

Chapter 7

Noether's Theorem

7.1 Continuous Symmetry Implies Conserved Charges

Consider a particle moving in two dimensions under the influence of an external potential $U(r)$. The potential is a function only of the magnitude of the vector \mathbf{r} . The Lagrangian is then

$$L = T - U = \frac{1}{2}m (\dot{r}^2 + r^2 \dot{\phi}^2) - U(r) , \quad (7.1)$$

where we have chosen generalized coordinates (r, ϕ) . The momentum conjugate to ϕ is $p_\phi = m r^2 \dot{\phi}$. The generalized force F_ϕ clearly vanishes, since L does not depend on the coordinate ϕ . (One says that L is 'cyclic' in ϕ .) Thus, although $r = r(t)$ and $\phi = \phi(t)$ will in general be time-dependent, the combination $p_\phi = m r^2 \dot{\phi}$ is constant. This is the conserved angular momentum about the \hat{z} axis.

In general, whenever the system exhibits a *continuous symmetry*, there is an associated *conserved charge*. (The terminology 'charge' is from field theory.) Indeed, this is a rigorous result, known as *Noether's Theorem*. Consider a one-parameter family of transformations,

$$q_\sigma \longrightarrow \tilde{q}_\sigma(q, \zeta) , \quad (7.2)$$

where ζ is the continuous parameter. Suppose further (without loss of generality) that at $\zeta = 0$ this transformation is the identity, *i.e.* $\tilde{q}_\sigma(q, \zeta) = q_\sigma$. The transformation may be nonlinear in the generalized coordinates. Suppose further that the Lagrangian

L is invariant under the replacement $q \rightarrow \tilde{q}$. Then we must have

$$\begin{aligned}
0 &= \frac{d}{d\zeta} \Big|_{\zeta=0} L(\tilde{q}, \dot{\tilde{q}}, t) = \frac{\partial L}{\partial q_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \Big|_{\zeta=0} + \frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \dot{\tilde{q}}_\sigma}{\partial \zeta} \Big|_{\zeta=0} \\
&= \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \right) \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \Big|_{\zeta=0} + \frac{\partial L}{\partial \dot{q}_\sigma} \frac{d}{dt} \left(\frac{\partial \tilde{q}_\sigma}{\partial \zeta} \right) \Big|_{\zeta=0} \\
&= \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \right) \Big|_{\zeta=0}.
\end{aligned} \tag{7.3}$$

Thus, there is an associated conserved charge

$$\Lambda = \frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \Big|_{\zeta=0}. \tag{7.4}$$

7.1.1 Examples of one-parameter families of transformations

Consider the Lagrangian

$$L = \frac{1}{2}m(\dot{x}^2 + \dot{y}^2) - U(\sqrt{x^2 + y^2}). \tag{7.5}$$

In two-dimensional polar coordinates, we have

$$L = \frac{1}{2}m(\dot{r}^2 + r^2\dot{\phi}^2) - U(r), \tag{7.6}$$

and we may now define

$$\tilde{r}(\zeta) = r \tag{7.7}$$

$$\tilde{\phi}(\zeta) = \phi + \zeta. \tag{7.8}$$

Note that $\tilde{r}(0) = r$ and $\tilde{\phi}(0) = \phi$, *i.e.* the transformation is the identity when $\zeta = 0$. We now have

$$\Lambda = \sum_{\sigma} \frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \Big|_{\zeta=0} = \frac{\partial L}{\partial \dot{r}} \frac{\partial \tilde{r}}{\partial \zeta} \Big|_{\zeta=0} + \frac{\partial L}{\partial \dot{\phi}} \frac{\partial \tilde{\phi}}{\partial \zeta} \Big|_{\zeta=0} = mr^2\dot{\phi}. \tag{7.9}$$

Another way to derive the same result which is somewhat instructive is to work out the transformation in Cartesian coordinates. We then have

$$\tilde{x}(\zeta) = x \cos \zeta - y \sin \zeta \tag{7.10}$$

$$\tilde{y}(\zeta) = x \sin \zeta + y \cos \zeta. \tag{7.11}$$

Thus,

$$\frac{\partial \tilde{x}}{\partial \zeta} = -y(\zeta) \quad , \quad \frac{\partial \tilde{y}}{\partial \zeta} = x(\zeta) \quad (7.12)$$

and

$$\Lambda = \left. \frac{\partial L}{\partial \dot{x}} \frac{\partial \tilde{x}}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial L}{\partial \dot{y}} \frac{\partial \tilde{y}}{\partial \zeta} \right|_{\zeta=0} = m(x\dot{y} - y\dot{x}) . \quad (7.13)$$

But

$$m(x\dot{y} - y\dot{x}) = m\hat{\mathbf{z}} \cdot \mathbf{r} \times \dot{\mathbf{r}} = mr^2\dot{\phi} . \quad (7.14)$$

As another example, consider the potential

$$U(\rho, \phi, z) = V(\rho, a\phi + z) , \quad (7.15)$$

where (ρ, ϕ, z) are cylindrical coordinates for a particle of mass m , and where a is a constant with dimensions of length. The Lagrangian is

$$\frac{1}{2}m(\dot{\rho}^2 + \rho^2\dot{\phi}^2 + \dot{z}^2) - V(\rho, a\phi + z) . \quad (7.16)$$

This model possesses a helical symmetry, with a one-parameter family

$$\tilde{\rho}(\zeta) = \rho \quad (7.17)$$

$$\tilde{\phi}(\zeta) = \phi + \zeta \quad (7.18)$$

$$\tilde{z}(\zeta) = z - \zeta a . \quad (7.19)$$

Note that

$$a\tilde{\phi} + \tilde{z} = a\phi + z , \quad (7.20)$$

so the potential energy, and the Lagrangian as well, is invariant under this one-parameter family of transformations. The conserved charge for this symmetry is

$$\Lambda = \left. \frac{\partial L}{\partial \dot{\rho}} \frac{\partial \tilde{\rho}}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial L}{\partial \dot{\phi}} \frac{\partial \tilde{\phi}}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial L}{\partial \dot{z}} \frac{\partial \tilde{z}}{\partial \zeta} \right|_{\zeta=0} = m\rho^2\dot{\phi} - ma\dot{z} . \quad (7.21)$$

We can check explicitly that Λ is conserved, using the equations of motion

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{\phi}} = \frac{d}{dt} (m\rho^2\dot{\phi}) = \frac{\partial L}{\partial \phi} = -a \frac{\partial V}{\partial z} \quad (7.22)$$

$$\frac{d}{dt} \frac{\partial L}{\partial \dot{z}} = \frac{d}{dt} (m\dot{z}) = \frac{\partial L}{\partial z} = -\frac{\partial V}{\partial z} . \quad (7.23)$$

Thus,

$$\dot{\Lambda} = \frac{d}{dt} (m\rho^2\dot{\phi}) - a \frac{d}{dt} (m\dot{z}) = 0 . \quad (7.24)$$

7.2 Conservation of Linear and Angular Momentum

Suppose that the Lagrangian of a mechanical system is invariant under a uniform translation of all particles in the $\hat{\mathbf{n}}$ direction. Then our one-parameter family of transformations is given by

$$\tilde{\mathbf{x}}_a = \mathbf{x}_a + \zeta \hat{\mathbf{n}} , \quad (7.25)$$

and the associated conserved Noether charge is

$$\Lambda = \sum_a \frac{\partial L}{\partial \dot{\mathbf{x}}_a} \cdot \hat{\mathbf{n}} = \hat{\mathbf{n}} \cdot \mathbf{P} , \quad (7.26)$$

where $\mathbf{P} = \sum_a \mathbf{p}_a$ is the *total momentum* of the system.

If the Lagrangian of a mechanical system is invariant under rotations about an axis $\hat{\mathbf{n}}$, then

$$\begin{aligned} \tilde{\mathbf{x}}_a &= R(\zeta, \hat{\mathbf{n}}) \mathbf{x}_a \\ &= \mathbf{x}_a + \zeta \hat{\mathbf{n}} \times \mathbf{x}_a + \mathcal{O}(\zeta^2) , \end{aligned} \quad (7.27)$$

where we have expanded the rotation matrix $R(\zeta, \hat{\mathbf{n}})$ in powers of ζ . The conserved Noether charge associated with this symmetry is

$$\Lambda = \sum_a \frac{\partial L}{\partial \dot{\mathbf{x}}_a} \cdot \hat{\mathbf{n}} \times \mathbf{x}_a = \hat{\mathbf{n}} \cdot \sum_a \mathbf{x}_a \times \mathbf{p}_a = \hat{\mathbf{n}} \cdot \mathbf{L} , \quad (7.28)$$

where \mathbf{L} is the *total angular momentum* of the system.

7.3 Advanced discussion : Invariance of L vs. invariance of S

Observant readers might object that demanding invariance of L is too strict. We should instead be demanding invariance of the action S ¹. Suppose S is invariant under

$$t \rightarrow \tilde{t}(q, t, \zeta) \quad (7.29)$$

$$q_\sigma(t) \rightarrow \tilde{q}_\sigma(q, t, \zeta) . \quad (7.30)$$

¹Indeed, we should be demanding that S only change by a function of the endpoint values.

Then invariance of S means

$$S = \int_{t_a}^{t_b} dt L(q, \dot{q}, t) = \int_{\tilde{t}_a}^{\tilde{t}_b} dt L(\tilde{q}, \dot{\tilde{q}}, t) . \quad (7.31)$$

Note that t is a dummy variable of integration, so it doesn't matter whether we call it t or \tilde{t} . The endpoints of the integral, however, do change under the transformation. Now consider an infinitesimal transformation, for which $\delta t = \tilde{t} - t$ and $\delta q = \tilde{q}(\tilde{t}) - q(t)$ are both small. Invariance of S means

$$S = \int_{t_a}^{t_b} dt L(q, \dot{q}, t) = \int_{t_a + \delta t_a}^{t_b + \delta t_b} dt \left\{ L(q, \dot{q}, t) + \frac{\partial L}{\partial q_\sigma} \bar{\delta} q_\sigma + \frac{\partial L}{\partial \dot{q}_\sigma} \bar{\delta} \dot{q}_\sigma + \dots \right\} , \quad (7.32)$$

where

$$\begin{aligned} \bar{\delta} q_\sigma(t) &\equiv \tilde{q}_\sigma(t) - q_\sigma(t) \\ &= \tilde{q}_\sigma(\tilde{t}) - \tilde{q}_\sigma(\tilde{t}) + \tilde{q}_\sigma(t) - q_\sigma(t) \\ &= \delta q_\sigma - \dot{q}_\sigma \delta t + \mathcal{O}(\delta q \delta t) \end{aligned} \quad (7.33)$$

Subtracting the top line from the bottom, we obtain

$$\begin{aligned} 0 &= L_b \delta t_b - L_a \delta t_a + \frac{\partial L}{\partial \dot{q}_\sigma} \Big|_b \bar{\delta} q_{\sigma,b} - \frac{\partial L}{\partial \dot{q}_\sigma} \Big|_a \bar{\delta} q_{\sigma,a} + \int_{t_a + \delta t_a}^{t_b + \delta t_b} dt \left\{ \frac{\partial L}{\partial q_\sigma} - \frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \right) \right\} \bar{\delta} q(t) \\ &= \int_{t_a}^{t_b} dt \frac{d}{dt} \left\{ \left(L - \frac{\partial L}{\partial \dot{q}_\sigma} \dot{q}_\sigma \right) \delta t + \frac{\partial L}{\partial \dot{q}_\sigma} \delta q_\sigma \right\} . \end{aligned} \quad (7.34)$$

Thus, if $\zeta \equiv \delta \zeta$ is infinitesimal, and

$$\delta t = A(q, t) \delta \zeta \quad (7.35)$$

$$\delta q_\sigma = B_\sigma(q, t) \delta \zeta , \quad (7.36)$$

then the conserved charge is

$$\begin{aligned} \Lambda &= \left(L - \frac{\partial L}{\partial \dot{q}_\sigma} \dot{q}_\sigma \right) A(q, t) + \frac{\partial L}{\partial \dot{q}_\sigma} B_\sigma(q, t) \\ &= -H(q, p, t) A(q, t) + p_\sigma B_\sigma(q, t) . \end{aligned} \quad (7.37)$$

Thus, when $A = 0$, we recover our earlier results, obtained by assuming invariance of L . Note that conservation of H follows from time translation invariance: $t \rightarrow t + \zeta$, for which $A = 1$ and $B_\sigma = 0$. Here we have written

$$H = p_\sigma \dot{q}_\sigma - L , \quad (7.38)$$

and expressed it in terms of the momenta p_σ , the coordinates q_σ , and time t . H is called the *Hamiltonian*.

7.3.1 The Hamiltonian

The Lagrangian is a function of generalized coordinates, velocities, and time. The canonical momentum conjugate to the generalized coordinate q_σ is

$$p_\sigma = \frac{\partial L}{\partial \dot{q}_\sigma} . \quad (7.39)$$

The Hamiltonian is a function of coordinates, momenta, and time. It is defined as the Legendre transform of L :

$$H(q, p, t) = \sum_\sigma p_\sigma \dot{q}_\sigma - L . \quad (7.40)$$

Let's examine the differential of H :

$$\begin{aligned} dH &= \sum_\sigma \left(\dot{q}_\sigma dp_\sigma + p_\sigma d\dot{q}_\sigma - \frac{\partial L}{\partial q_\sigma} dq_\sigma - \frac{\partial L}{\partial \dot{q}_\sigma} d\dot{q}_\sigma \right) - \frac{\partial L}{\partial t} dt \\ &= \sum_\sigma \left(\dot{q}_\sigma dp_\sigma - \frac{\partial L}{\partial q_\sigma} dq_\sigma \right) - \frac{\partial L}{\partial t} dt , \end{aligned} \quad (7.41)$$

where we have invoked the definition of p_σ to cancel the coefficients of $d\dot{q}_\sigma$. Since $\dot{p}_\sigma = \partial L / \partial q_\sigma$, we have *Hamilton's equations of motion*,

$$\dot{q}_\sigma = \frac{\partial H}{\partial p_\sigma} , \quad \dot{p}_\sigma = -\frac{\partial H}{\partial q_\sigma} . \quad (7.42)$$

Thus, we can write

$$dH = \sum_\sigma \left(\dot{q}_\sigma dp_\sigma - \dot{p}_\sigma dq_\sigma \right) - \frac{\partial H}{\partial t} dt . \quad (7.43)$$

Dividing by dt , we obtain

$$\frac{dH}{dt} = -\frac{\partial H}{\partial t} , \quad (7.44)$$

which says that the Hamiltonian is *conserved* (i.e. it does not change with time) whenever there is no *explicit* time dependence to L .

Example #1 : For a simple $d = 1$ system with $L = \frac{1}{2}m\dot{x}^2 - U(x)$, we have $p = m\dot{x}$ and

$$H = p\dot{x} - L = \frac{1}{2}m\dot{x}^2 + U(x) = \frac{p^2}{2m} + U(x) . \quad (7.45)$$

Example #2 : Consider now the mass point – wedge system analyzed above, with

$$L = \frac{1}{2}(M + m)\dot{X}^2 + m\dot{X}\dot{x} + \frac{1}{2}m(1 + \tan^2\alpha)\dot{x}^2 - mgx \tan\alpha , \quad (7.46)$$

The canonical momenta are

$$P = \frac{\partial L}{\partial \dot{X}} = (M + m)\dot{X} + m\dot{x} \quad (7.47)$$

$$p = \frac{\partial L}{\partial \dot{x}} = m\dot{X} + m(1 + \tan^2\alpha)\dot{x} . \quad (7.48)$$

The Hamiltonian is given by

$$\begin{aligned} H &= P\dot{X} + p\dot{x} - L \\ &= \frac{1}{2}(M + m)\dot{X}^2 + m\dot{X}\dot{x} + \frac{1}{2}m(1 + \tan^2\alpha)\dot{x}^2 + mgx \tan\alpha . \end{aligned} \quad (7.49)$$

However, this is not quite H , since $H = H(X, x, P, p, t)$ must be expressed in terms of the coordinates and the *momenta* and not the coordinates and velocities. So we must eliminate \dot{X} and \dot{x} in favor of P and p . We do this by inverting the relations

$$\begin{pmatrix} P \\ p \end{pmatrix} = \begin{pmatrix} M + m & m \\ m & m(1 + \tan^2\alpha) \end{pmatrix} \begin{pmatrix} \dot{X} \\ \dot{x} \end{pmatrix} \quad (7.50)$$

to obtain

$$\begin{pmatrix} \dot{X} \\ \dot{x} \end{pmatrix} = \frac{1}{m(M + (M + m)\tan^2\alpha)} \begin{pmatrix} m(1 + \tan^2\alpha) & -m \\ -m & M + m \end{pmatrix} \begin{pmatrix} P \\ p \end{pmatrix} . \quad (7.51)$$

Substituting into 7.49, we obtain

$$H = \frac{M + m}{2m} \frac{P^2 \cos^2\alpha}{M + m \sin^2\alpha} - \frac{Pp \cos^2\alpha}{M + m \sin^2\alpha} + \frac{p^2}{2(M + m \sin^2\alpha)} + mgx \tan\alpha . \quad (7.52)$$

Notice that $\dot{P} = 0$ since $\frac{\partial L}{\partial X} = 0$. P is the total horizontal momentum of the system (wedge plus particle) and it is conserved.

7.3.2 Is $H = T + U$?

The most general form of the kinetic energy is

$$\begin{aligned} T &= T_2 + T_1 + T_0 \\ &= \frac{1}{2}T_{\sigma\sigma'}^{(2)}(q, t) \dot{q}_\sigma \dot{q}_{\sigma'} + T_\sigma^{(1)}(q, t) \dot{q}_\sigma + T^{(0)}(q, t) , \end{aligned} \quad (7.53)$$

where $T_n(q, \dot{q}, t)$ is homogeneous of degree n in the velocities². We assume a potential energy of the form

$$\begin{aligned} U &= U_1 + U_0 \\ &= U_\sigma^{(1)}(q, t) \dot{q}_\sigma + U^{(0)}(q, t) , \end{aligned} \quad (7.54)$$

which allows for velocity-dependent forces, as we have with charged particles moving in an electromagnetic field. The Lagrangian is then

$$L = T - U = \frac{1}{2}T_{\sigma\sigma'}^{(2)}(q, t) \dot{q}_\sigma \dot{q}_{\sigma'} + T_\sigma^{(1)}(q, t) \dot{q}_\sigma + T^{(0)}(q, t) - U_\sigma^{(1)}(q, t) \dot{q}_\sigma - U^{(0)}(q, t) . \quad (7.55)$$

We have assumed $U(q, t)$ is velocity-independent, but the above form for $L = T - U$ is quite general. (*E.g.* any velocity-dependence in U can be absorbed into the $B_\sigma \dot{q}_\sigma$ term.) The canonical momentum conjugate to q_σ is

$$p_\sigma = \frac{\partial L}{\partial \dot{q}_\sigma} = T_{\sigma\sigma'}^{(2)} \dot{q}_{\sigma'} + T_\sigma^{(1)}(q, t) - U_\sigma^{(1)}(q, t) \quad (7.56)$$

which is inverted to give

$$\dot{q}_\sigma = T_{\sigma\sigma'}^{(2)-1} \left(p_{\sigma'} - T_{\sigma'}^{(1)} + U_{\sigma'}^{(1)} \right) . \quad (7.57)$$

The Hamiltonian is then

$$\begin{aligned} H &= p_\sigma \dot{q}_\sigma - L \\ &= \frac{1}{2} T_{\sigma\sigma'}^{(2)-1} \left(p_\sigma - T_\sigma^{(1)} + U_\sigma^{(1)} \right) \left(p_{\sigma'} - T_{\sigma'}^{(1)} + U_{\sigma'}^{(1)} \right) - T_0 + U_0 \end{aligned} \quad (7.58)$$

$$= T_2 - T_0 + U_0 . \quad (7.59)$$

If T_0 , T_1 , and U_1 vanish, *i.e.* if $T(q, \dot{q}, t)$ is a homogeneous function of degree two in the generalized velocities, and $U(q, t)$ is velocity-independent, then $H = T + U$. But if T_0 or T_1 is nonzero, or the potential is velocity-dependent, then $H \neq T + U$.

²A homogeneous function of degree k satisfies $f(\lambda x_1, \dots, \lambda x_n) = \lambda^k f(x_1, \dots, x_n)$. It is then easy to prove *Euler's theorem*, $\sum_{i=1}^n x_i \frac{\partial f}{\partial x_i} = k f$.

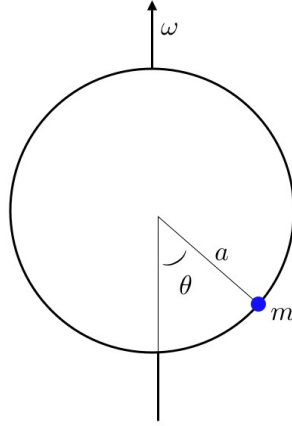


Figure 7.1: A bead of mass m on a rotating hoop of radius a .

7.3.3 Example: A Bead on a Rotating Hoop

Consider a bead of mass m constrained to move along a hoop of radius a . The hoop is further constrained to rotate with angular velocity ω about the \hat{z} -axis, as shown in Fig. 7.1.

The most convenient set of generalized coordinates is spherical polar (r, θ, ϕ) , in which case

$$\begin{aligned} T &= \frac{1}{2}m(\dot{r}^2 + r^2\dot{\theta}^2 + r^2\sin^2\theta\dot{\phi}^2) \\ &= \frac{1}{2}ma^2(\dot{\theta}^2 + \omega^2\sin^2\theta) . \end{aligned} \quad (7.60)$$

Thus, $T_2 = \frac{1}{2}ma^2\dot{\theta}^2$ and $T_0 = \frac{1}{2}ma^2\omega^2\sin^2\theta$. The potential energy is $U(\theta) = mga(1 - \cos\theta)$. The momentum conjugate to θ is $p_\theta = ma^2\dot{\theta}$, and thus

$$\begin{aligned} H(\theta, p) &= T_2 - T_0 + U \\ &= \frac{1}{2}ma^2\dot{\theta}^2 - \frac{1}{2}ma^2\omega^2\sin^2\theta + mga(1 - \cos\theta) \\ &= \frac{p_\theta^2}{2ma^2} - \frac{1}{2}ma^2\omega^2\sin^2\theta + mga(1 - \cos\theta) . \end{aligned} \quad (7.61)$$

For this problem, we can define the *effective potential*

$$\begin{aligned} U_{\text{eff}}(\theta) &\equiv U - T_0 = mga(1 - \cos\theta) - \frac{1}{2}ma^2\omega^2\sin^2\theta \\ &= mga\left(1 - \cos\theta - \frac{\omega^2}{2\omega_0^2}\sin^2\theta\right) , \end{aligned} \quad (7.62)$$

where $\omega_0 \equiv g/a^2$. The Lagrangian may then be written

$$L = \frac{1}{2}ma^2\dot{\theta}^2 - U_{\text{eff}}(\theta) , \quad (7.63)$$

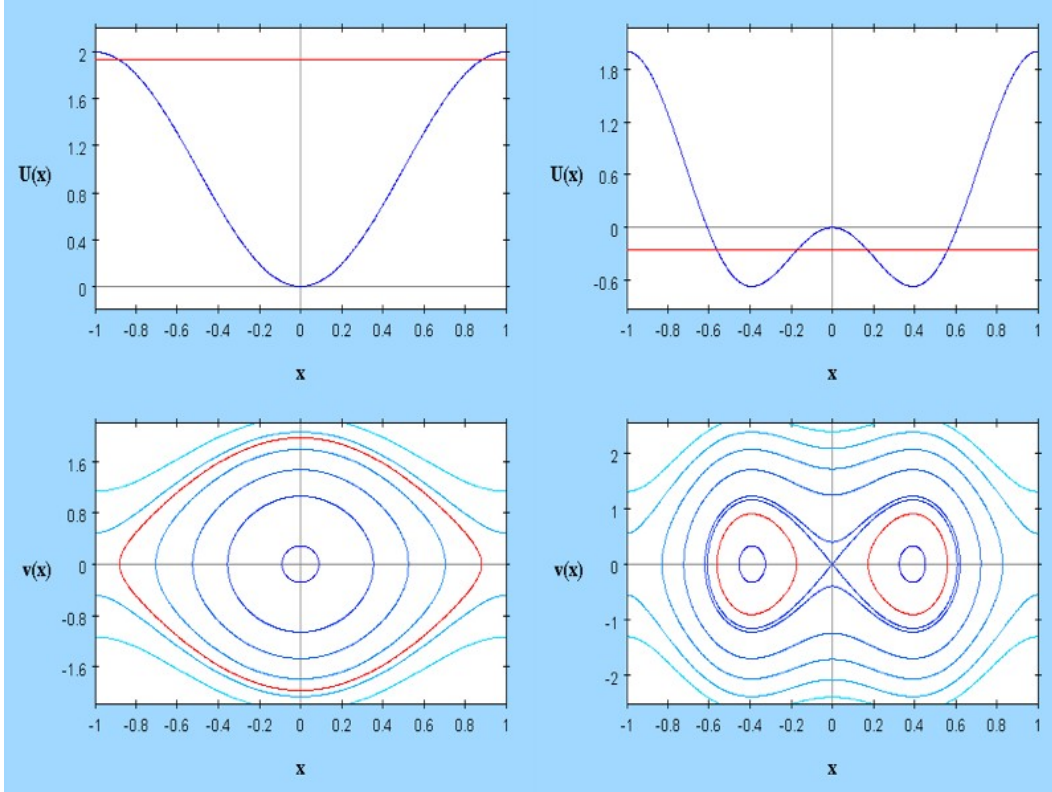


Figure 7.2: The effective potential $U_{\text{eff}}(\theta) = mga[1 - \cos \theta - \frac{\omega^2}{2\omega_0^2} \sin^2 \theta]$. (The dimensionless potential $\tilde{U}_{\text{eff}}(x) = U_{\text{eff}}/mga$ is shown, where $x = \theta/\pi$.) Left panels: $\omega = \frac{1}{2}\sqrt{3}\omega_0$. Right panels: $\omega = \sqrt{3}\omega_0$.

and thus the equations of motion are

$$ma^2\ddot{\theta} = -\frac{\partial U_{\text{eff}}}{\partial \theta} . \quad (7.64)$$

Equilibrium is achieved when $U'_{\text{eff}}(\theta) = 0$, which gives

$$\frac{\partial U_{\text{eff}}}{\partial \theta} = mga \sin \theta \left\{ 1 - \frac{\omega^2}{\omega_0^2} \cos \theta \right\} = 0 , \quad (7.65)$$

i.e. $\theta^* = 0$, $\theta^* = \pi$, or $\theta^* = \pm \cos^{-1}(\omega_0^2/\omega^2)$, where the last pair of equilibria are present only for $\omega^2 > \omega_0^2$. The stability of these equilibria is assessed by examining the sign of $U''_{\text{eff}}(\theta^*)$. We have

$$U''_{\text{eff}}(\theta) = mga \left\{ \cos \theta - \frac{\omega^2}{\omega_0^2} (2 \cos^2 \theta - 1) \right\} . \quad (7.66)$$

Thus,

$$U''_{\text{eff}}(\theta^*) = \begin{cases} mga \left(1 - \frac{\omega^2}{\omega_0^2}\right) & \text{at } \theta^* = 0 \\ -mga \left(1 + \frac{\omega^2}{\omega_0^2}\right) & \text{at } \theta^* = \pi \\ mga \left(\frac{\omega^2}{\omega_0^2} - \frac{\omega_0^2}{\omega^2}\right) & \text{at } \theta^* = \pm \cos^{-1} \left(\frac{\omega_0^2}{\omega^2}\right) . \end{cases} \quad (7.67)$$

Thus, $\theta^* = 0$ is stable for $\omega^2 < \omega_0^2$ but becomes unstable when the rotation frequency ω is sufficiently large, *i.e.* when $\omega^2 > \omega_0^2$. In this regime, there are two new equilibria, at $\theta^* = \pm \cos^{-1}(\omega_0^2/\omega^2)$, which are both stable. The equilibrium at $\theta^* = \pi$ is always unstable, independent of the value of ω . The situation is depicted in Fig. 7.2.

7.4 Charged Particle in a Magnetic Field

Consider next the case of a charged particle moving in the presence of an electromagnetic field. The particle's potential energy is

$$U(\mathbf{r}) = q\phi(\mathbf{r}, t) - \frac{q}{c} \mathbf{A}(\mathbf{r}, t) \cdot \dot{\mathbf{r}} , \quad (7.68)$$

which is velocity-dependent. The kinetic energy is $T = \frac{1}{2}m\dot{\mathbf{r}}^2$, as usual. Here $\phi(\mathbf{r})$ is the scalar potential and $\mathbf{A}(\mathbf{r})$ the vector potential. The electric and magnetic fields are given by

$$\mathbf{E} = -\nabla\phi - \frac{1}{c} \frac{\partial \mathbf{A}}{\partial t} , \quad \mathbf{B} = \nabla \times \mathbf{A} . \quad (7.69)$$

The canonical momentum is

$$\mathbf{p} = \frac{\partial L}{\partial \dot{\mathbf{r}}} = m\dot{\mathbf{r}} + \frac{q}{c} \mathbf{A} , \quad (7.70)$$

and hence the Hamiltonian is

$$\begin{aligned} H(\mathbf{r}, \mathbf{p}, t) &= \mathbf{p} \cdot \dot{\mathbf{r}} - L \\ &= m\dot{\mathbf{r}}^2 + \frac{q}{c} \mathbf{A} \cdot \dot{\mathbf{r}} - \frac{1}{2}m\dot{\mathbf{r}}^2 - \frac{q}{c} \mathbf{A} \cdot \dot{\mathbf{r}} + q\phi \\ &= \frac{1}{2}m\dot{\mathbf{r}}^2 + q\phi \\ &= \frac{1}{2m} \left(\mathbf{p} - \frac{q}{c} \mathbf{A}(\mathbf{r}, t) \right)^2 + q\phi(\mathbf{r}, t) . \end{aligned} \quad (7.71)$$

If \mathbf{A} and ϕ are time-independent, then $H(\mathbf{r}, \mathbf{p})$ is conserved.

Let's work out the equations of motion. We have

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{\mathbf{r}}} \right) = \frac{\partial L}{\partial \mathbf{r}} \quad (7.72)$$

which gives

$$m \ddot{\mathbf{r}} + \frac{q}{c} \frac{d\mathbf{A}}{dt} = -q \nabla \phi + \frac{q}{c} \nabla (\mathbf{A} \cdot \dot{\mathbf{r}}), \quad (7.73)$$

or, in component notation,

$$m \ddot{x}_i + \frac{q}{c} \frac{\partial A_i}{\partial x_j} \dot{x}_j + \frac{q}{c} \frac{\partial A_i}{\partial t} = -q \frac{\partial \phi}{\partial x_i} + \frac{q}{c} \frac{\partial A_j}{\partial x_i} \dot{x}_j, \quad (7.74)$$

which is to say

$$m \ddot{x}_i = -q \frac{\partial \phi}{\partial x_i} - \frac{q}{c} \frac{\partial A_i}{\partial t} + \frac{q}{c} \left(\frac{\partial A_j}{\partial x_i} - \frac{\partial A_i}{\partial x_j} \right) \dot{x}_j. \quad (7.75)$$

It is convenient to express the cross product in terms of the completely antisymmetric tensor of rank three, ϵ_{ijk} :

$$B_i = \epsilon_{ijk} \frac{\partial A_k}{\partial x_j}, \quad (7.76)$$

and using the result

$$\epsilon_{ijk} \epsilon_{imn} = \delta_{jm} \delta_{kn} - \delta_{jn} \delta_{km}, \quad (7.77)$$

we have $\epsilon_{ijk} B_i = \partial_j A_k - \partial_k A_j$, and

$$m \ddot{x}_i = -q \frac{\partial \phi}{\partial x_i} - \frac{q}{c} \frac{\partial A_i}{\partial t} + \frac{q}{c} \epsilon_{ijk} \dot{x}_j B_k, \quad (7.78)$$

or, in vector notation,

$$\begin{aligned} m \ddot{\mathbf{r}} &= -q \nabla \phi - \frac{q}{c} \frac{\partial \mathbf{A}}{\partial t} + \frac{q}{c} \dot{\mathbf{r}} \times (\nabla \times \mathbf{A}) \\ &= q \mathbf{E} + \frac{q}{c} \dot{\mathbf{r}} \times \mathbf{B}, \end{aligned} \quad (7.79)$$

which is, of course, the Lorentz force law.

7.5 Field Theory: Systems with Several Independent Variables

Suppose $\phi_a(\mathbf{x})$ depends on several independent variables: $\{x^1, x^2, \dots, x^n\}$. Furthermore, suppose

$$S[\{\phi_a(\mathbf{x})\}] = \int_{\Omega} d\mathbf{x} \mathcal{L}(\phi_a, \partial_{\mu} \phi_a, \mathbf{x}), \quad (7.80)$$

i.e. the Lagrangian density \mathcal{L} is a function of the fields ϕ_a and their partial derivatives $\partial\phi_a/\partial x^\mu$. Here Ω is a region in \mathbb{R}^K . Then the first variation of S is

$$\begin{aligned}\delta S &= \int_{\Omega} d\mathbf{x} \left\{ \frac{\partial \mathcal{L}}{\partial \phi_a} \delta \phi_a + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial \delta \phi_a}{\partial x^\mu} \right\} \\ &= \oint_{\partial \Omega} d\Sigma n^\mu \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \delta \phi_a - \int_{\Omega} d\mathbf{x} \left\{ \frac{\partial \mathcal{L}}{\partial \phi_a} - \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \right) \right\} \delta \phi_a ,\end{aligned}\quad (7.81)$$

where $\partial \Omega$ is the $(n - 1)$ -dimensional boundary of Ω , $d\Sigma$ is the differential surface area, and n^μ is the unit normal. If we demand $\partial \mathcal{L} / \partial (\partial_\mu \phi_a)|_{\partial \Omega} = 0$ or $\delta \phi_a|_{\partial \Omega} = 0$, the surface term vanishes, and we conclude

$$\frac{\delta S}{\delta \phi_a(\mathbf{x})} = \frac{\partial \mathcal{L}}{\partial \phi_a} - \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \right). \quad (7.82)$$

As an example, consider the case of a stretched string of linear mass density μ and tension τ . The action is a functional of the height $y(x, t)$, where the coordinate along the string, x , and time, t , are the two independent variables. The Lagrangian density is

$$\mathcal{L} = \frac{1}{2} \mu \left(\frac{\partial y}{\partial t} \right)^2 - \frac{1}{2} \tau \left(\frac{\partial y}{\partial x} \right)^2, \quad (7.83)$$

whence the Euler-Lagrange equations are

$$\begin{aligned}0 &= \frac{\delta S}{\delta y(x, t)} = -\frac{\partial}{\partial x} \left(\frac{\partial \mathcal{L}}{\partial y'} \right) - \frac{\partial}{\partial t} \left(\frac{\partial \mathcal{L}}{\partial \dot{y}} \right) \\ &= \tau \frac{\partial^2 y}{\partial x^2} - \mu \frac{\partial^2 y}{\partial t^2},\end{aligned}\quad (7.84)$$

where $y' = \frac{\partial y}{\partial x}$ and $\dot{y} = \frac{\partial y}{\partial t}$. Thus, $\mu \ddot{y} = \tau y''$, which is the Helmholtz equation. We've assumed boundary conditions where $\delta y(x_a, t) = \delta y(x_b, t) = \delta y(x, t_a) = \delta y(x, t_b) = 0$.

The Lagrangian density for an electromagnetic field with sources is

$$\mathcal{L} = -\frac{1}{16\pi} F_{\mu\nu} F^{\mu\nu} - J_\mu A^\mu. \quad (7.85)$$

The equations of motion are then

$$\frac{\partial \mathcal{L}}{\partial A^\nu} - \frac{\partial}{\partial x^\nu} \left(\frac{\partial \mathcal{L}}{\partial (\partial^\mu A^\nu)} \right) = 0 \quad \Rightarrow \quad \partial_\mu F^{\mu\nu} = 4\pi J^\nu, \quad (7.86)$$

which are Maxwell's equations.

Recall the result of Noether's theorem for mechanical systems:

$$\frac{d}{dt} \left(\frac{\partial L}{\partial \dot{q}_\sigma} \frac{\partial \tilde{q}_\sigma}{\partial \zeta} \right)_{\zeta=0} = 0 , \quad (7.87)$$

where $\tilde{q}_\sigma = \tilde{q}_\sigma(q, \zeta)$ is a one-parameter (ζ) family of transformations of the generalized coordinates which leaves L invariant. We generalize to field theory by replacing

$$q_\sigma(t) \longrightarrow \phi_a(\mathbf{x}, t) , \quad (7.88)$$

where $\{\phi_a(\mathbf{x}, t)\}$ are a set of fields, which are functions of the independent variables $\{x, y, z, t\}$. We will adopt covariant relativistic notation and write for four-vector $x^\mu = (ct, x, y, z)$. The generalization of $d\Lambda/dt = 0$ is

$$\frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \right)_{\zeta=0} = 0 , \quad (7.89)$$

where there is an implied sum on both μ and a . We can write this as $\partial_\mu J^\mu = 0$, where

$$J^\mu \equiv \frac{\partial \mathcal{L}}{\partial (\partial_\mu \phi_a)} \frac{\partial \tilde{\phi}_a}{\partial \zeta} \Big|_{\zeta=0} . \quad (7.90)$$

We call $\Lambda = J^0/c$ the *total charge*. If we assume $\mathbf{J} = 0$ at the spatial boundaries of our system, then integrating the conservation law $\partial_\mu J^\mu$ over the spatial region Ω gives

$$\frac{d\Lambda}{dt} = \int_{\Omega} d^3x \partial_0 J^0 = - \int_{\Omega} d^3x \nabla \cdot \mathbf{J} = - \oint_{\partial\Omega} d\Sigma \hat{\mathbf{n}} \cdot \mathbf{J} = 0 , \quad (7.91)$$

assuming $\mathbf{J} = 0$ at the boundary $\partial\Omega$.

As an example, consider the case of a complex scalar field, with Lagrangian density³

$$\mathcal{L}(\psi, \psi^*, \partial_\mu \psi, \partial_\mu \psi^*) = \frac{1}{2} K (\partial_\mu \psi^*)(\partial^\mu \psi) - U(\psi^* \psi) . \quad (7.92)$$

This is invariant under the transformation $\psi \rightarrow e^{i\zeta} \psi$, $\psi^* \rightarrow e^{-i\zeta} \psi^*$. Thus,

$$\frac{\partial \tilde{\psi}}{\partial \zeta} = i e^{i\zeta} \psi \quad , \quad \frac{\partial \tilde{\psi}^*}{\partial \zeta} = -i e^{-i\zeta} \psi^* , \quad (7.93)$$

and, summing over both ψ and ψ^* fields, we have

$$\begin{aligned} J^\mu &= \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi)} \cdot (i\psi) + \frac{\partial \mathcal{L}}{\partial (\partial_\mu \psi^*)} \cdot (-i\psi^*) \\ &= \frac{K}{2i} (\psi^* \partial^\mu \psi - \psi \partial^\mu \psi^*) . \end{aligned} \quad (7.94)$$

The potential, which depends on $|\psi|^2$, is independent of ζ . Hence, this form of conserved 4-current is valid for an entire class of potentials.

³We raise and lower indices using the Minkowski metric $g_{\mu\nu} = \text{diag}(+, -, -, -)$.

7.5.1 Gross-Pitaevskii Model

As one final example of a field theory, consider the Gross-Pitaevskii model, with

$$\mathcal{L} = i\hbar \psi^* \frac{\partial \psi}{\partial t} - \frac{\hbar^2}{2m} \nabla \psi^* \cdot \nabla \psi - g (|\psi|^2 - n_0)^2. \quad (7.95)$$

This describes a Bose fluid with repulsive short-ranged interactions. Here $\psi(\mathbf{x}, t)$ is again a complex scalar field, and ψ^* is its complex conjugate. Using the Leibniz rule, we have

$$\begin{aligned} \delta S[\psi^*, \psi] &= S[\psi^* + \delta\psi^*, \psi + \delta\psi] \\ &= \int dt \int d^d x \left\{ i\hbar \psi^* \frac{\partial \delta\psi}{\partial t} + i\hbar \delta\psi^* \frac{\partial \psi}{\partial t} - \frac{\hbar^2}{2m} \nabla \psi^* \cdot \nabla \delta\psi - \frac{\hbar^2}{2m} \nabla \delta\psi^* \cdot \nabla \psi \right. \\ &\quad \left. - 2g (|\psi|^2 - n_0) (\psi^* \delta\psi + \psi \delta\psi^*) \right\} \\ &= \int dt \int d^d x \left\{ \left[-i\hbar \frac{\partial \psi^*}{\partial t} + \frac{\hbar^2}{2m} \nabla^2 \psi^* - 2g (|\psi|^2 - n_0) \psi^* \right] \delta\psi \right. \\ &\quad \left. + \left[i\hbar \frac{\partial \psi}{\partial t} + \frac{\hbar^2}{2m} \nabla^2 \psi - 2g (|\psi|^2 - n_0) \psi \right] \delta\psi^* \right\}, \quad (7.96) \end{aligned}$$

where we have integrated by parts where necessary and discarded the boundary terms. Extremizing $S[\psi^*, \psi]$ therefore results in the *nonlinear Schrödinger equation* (NLSE),

$$i\hbar \frac{\partial \psi}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi + 2g (|\psi|^2 - n_0) \psi \quad (7.97)$$

as well as its complex conjugate,

$$-i\hbar \frac{\partial \psi^*}{\partial t} = -\frac{\hbar^2}{2m} \nabla^2 \psi^* + 2g (|\psi|^2 - n_0) \psi^*. \quad (7.98)$$

Note that these equations are indeed the Euler-Lagrange equations:

$$\frac{\delta S}{\delta \psi} = \frac{\partial \mathcal{L}}{\partial \psi} - \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial \partial_\mu \psi} \right) \quad (7.99)$$

$$\frac{\delta S}{\delta \psi^*} = \frac{\partial \mathcal{L}}{\partial \psi^*} - \frac{\partial}{\partial x^\mu} \left(\frac{\partial \mathcal{L}}{\partial \partial_\mu \psi^*} \right), \quad (7.100)$$

with $x^\mu = (t, \mathbf{x})$ ⁴ Plugging in

$$\frac{\partial \mathcal{L}}{\partial \psi} = -2g (|\psi|^2 - n_0) \psi^*, \quad \frac{\partial \mathcal{L}}{\partial \partial_t \psi} = i\hbar \psi^*, \quad \frac{\partial \mathcal{L}}{\partial \nabla \psi} = -\frac{\hbar^2}{2m} \nabla \psi^* \quad (7.101)$$

⁴In the nonrelativistic case, there is no utility in defining $x^0 = ct$, so we simply define $x^0 = t$.

and

$$\frac{\partial \mathcal{L}}{\partial \psi^*} = i\hbar \psi - 2g(|\psi|^2 - n_0)\psi \quad , \quad \frac{\partial \mathcal{L}}{\partial \partial_t \psi^*} = 0 \quad , \quad \frac{\partial \mathcal{L}}{\partial \nabla \psi^*} = -\frac{\hbar^2}{2m} \nabla \psi \quad , \quad (7.102)$$

we recover the NLSE and its conjugate.

The Gross-Pitaevskii model also possesses a U(1) invariance, under

$$\psi(\mathbf{x}, t) \rightarrow \tilde{\psi}(\mathbf{x}, t) = e^{i\zeta} \psi(\mathbf{x}, t) \quad , \quad \psi^*(\mathbf{x}, t) \rightarrow \tilde{\psi}^*(\mathbf{x}, t) = e^{-i\zeta} \psi^*(\mathbf{x}, t) \quad . \quad (7.103)$$

Thus, the conserved Noether current is then

$$J^\mu = \left. \frac{\partial \mathcal{L}}{\partial \partial_\mu \psi} \frac{\partial \tilde{\psi}}{\partial \zeta} \right|_{\zeta=0} + \left. \frac{\partial \mathcal{L}}{\partial \partial_\mu \psi^*} \frac{\partial \tilde{\psi}^*}{\partial \zeta} \right|_{\zeta=0}$$

$$J^0 = -\hbar |\psi|^2 \quad (7.104)$$

$$\mathbf{J} = -\frac{\hbar^2}{2im} (\psi^* \nabla \psi - \psi \nabla \psi^*) \quad . \quad (7.105)$$

Dividing out by \hbar , taking $J^0 \equiv -\hbar \rho$ and $\mathbf{J} \equiv -\hbar \mathbf{j}$, we obtain the continuity equation,

$$\frac{\partial \rho}{\partial t} + \nabla \cdot \mathbf{j} = 0 \quad , \quad (7.106)$$

where

$$\rho = |\psi|^2 \quad , \quad \mathbf{j} = \frac{\hbar}{2im} (\psi^* \nabla \psi - \psi \nabla \psi^*) \quad . \quad (7.107)$$

are the particle density and the particle current, respectively.