8}} \cos^2\left(\frac{2\pi x}{L}\right) dx = \left(\frac{4}{L}\right) \left(\frac{1}{2}\right) \int_0^{\frac{L}{8}} \left(1 + \cos\left(\frac{4\pi x}{L}\right)\right) dx \\ &= \left(\frac{2}{L}\right) \left(\frac{L}{8}\right) + \left(\frac{2}{L}\right) \left(\frac{L}{4\pi}\right) \sin\left(\frac{4\pi x}{L}\right) \Bigg|_0^{\frac{L}{8}} = \frac{1}{4} + \frac{1}{2\pi} = 0.409 \end{aligned}$$

6-3 (a) $A \sin\left(\frac{2\pi x}{\lambda}\right) = A \sin(5 \times 10^{10} x)$ so $\left(\frac{2\pi}{\lambda}\right) = 5 \times 10^{10} \text{ m}^{-1}$,
 $\lambda = \frac{2\pi}{5 \times 10^{10}} = 1.26 \times 10^{-10} \text{ m}$.

$$\text{(b)} \quad p = \frac{h}{\lambda} = \frac{6.626 \times 10^{-34} \text{ Js}}{1.26 \times 10^{-10} \text{ m}} = 5.26 \times 10^{-24} \text{ kg m/s}$$

$$(c) \quad K = \frac{p^2}{2m} \quad m = 9.11 \times 10^{-31} \text{ kg}$$

$$K = \frac{(5.26 \times 10^{-24} \text{ kg m/s})^2}{(2 \times 9.11 \times 10^{-31} \text{ kg})} = 1.52 \times 10^{-17} \text{ J}$$

$$K = \frac{1.52 \times 10^{-17} \text{ J}}{1.6 \times 10^{-19} \text{ J/eV}} = 95 \text{ eV}$$

6-4 The time development of Ψ is given by Equation 6.8 or

$$\Psi(x, t) = \int a(k) e^{i\{kx - \omega(k)t\}} dk = \left(\frac{C\alpha}{\sqrt{\pi}} \right) \int_{-\infty}^{\infty} e^{\{ikx - i\omega(k)t - \alpha^2 k^2\}} dk,$$

with $\omega(k) = \frac{\hbar k^2}{2m}$ for a free particle of mass m . As in Example 6.3, the integral may be reduced to a recognizable form by completing the square in the exponent. Since $\omega(k)t = \left(\frac{\hbar t}{2m}\right)k^2$, we group this term together with $\alpha^2 k^2$ by introducing $\beta^2 = \alpha^2 + \frac{i\hbar t}{2m}$ to get

$$ikx - \omega(k)t - \alpha^2 k^2 = -\left(\beta k - \frac{ix}{2\beta}\right)^2 - \frac{x^2}{4\beta^2}.$$

Then, changing variables to $z = \beta k - \frac{ix}{2\beta}$ gives

$$\Psi(x, t) = \left(\frac{C\alpha}{\beta\sqrt{\pi}} \right) e^{-x^2/4\beta^2} \int_{-\infty}^{\infty} e^{-z^2} dz = \left(\frac{C\alpha}{\beta} \right) e^{-x^2/4\beta^2}.$$

To interpret this result, we must recognize that β is complex and separate real and imaginary parts. Thus, $|\beta^2|^2 = \left| \alpha^2 + \frac{i\hbar t}{2m} \right|^2 = \alpha^4 + \left(\frac{\hbar t}{2m} \right)^2$ and the exponent for Ψ is

$$\frac{x^2}{4\beta^2} = \frac{x^2 \left(\alpha^2 - \frac{i\hbar t}{2m} \right)}{4|\beta^2|^2} = \frac{x^2}{4 \left[\alpha^2 + \left(\frac{\hbar t}{2m\alpha} \right)^2 \right]} + (\text{imaginary terms})$$

then

$$|\Psi(x, t)| = \frac{C\alpha}{\left(\alpha^4 + \left(\frac{\hbar t}{2m} \right)^2 \right)^{1/4}} e^{\left\{ -x^2 / \left[4 \left(\alpha^2 + \left(\frac{\hbar t}{2m\alpha} \right)^2 \right) \right] \right\}}.$$

We see that apart from a phase factor, $\Psi(x, t)$ is still a gaussian but with amplitude

diminished by $\frac{\alpha}{\left(\alpha^4 + \left(\frac{\hbar t}{2m} \right)^2 \right)^{1/4}}$ and a width $\Delta x(t) = \left(\alpha^2 + \left(\frac{\hbar t}{2m\alpha} \right)^2 \right)^{1/2}$ where $\alpha = \Delta x(0)$ is

the initial width.