

Solutions Assignment 3

3.27 (a) The angular momentum of an object is $\vec{\ell} = \vec{r} \times \vec{p}$. Choosing the orbit of the planet to lie in the $x - y$ plane then \vec{r} and \vec{p} both lie in this plane as well. In polar coordinates we find that

$$\begin{aligned}\dot{\vec{r}} &= \frac{d}{dt}(r\hat{r}) = \frac{dr}{dt}\hat{r} + r\frac{d\phi}{dt}\frac{d}{d\phi}\hat{r}, \\ \dot{\vec{r}} &= \dot{r}\hat{r} + r\dot{\phi}\hat{\phi}.\end{aligned}$$

From the expressions for \hat{r} and $\hat{\phi}$ we find $\hat{r} \times \hat{\phi} = \hat{z}$. Thus the angular momentum is

$$\vec{\ell} = \vec{r} \times \vec{p} = mr^2\dot{\phi}\hat{z} = mr^2\omega\hat{z},$$

where $\omega = d\phi/dt$.

(b) For a differential angular displacement, $d\phi$, the area inside the triangle is

$$dA = \frac{1}{2}r^2d\phi.$$

Dividing by dt we find

$$\frac{dA}{dt} = \frac{1}{2}r^2\frac{d\phi}{dt} = \frac{1}{2}r^2\omega = \frac{\ell}{2m}.$$

Since angular momentum is conserved for central forces, we see that dA/dt is constant. Thus an orbiting planet sweeps out equal areas in equal times.

3.32 The moment of inertia of a uniform solid sphere of mass M and radius R rotating about an axis (taken to be the z axis) that passes through its center is found from

$$I = \int r_z^2 dM = \rho \int \rho_z^2 dV = \frac{3M}{4\pi R^3} \int_0^R \int_0^\pi \int_0^{2\pi} r_z^2 r^2 \sin\theta d\phi d\theta dr.$$

where r_z is the perpendicular distance from the z axis. In spherical coordinates $r_z = r \sin\theta$ and the momentum of inertia is

$$\begin{aligned}I &= \frac{3M}{4\pi R^3} \int_0^R \int_0^\pi \int_0^{2\pi} r^4 \sin^3\theta d\phi d\theta dr = \frac{3M}{2R^3} \int_0^R \int_0^\pi r^4 (1 - \cos^2\theta) \sin\theta d\theta dr, \\ I &= \frac{3M}{2R^3} \frac{R^5}{5} \left(2 - \frac{2}{3}\right) = \frac{2}{5}MR^2.\end{aligned}$$

3.35 (b) The torque about the point P , the point of contact between the disc and the inclined plane, for a disk of mass M and radius R is

$$\vec{\Gamma}^{ext} = \vec{r} \times \vec{F} = R\hat{y} \times Mg(\sin\gamma\hat{x} - \cos\gamma\hat{y}),$$

where γ is the angle of the plane relative to horizontal. Note that both the normal force and the frictional force pass through the point P and contribute no additional torques. We have defined the x axis to point down the inclined plane and the y axis normal to the plane. Performing the cross product we find

$$\dot{\vec{L}} = \vec{\Gamma}^{ext} = -MgR \sin \gamma \hat{z}.$$

The minus sign comes from defining the z axis to point out of the plane and from the right hand rule, the disc rotates in the negative z direction. The rate of change for the magnitude of the angular momentum is

$$\dot{L} = I\dot{\omega} = MgR \sin \gamma.$$

As stated in the problem, the momentum of inertia for a disk rotating about its circumference is $I = 3MR^2/2$, thus

$$\frac{3}{2}MR^2\dot{\omega} = MgR \sin \gamma \rightarrow \dot{\omega} = \frac{2 \sin \gamma}{3R}.$$

From the no slip condition

$$\dot{v} = R\dot{\omega} = \frac{2}{3} \sin \gamma.$$

(c) The torque about the CM for the disc is

$$\vec{\Gamma}^{ext} = \vec{r} \times \vec{F} = -R\hat{y} \times F_f(-\hat{x}) = -RF_f\hat{z},$$

where F_f is the magnitude of the frictional force. From the x component of Newton's second law

$$M\dot{v} = Mg \sin \gamma - F_f \rightarrow F_f = M\dot{v} - Mg \sin \gamma.$$

Since the momentum of inertia for the disc about its CM is $I = MR^2/2$ we have (including the no slip condition, $R\dot{\omega} = \dot{v}$)

$$\begin{aligned} \Gamma^{ext} &= \dot{L} = I\dot{\omega} = \frac{1}{2}MR^2\dot{\omega} = -R(M\dot{v} - Mg \sin \gamma) \\ \frac{1}{2}M\dot{v} &= -M\dot{v} + Mg \sin \gamma \\ \dot{v} &= \frac{2}{3} \sin \gamma. \end{aligned}$$

Note that in keeping track of signs, $\dot{\omega}$ is defined to be positive when the disk is rotating so that it proceeds up the plane (think of the right hand rule).

3.37 The derivation of

$$\frac{d}{dt} \vec{L} \text{ (about CM)} = \vec{\Gamma}^{\text{ext}} \text{ (about CM)},$$

was done in the discussion session, but is repeated here for reference. Defining \vec{r}'_{α} as the position of mass particle α relative to the center of mass, $\vec{r}'_{\alpha} = \vec{r}_{\alpha} - \vec{R}$, the rate of change of the angular momentum about the center of mass (CM) is expressed as

$$\begin{aligned} \frac{d}{dt} \vec{L}_{\text{CM}} &= \sum_{\alpha} \frac{d}{dt} \left(\vec{r}'_{\alpha} \times m_{\alpha} \dot{\vec{r}}'_{\alpha} \right) = \sum_{\alpha} \vec{r}'_{\alpha} \times m_{\alpha} \ddot{\vec{r}}'_{\alpha}, \\ \frac{d}{dt} \vec{L}_{\text{CM}} &= \sum_{\alpha} m_{\alpha} \vec{r}'_{\alpha} \times \left(\ddot{\vec{r}}_{\alpha} - \ddot{\vec{R}} \right), \\ \frac{d}{dt} \vec{L}_{\text{CM}} &= \sum_{\alpha} \vec{r}'_{\alpha} \times \vec{F}_{\alpha} - \left(\sum_{\alpha} m_{\alpha} \vec{r}'_{\alpha} \right) \times \ddot{\vec{R}}. \end{aligned}$$

It is very straightforward to show that from the definition of the CM that $\sum_{\alpha} m_{\alpha} \vec{r}'_{\alpha} = 0$. Thus the expression reduces to

$$\frac{d}{dt} \vec{L}_{\text{CM}} = \sum_{\alpha} \vec{r}'_{\alpha} \times \vec{F}_{\alpha} = \vec{\Gamma}_{\text{CM}}^{\text{ext}}$$

Note that this is true even if the CM is accelerating, $\ddot{\vec{R}} \neq 0$!

4.3 (a) First consider the line integral for $\vec{F} = -y\hat{x} + x\hat{y}$ along the path POQ . This path integral is expressed as

$$W = \int_P^Q \vec{F} \cdot d\vec{r} = - \int_1^0 (y=0) dx + \int_0^1 (x=0) dy = 0.$$

(b) Now consider the line integral for $\vec{F} = -y\hat{x} + x\hat{y}$ along the path PQ . This path integral is expressed as

$$\begin{aligned} W &= \int_P^Q \vec{F} \cdot d\vec{r} = \int_{(1,0)}^{(0,1)} (F_x(x,y) dx + F_y(x,y) dy), \\ W &= \int_{(0,1)}^{(0,1)} -y dx + x dy. \end{aligned}$$

Along this path, $y = 1 - x$, so that

$$W = \int_1^0 -(1-x) dx - \int_1^0 x dx = - \int_1^0 dx = 1.$$

(c) Now consider the line integral for $\vec{F} = -y\hat{x} + x\hat{y}$ along the circular path PQ . First we will rewrite the force in polar coordinates along the unit circle. In these coordinates $x = \cos \phi$ and $y = \sin \phi$; and $\hat{x} = \cos \phi \hat{r} - \sin \phi \hat{\phi}$ and $\hat{y} = \sin \phi \hat{r} + \cos \phi \hat{\phi}$ so that

$$\begin{aligned}\vec{F} &= -y\hat{x} + x\hat{y} = -\sin \phi (\cos \phi \hat{r} - \sin \phi \hat{\phi}) + \cos \phi (\sin \phi \hat{r} + \cos \phi \hat{\phi}), \\ \vec{F} &= \hat{\phi}.\end{aligned}$$

The path integral then becomes

$$W = \int_P^Q \vec{F} \cdot d\vec{r} = \int_0^{\pi/2} d\phi = \pi/2.$$

4.4 (a) Since it is a radial force, the angular momentum is conserved so that

$$\ell = mr_o^2 \omega_o = mr^2 \omega \rightarrow \omega = \frac{\ell}{mr^2} = \frac{r_o^2}{r^2} \omega_o.$$

(b) From lecture 2, we know that (assuming \ddot{r} is small) the force on the string is $F = ma_\phi = mr\omega^2$. The work pulling on the string is then

$$W = \int_{r_o}^r F dr = - \int_{r_o}^r mr\omega^2 dr = - \int_{r_o}^r \frac{\ell^2}{mr^3} dr = \frac{1}{2} \frac{\ell^2}{m} \left(\frac{1}{r^2} - \frac{1}{r_o^2} \right) = \frac{1}{2} mr_o^4 \omega_o^2 \left(\frac{1}{r^2} - \frac{1}{r_o^2} \right),$$

where the minus sign comes from F being in the opposite direction of dr .

(c) The change in KE is

$$\Delta KE = \frac{1}{2} \frac{\ell^2}{m} \left(\frac{1}{r^2} - \frac{1}{r_o^2} \right) = \frac{1}{2} mr_o^4 \omega_o^2 \left(\frac{1}{r^2} - \frac{1}{r_o^2} \right),$$

which is the same as the work done as it had to be.

4.7 (a) The force of gravity on Planet X is $\vec{F} = -m\gamma y^2 \hat{y}$. The work done by gravity moving the mass m from \vec{r}_1 to \vec{r}_2 is

$$\begin{aligned}W(\vec{r}_1 \rightarrow \vec{r}_2) &= \int_{\vec{r}_1}^{\vec{r}_2} \vec{F} \cdot d\vec{r} = - \int_{\vec{r}_1}^{\vec{r}_2} m\gamma y^2 dy = - \int_{y_1}^{y_2} m\gamma y^2 dy, \\ W(\vec{r}_1 \rightarrow \vec{r}_2) &= -\frac{1}{3} m\gamma (y_2^3 - y_1^3).\end{aligned}$$

Since the work done only depends on the end points, it is a conservative force. The potential energy for this gravitational field is

$$U(y) = \frac{1}{3} m\gamma y^3$$

(c) The energy for a stationary mass at a height h is

$$E = \frac{1}{3} m\gamma h^3 = \frac{1}{2} m\dot{y}^2 + \frac{1}{3} m\gamma y^3,$$

where y is measured from the ground. When $y = 0$, the velocity is

$$\dot{y} = \sqrt{2\gamma h^3/3}.$$

4.18 (a) From equation (4.35) in the text $df = \nabla f \cdot d\vec{r}$. If the differential displacement vector $d\vec{r}$ lies in a surface defined by $f = \text{const.}$ then $df = 0$. From that we see

$$\nabla f \cdot d\vec{r} = 0,$$

when $d\vec{r}$ lies in a surface defined by $f = \text{const.}$, thus ∇f is \perp to a surface of constant f .

(b) Now let $d\vec{r} = \epsilon \hat{u}$ where ϵ is small and \hat{u} is a unit vector that points in an arbitrary direction. Thus

$$df = \nabla f \cdot d\vec{r} = \epsilon \nabla f \cdot \hat{u} = \epsilon |\nabla f| \cos \theta,$$

where df is the change in f in the direction of \hat{u} . This is a maximum when $\cos \theta = 1$ which occurs when \hat{u} points in the direction of ∇f .

4.19 (a) A surface defined by $f = x^2 + 4y^2 = \text{const.}$ is a elliptic surface. The intersection of any plane orthogonal to the z axis with this surface will form an ellipse with the semimajor axis, which is parallel to the x axis, being twice that of the semiminor axis, which is parallel to the y axis.

(b) The unit normal to this surface is in the direction of ∇f which yields

$$\begin{aligned} \nabla f &= 2x\hat{x} + 8y\hat{y} \rightarrow \hat{n} = \frac{x\hat{x} + 4y\hat{y}}{\sqrt{x^2 + 16y^2}} \\ \hat{n}(1, 1, 1) &= \frac{\hat{x} + 4\hat{y}}{\sqrt{17}}. \end{aligned}$$

Moving in the direction of \hat{n} will maximize the rate of change in f .

4.23 First all three forces only depend on position. It only remains to check $\nabla \times \vec{F}$.

(a) If $\vec{F} = kx\hat{x} + 2ky\hat{y} + 3kz\hat{z}$, then

$$\nabla \times \vec{F} = \left(\frac{\partial F_z}{\partial y} - \frac{\partial F_y}{\partial z} \right) \hat{x} + \left(\frac{\partial F_x}{\partial z} - \frac{\partial F_z}{\partial x} \right) \hat{y} + \left(\frac{\partial F_y}{\partial x} - \frac{\partial F_x}{\partial y} \right) \hat{z} = 0.$$

This force is conservative. The corresponding potential is

$$U = -k \left(\frac{1}{2}x^2 + y^2 + \frac{3}{2}z^2 \right).$$

(b) If $\vec{F} = ky\hat{x} + kx\hat{y}$, then

$$\begin{aligned} \nabla \times \vec{F} &= \left(\frac{\partial F_z}{\partial y} - \frac{\partial F_y}{\partial z} \right) \hat{x} + \left(\frac{\partial F_x}{\partial z} - \frac{\partial F_z}{\partial x} \right) \hat{y} + \left(\frac{\partial F_y}{\partial x} - \frac{\partial F_x}{\partial y} \right) \hat{z}, \\ \nabla \times \vec{F} &= k \left(\frac{\partial x}{\partial x} - \frac{\partial y}{\partial y} \right) \hat{z} = 0. \end{aligned}$$

This force is conservative. The potential is

$$U(x, y) = -kxy.$$

(c) If $\vec{F} = -ky\hat{x} + kx\hat{y}$, then

$$\nabla \times \vec{F} = \left(\frac{\partial F_z}{\partial y} - \frac{\partial F_y}{\partial z} \right) \hat{x} + \left(\frac{\partial F_x}{\partial z} - \frac{\partial F_z}{\partial x} \right) \hat{y} + \left(\frac{\partial F_y}{\partial x} - \frac{\partial F_x}{\partial y} \right) \hat{z},$$

$$\nabla \times \vec{F} = \left(k \frac{\partial x}{\partial x} + k \frac{\partial y}{\partial y} \right) \hat{z} = 2k\hat{z} \neq 0.$$

This force is not conservative.

4.34 (a) The vertical distance that the mass hangs down on a pendulum of length ℓ is $\ell \cos \phi$. Since at equilibrium it is a length ℓ below the pivot the potential energy measured above equilibrium is

$$U(\phi) = mg\ell(1 - \cos \phi).$$

The total energy is

$$E = \frac{1}{2}m \left(\ell \dot{\phi} \right)^2 + mg\ell(1 - \cos \phi).$$

(b) Differentiating this expression wrt t yields

$$m\ell^2 \dot{\phi} \ddot{\phi} + mg\ell \sin \phi \dot{\phi} = 0,$$

or

$$m\ell^2 \ddot{\phi} + mg\ell \sin \phi = 0 \rightarrow I\alpha = -F\ell \sin \phi = -|\vec{r} \times mg\hat{y}| = \Gamma.$$

(c) If $\phi \ll 1$ then $\sin \phi \simeq \phi$ and our EOM becomes

$$\ddot{\phi} + \frac{g}{\ell} \phi = 0.$$

This differential equation has as solutions

$$\phi = A \sin \omega t + B \cos \omega t,$$

where $\omega^2 = g/\ell$ or a period of $\tau = 2\pi\sqrt{\ell/g}$.