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Pure Gravity

or

Particles at Play

Volume 1:

Writing an N-Body Code

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Chapter 8

A More Modular N-Body Hermite Code

8.1 Starting a Tool Box

In this chapter we will discuss in detail a more modular version of the Hermite code hermite6.C, developed in the previous chapter. The new version is called nbody_sh1.C. Here 'sh' stands for the shared but variable time step choice, and the number 1 indicates again that this is the first version. This code will be the first tool of a tool box that we will continue to develop in the rest of this book, as well as in following books in this series. From now on, each tool will adhere to our N-body I/O format, specified in the previous chapter (and possibly more fancy formats as well, but we will keep those more advanced versions compatible with our current bare bones format). In addition, each tool will have extensive comments, explaining both the usage and the internal structure of the code.

The new code, nbody_sh1.C, has roughly four times more lines than the previous version, hermite6.C. Almost half of these lines are either comments of blank lines, both of which help to make the code more readable and more understandable. The fact that the code itself still has more than twice the length of the previous version stems from several factors. First, the new code has nine functions, besides main(), while the old code had only two. Second, there are seven command line options, rather than two. Third, we now declare all functions at the top of the file. Finally, there is more diagnostics output than we had before.

Below, the full code is presented, one function at a time.

8.2 Gravitylab

Our aim is to build a powerful software environment for experiments in stellar dynamics of dense stellar systems. The idea is to build a virtual laboratory, which we will call *gravitylab*. From now on, each new tool in our tool box will have a distinctive 'gravitylab' header:

We will not show these headers in future code listings, but they will be there in the source code for other tools. The three stars moving on a figure-8 orbit are inspired by the solution presented in chapter 5. They are being observed at bottom left by the small figure looking through a telescope.

Note the time stamp at the very first line. This is a handy feature of the *emacs* editor that we have used to write this book. When you add the line "(add-hook 'write-file-hooks 'time-stamp)" to the .emacs startup file, the date and time and user name will be updated automatically each time you write the file to disk.

8.3 Introductory Comments

Immediately following the gravitylab header, we see a lengthy comment block:

```
(the same for all particles but changing in time), using
                 the Hermite integration scheme.
                 ref.: Hut, P., Makino, J. & McMillan, S., 1995,
                       Astrophysical Journal Letters 443, L93-L96.
 note: in this first version, all functions are included in one file,
        without any use of a special library or header files.
 usage: nbody_sh1 [-h (for help)] [-d step_size_control_parameter]
                    [-e diagnostics_interval] [-o output_interval]
                    [-t total_duration] [-i (start output at t = 0)]
                    [-x (extra debugging diagnostics)]
          "step_size_control_parameter" is a coefficient determining the
            the size of the shared but variable time step for all particles
          "diagnostics_interval" is the time between output of diagnostics,
             in the form of kinetic, potential, and total energy; with the
            -x option, a dump of the internal particle data is made as well
          "output_interval" is the time between successive snapshot outputs
          "total_duration" is the integration time, until the program stops
         Input and output are written from the standard i/o streams. Since
         all options have sensible defaults, the simplest way to run the code
          is by only specifying the i/o files for the N-body snapshots:
            nbody_sh1 < data.in > data.out
         The diagnostics information will then appear on the screen.
         To capture the diagnostics information in a file, capture the
          standard error stream as follows:
             (nbody_sh1 < data.in > data.out) >& data.err
* Note: if any of the times specified in the -e, -o, or -t options are not an
        an integer multiple of "step", output will occur slightly later than
        predicted, after a full time step has been taken. And even if they
        are integer multiples, round-off error may induce one extra step.
* External data format:
     The program expects input of a single snapshot of an N-body system,
```

It starts with the name of the file, a brief summary with a reference to the literature, followed by a detailed description of how to use the code. For a typical user, this is all the information needed. As long as the user combines <code>nbody_sh1</code> with other tools from gravitylab, there is even no need to understand the external data format, in which the N-body snapshots are written to and read from files. For those users interested in such details, as well as in the internal format in which the data are stored during the execution of the code, the comment block contains format information near the end. The last few lines list the history and version numbers of the code.

8.4 Include Statements, Function Declarations, etc.

The first lines of real code start right after the introductory comments:

```
// standard floating-point data type
const int NDIM = 3;
                                           // number of spatial dimensions
void correct_step(real pos[][NDIM], real vel[][NDIM],
                  const real acc[][NDIM], const real jerk[][NDIM],
                  const real old_pos[][NDIM], const real old_vel[][NDIM],
                  const real old_acc[][NDIM], const real old_jerk[][NDIM],
                  int n, real dt);
void evolve(const real mass[], real pos[][NDIM], real vel[][NDIM],
            int n, real & t, real dt_param, real dt_dia, real dt_out,
            real dt_tot, bool init_out, bool x_flag);
void evolve_step(const real mass[], real pos[][NDIM], real vel[][NDIM],
                 real acc[][NDIM], real jerk[][NDIM], int n, real & t,
                 real dt, real & epot, real & coll_time);
void get_acc_jerk_pot_coll(const real mass[], const real pos[][NDIM],
                           const real vel[][NDIM], real acc[][NDIM],
                           real jerk[][NDIM], int n, real & epot,
                           real & coll_time);
void get_snapshot(real mass[], real pos[][NDIM], real vel[][NDIM], int n);
void predict_step(real pos[][NDIM], real vel[][NDIM],
                  const real acc[][NDIM], const real jerk[][NDIM],
                  int n, real dt);
void put_snapshot(const real mass[], const real pos[][NDIM],
                  const real vel[][NDIM], int n, real t);
bool read_options(int argc, char *argv[], real & dt_param, real & dt_dia,
                  real & dt_out, real & dt_tot, bool & i_flag, bool & x_flag);
void write_diagnostics(const real mass[], const real pos[][NDIM],
                       const real vel[][NDIM], const real acc[][NDIM],
                       const real jerk[][NDIM], int n, real t, real epot,
                       int nsteps, real & einit, bool init_flag,
                       bool x_flag);
```

We start with #include statements to various libraries. The comments on each line mention some of the functions used from those libraries. If we would leave out one of these include statements, the corresponding functions listed could not be linked, and the compiler would issue an error.

The next statement indicates that we used the standard C++ namespace. Later, when gravitylab will have grown sufficiently large, it may be useful to create our own namespaces, in order to avoid collisions with other programs that may use names that are the same as we have chosen. Right now it is too early to worry about such complications.

The typedef statement defines the word real as an alternative for the build-in function type double. From now on we will only use the name real to indicate the

standard floating point type double. It is far more logical to talk about real numbers of type real, together with the integers of type int, without using the archaic term 'double' that stems from the expression 'double precision' (long ago, the standard precision for floating point calculations used four bytes per floating point word, leading to the expression double precision for the now-standard eight-byte word length).

Next we introduce the symbol NDIM for the number of dimensions. So far we have simply used the number 3 in our loops over Cartesian coordinates, but it is much better not to have any magic numbers in a code, where a magic number is defined as anything that is not 0 or 1. The term "NDIM" for the number of dimensions is far clearer than a blind "3" in the middle of a piece of code. A second advantage of introducing a symbol, rather than magic numbers, is that we can change the symbol at one place, while guaranteeing its substitution everywhere else in the code. In the vast majority of cases, we will do our simulations in three spatial dimensions, hence the assignment here of the number 3 to NDIM here, but we will also encounter cases where we want to do some experimentation in one or two dimensions. In that case, changing 3 to 1 or 2 in this line is all we need to do (apart from making sure that we have not used uniquely three-dimensional constructs elsewhere in the code, such as for example the use of 3D spherical harmonics).

Note that older C-style usage would have defined NDIM through the macro definition "#define NDIM 3". Nowadays, however, it is considered good form to use the C++ expression "const int NDIM = 3;". Although the use of a #define macro in this case is quite innocent, there are many other cases where the use of macros can lead to code that is prone to confusing errors that are hard to debug. Therefore, as a matter of style it is a good idea to avoid them as much as possible.

The following nine function declarations are necessary if we want to have the freedom to define them in an arbitrary way in the rest of the file. The problem is that the C compiler goes through the file in one single pass, from top to bottom. As long as each function is invoked only after it has been seen by the compiler, there is no problem. In the codes hermite4.C through hermite6.C, the two functions listed at the top of the files were invoked only by main(), which was listed last. In general, however, with many functions there may not be a unique flow of functions calls. Besides, it is easier to follow the logic of the code if we can start with main() at the top of the file. The latter immediately implies that we will have to declare all functions mentioned in main().

This need for redundant information in the form of declarations is a weakness of C++. In general, any time that a computer language forces you to duplicate information, it brings with it the danger of errors creeping in. It is easy to change the definition of a function without changing the declaration, or *vice versa*. In some cases, the compiler may catch this, but there may be other cases where overloading of function names with different argument sets makes it impossible for the compiler to catch such mistakes. Unfortunately, we will have to live with this situation.

Another example of redundant information in our program is the description of the command line options. Almost the same words appear once in the 'usage' part of the initial commments, and twice in the function read_options() (for the help option and the unknown option). It is possible to capture that information in a string at the top of the program, and to echo that string in read_options(). We will make such a modification later.

8.5 The Function main()

```
/*-----
* main -- reads option values, reads a snapshot, and launches the
            integrator
*-----
*/
int main(int argc, char *argv[])
   real dt_param = 0.03;
                           // control parameter to determine time step size
   real dt_dia = 1;
                           // time interval between diagnostics output
   real dt_out = 1;
                           // time interval between output of snapshots
   real dt_tot = 10;
                           // duration of the integration
   bool init_out = false;
                           // if true: snapshot output with start at t = 0
                                      with an echo of the input snapshot
                           //
   bool x_flag = false;
                           // if true: extra debugging diagnostics output
   if (! read_options(argc, argv, dt_param, dt_dia, dt_out, dt_tot, init_out,
                    x_flag))
                             // halt criterion detected by read_options()
       return 1:
                             // N, number of particles in the N-body system
   int n;
   cin >> n;
   real t;
                             // time
   cin >> t;
   real * mass = new real[n];
                                          // masses for all particles
   real (* pos)[NDIM] = new real[n][NDIM];
                                          // positions for all particles
   real (* vel)[NDIM] = new real[n][NDIM];
                                         // velocities for all particles
   get_snapshot(mass, pos, vel, n);
   evolve(mass, pos, vel, n, t, dt_param, dt_dia, dt_out, dt_tot, init_out,
         x_flag);
   delete[] mass;
```

```
delete[] pos;
  delete[] vel;
}
```

The first six variables declared at the top of main() receive their values from the function read_options() which reads the Unix style command line arguments. Note that each variable has a default value, which is retained unless it is changed explicitly by the corresponding command. We discuss the usage of command line options in the next section.

If the function read_options() detects a request for help, or the invocation of a non-existent option, it will return the Boolean value false. In that case the statement !read_options() is true, and program execution is halted. In C++, returning the value 0 indicates normal successful completion of the main() program, while any other value indicates a failure of some kind or other. For simplicity we return here the value 1.

Once the options are interpreted, we are ready to read the N-body snapshot from the standard input (which typically is redirected to read either the contents of a file, as in nbody_sh1 < data.in or to receive data from another program through a pipe, as in generate_data | nbody_sh1). Once the number of particles n has been read in, we can allocate storage space to contain the masses and dynamical information for all n particles, as we have seen in the previous chapter. The actual initialization of the arrays is carried out by the function get_snapshot().

The real work is then delegated to the function evolve(), which oversees the evolution in time of the N-body system. When the call to evolve() returns, there is nothing left to be done. For good form we then deallocate the memory that we had dynamically allocated with the new operator. Note the square brackets in delete, which tell the compiler to delete the full memory assigned to the arrays. If we would leave this out, for example in a statement delete[] mass, we would only free the memory for mass[0]. This would constitute a memory leak, since the rest of the array will still be allocated, but it will be no longer usable in our program. In our particular case, this is no problem since we are about to terminate the program anyway, but in more complex cases, such as we will encounter in the function evolve(), it will be important to not create memory leaks.

8.6 Command Line Options

There are six command line options, Unix style, from which we can choose. All essential options have default values, so it is perfectly possible to run our code without specifying any of them. For example, if we start with an N-body snapshot in an input file data.in, we can run the code to produce a stream of snapshot data in the output file data.out, by typing:

```
|gravity> nbody_sh1 < data.in > data.out
```

This will have the exact same effect as if we would have specified the default values for the four main options, namely the time step control parameter (0.03), the interval between diagnostics output (1 time unit), the interval between output of snapshots (1 time unit), and the duration of the integration (10 time units):

```
|gravity> nbody_sh1 -d 0.03 -e 1 -o 1 -t 10 < data.in > data.out
```

If we would like to have three times smaller time steps, twice as many diagnostics outputs and with additional information, snapshot output intervals of 5 time units but starting at t=0, and a total run time of 30 time units, we have to give the following command:

```
|gravity> nbody_sh1 -d 0.01 -e 0.5 -x -o 5 -i -t 30 < data.in > data.out
```

The order of the arguments is unimportant, but each option that expects a value (the -d, -e, -o, -t options) should be immediately followed by its corresponding value. By the way, the value 0.03 as the default for the scale of the time step parameter is somewhat arbitrary. In practice, a value of 0.1 is often found to be too large, while 0.01 is often overkill. For example, when we start from the initial conditions for three stars on a figure 8 orbit, running $nbody_sh1$ with all default values in place, we wind up at time t = 10 with a relative energy error of order 10^{-7} .

Of course, the optimal choice of values depend strongly on the particular application, and the default values are only a hint, in a blind attempt to come up with at least somewhat reasonable starting values. It is up to the user to make sure that these values are appropriate in a given situation, and if not, to supply a better value after some experimentation.

The help option can be invoked by typing:

```
|gravity> nbody_sh1 -h
```

This will not result in program execution, only in the printing of a short message that lays out the various command line option choices. A similar message will appear when we attempt to supply an non-existent option, for example:

All this behavior can be inspected in the function read_options():

```
read_options -- reads the command line options, and implements them.
   note: when the help option -h is invoked, the return value is set to false,
          to prevent further execution of the main program; similarly, if an
          unknown option is used, the return value is set to false.
*/
bool read_options(int argc, char *argv[], real & dt_param, real & dt_dia,
                  real & dt_out, real & dt_tot, bool & i_flag, bool & x_flag)
    int c;
   while ((c = getopt(argc, argv, "hd:e:o:t:ix")) != -1)
        switch(c){
            case 'h': cerr << "usage: " << argv[0]</pre>
                           << " [-h (for help)]"
                           << " [-d step_size_control_parameter]\n"</pre>
                           << " [-e diagnostics_interval]"</pre>
                           << " [-o output_interval]\n"
                           << "
                                        [-t total_duration]"
                           << " [-i (start output at t = 0)]\n"
                           << "
                                       [-x (extra debugging diagnostics)]"
                           << endl;
                      return false;
                                            // execution should stop after help
            case 'd': dt_param = atof(optarg);
                      break;
            case 'e': dt_dia = atof(optarg);
                      break;
            case 'i': i_flag = true;
                      break;
            case 'o': dt_out = atof(optarg);
                      break:
            case 't': dt_tot = atof(optarg);
                      break;
            case 'x': x_flag = true;
                      break;
            case '?': cerr << "usage: " << argv[0]</pre>
                           << " [-h (for help)]"
                           << " [-d step_size_control_parameter]\n"</pre>
                           << "
                                       [-e diagnostics_interval]"
                           << " [-o output_interval] \n"
                           << "
                                        [-t total_duration]"
                           << " [-i (start output at t = 0)]\n"
                           << "
                                        [-x (extra debugging diagnostics)]"
```

Note that the six variables corresponding to the command line arguments are all passed by reference, so that the results are available to the calling program main().

The function getopt() is a standard C library function that can be used equally well in C++ programs. Its third argument is a string which lists all command line options. Each option can only consist of a single letter. Those letters that should be followed by a value to be read in are indicated by a colon immediately following the letter. The string "hd:e:o:t:ix" tells us that options h, i and x do not expect additional values, while options d, e, o and t are to be followed with an argument, all of which are of type real in our particle case. All option arguments are by default passed as ASCII strings, so we need the function atof() to convert the ASCII information into the proper floating point value, as we already saw in the previous chapter.

Notice that each case in the body of the switch statement is ended by either a return statement or a break statement. The latter is necessary, since the default behavior of switch is to 'fall through' from one case to the next, something that is clearly not desirable here. After we jump out of the switch statement through a break command, we encounter the last statement, "return true;" which tells the calling program that all is well, and that execution can continue.

8.7 Snapshot Input

The code for snapshot input is straightforward:

Note that we do not check here whether a complete snapshot is being offered on the standard input stream in the right format. It would be better to verify, for example, that new lines

n occur in the correct places, separating each particle, and that no end-of-file condition is encountered before the whole N-body snapshot is read in. In later versions we will provide more complete error checking.

8.8 Snapshot Output

The code for snapshot output is similarly simple:

```
/*-----
   put_snapshot -- writes a single snapshot on the output stream cout.
   note: unlike get_snapshot(), put_snapshot handles particle number and time
 */
void put_snapshot(const real mass[], const real pos[][NDIM],
                 const real vel[][NDIM], int n, real t)
{
   cout.precision(16);
                                            // full double precision
   cout << n << endl;</pre>
                                            // N, total particle number
   cout << t << endl;</pre>
                                            // current time
   for (int i = 0; i < n; i++){
       cout << mass[i];</pre>
                                            // mass of particle i
       for (int k = 0; k < NDIM; k++)
           cout << ' ' << pos[i][k];</pre>
                                            // position of particle i
       for (int k = 0; k < NDIM; k++)
           cout << ' ' << vel[i][k];</pre>
                                            // velocity of particle i
       cout << endl;</pre>
   }
}
```

Note that the masses, positions, and velocities are all declared as const in the declaration of the function arguments. This means that this function is not allowed

to change the values of those particular arguments. Being able to specify function arguments as const is a very useful C++ feature. It can help the compiler by providing extra information; it allows the compiler to flag an error if in the body of the function an attempt is made to change one of those arguments erroneously; and most importantly, it gives the human reader useful information about the intentions of the programmer.

For all these reasons, it is important to be consistent in the use of const specifications, and to always use const wherever we can. When we do this, we thereby imply that the absence of a const specifier for an argument means that we do want to affect the value of that particular argument. For example, in the previous function get_snapshot(), masses, positions, and velocities are not preceded by const. Indeed, all three arrays are being initialized in that function, and it is useful to be able to anticipate that already from looking at the argument list, either here or at the top of the file where all functions are declared.

The first line of the body of the function sets the precision for all subsequent output. It turns out that eight-byte double precision corresponds to about 16 digits of relative accuracy. If we would output less than 16 significant digits for each real variable, we would lose information. A subsequent program reading in the snapshot that we have just written out would not have access to the full information that we had before we wrote our data. On the other hand, if we would output those numbers with more than 16 digits, the extra digits would be effective garbage. While this doesn't hurt, it is a waste of space (and possibly later processing time) to go beyond 16 digits.

8.9 Reporting Diagnostics

Here is the code for the function that writes diagnostics to the standard error stream. Note the declarations of arguments: all arrays are specified to be <code>const</code>, which is appropriate since their values should only be reported, without changing them. The argument <code>einit</code> is passed by reference, since it will hold the initial value of the total energy of the system, information that should be passed back to the calling function. The other arguments are all passed by value.

```
/*----
* write_diagnostics -- writes diagnostics on the error stream cerr:

* current time; number of integration steps so far;

* kinetic, potential, and total energy; absolute and

* relative energy errors since the start of the run.

* If x_flag (x for eXtra data) is true, all internal

* data are dumped for each particle (mass, position,

* velocity, acceleration, and jerk).

*

* note: the kinetic energy is calculated here, while the potential energy is

* calculated in the function get_acc_jerk_pot_coll().
```

```
*-----
void write_diagnostics(const real mass[], const real pos[][NDIM],
                      const real vel[][NDIM], const real acc[][NDIM],
                      const real jerk[][NDIM], int n, real t, real epot,
                      int nsteps, real & einit, bool init_flag,
                      bool x_flag)
{
   real ekin = 0;
                                        // kinetic energy of the n-body system
   for (int i = 0; i < n; i++)
        for (int k = 0; k < NDIM; k++)
            ekin += 0.5 * mass[i] * vel[i][k] * vel[i][k];
   real etot = ekin + epot;
                                        // total energy of the n-body system
    if (init_flag)
                                        // at first pass, pass the initial
        einit = etot;
                                        // energy back to the calling function
    cerr << "at time t = " << t << " , after " << nsteps
        << " steps :\n E_kin = " << ekin
        << " , E_pot = " << epot
        << " , E_tot = " << etot << endl;
    cerr << "
        << "absolute energy error: E_tot - E_init = "</pre>
        << etot - einit << endl;
    cerr << "
        << "relative energy error: (E_tot - E_init) / E_init = "</pre>
        << (etot - einit) / einit << endl;
    if (x_flag){
        cerr << " for debugging purposes, here is the internal data "
            << "representation:\n";</pre>
        for (int i = 0; i < n; i++){
           cerr << "
                      internal data for particle " << i+1 << " : " << endl;
                        ";
           cerr << "
           cerr << mass[i];</pre>
           for (int k = 0; k < NDIM; k++)
               cerr << ', ' << pos[i][k];
           for (int k = 0; k < NDIM; k++)
               cerr << ' ' << vel[i][k];</pre>
           for (int k = 0; k < NDIM; k++)
               cerr << ' ' << acc[i][k];</pre>
           for (int k = 0; k < NDIM; k++)
               cerr << ' ' << jerk[i][k];</pre>
           cerr << endl;</pre>
        }
```

```
}
```

The only calculation performed in this function is that of the kinetic energy. The potential energy is determined in the function <code>get_acc_jerk_pot_coll()</code>. The <code>init_flag</code> is set to <code>true</code> when <code>write_diagnostics()</code> is evoked for the first time, at <code>t = 0</code>. In that case, we want to pass the value of the initial total energy back to the calling function <code>evolve()</code>, which can use that information to compare it with later measured values of the total energy, in order to determine the absolute and relative amounts of energy drifts, which are a good measure of numerical accuracy.

Note that we could have defined the initial energy einit as a static variable inside write_diagnostics(). For our present purpose that would be fine, but this type of programming may easily create a future limitation. If some day we would like to compare two different N-body systems, each of which evolves, we would get into a conflict if both of them would try to access the same static variable. Therefore, for the same reason we don't use global variables in the first place, we prefer to pass einit as a function variable.

8.10 Orbit Integration

We now come to the function that manages the orbit evolution, driving the Hermite integrator and scheduling the various output operations:

```
evolve -- integrates an N-body system, for a total duration dt_tot.
               Snapshots are sent to the standard output stream once every
               time interval dt_out. Diagnostics are sent to the standard
              error stream once every time interval dt_dia.
  note: the integration time step, shared by all particles at any given time,
         is variable. Before each integration step we use coll_time (short
         for collision time, an estimate of the time scale for any significant
         change in configuration to happen), multiplying it by dt_param (the
         accuracy parameter governing the size of dt in units of coll_time),
         to obtain the new time step size.
* Before moving any particles, we start with an initial diagnostics output
  and snapshot output if desired. In order to write the diagnostics, we
* first have to calculate the potential energy, with get_acc_jerk_pot_coll().
* That function also calculates accelerations, jerks, and an estimate for the
* collision time scale, all of which are needed before we can enter the main
  integration loop below.
```

```
In the main loop, we take as many integration time steps as needed to
* reach the next output time, do the output required, and continue taking
st integration steps and invoking output this way until the final time is
 * reached, which triggers a 'break' to jump out of the infinite loop set up
* with 'while(true)'.
*/
void evolve(const real mass[], real pos[][NDIM], real vel[][NDIM],
           int n, real & t, real dt_param, real dt_dia, real dt_out,
           real dt_tot, bool init_out, bool x_flag)
{
    cerr << "Starting a Hermite integration for a " << n
        << "-body system,\n from time t = " << t
        << " with time step control parameter dt_param = " << dt_param
        << " until time " << t + dt_tot
        << " ,\n with diagnostics output interval dt_dia = "
         << dt_dia << ",\n and snapshot output interval dt_out = "
        << dt_out << "." << endl;
   real (* acc)[NDIM] = new real[n][NDIM];
                                                    // accelerations and jerks
   real (* jerk)[NDIM] = new real[n][NDIM];
                                                    // for all particles
   real epot;
                                     // potential energy of the n-body system
   real coll_time;
                                     // collision (close encounter) time scale
   get_acc_jerk_pot_coll(mass, pos, vel, acc, jerk, n, epot, coll_time);
   int nsteps = 0;
                                 // number of integration time steps completed
                                 // initial total energy of the system
   real einit;
   write_diagnostics(mass, pos, vel, acc, jerk, n, t, epot, nsteps, einit,
                     true, x_flag);
   if (init_out)
                                                    // flag for initial output
       put_snapshot(mass, pos, vel, n, t);
   real t_dia = t + dt_dia;
                                     // next time for diagnostics output
   real t_out = t + dt_out;
                                     // next time for snapshot output
   real t_end = t + dt_tot;
                                     // final time, to finish the integration
   while (true){
       while (t < t_dia \&\& t < t_out \&\& t < t_end){
           real dt = dt_param * coll_time;
           evolve_step(mass, pos, vel, acc, jerk, n, t, dt, epot, coll_time);
           nsteps++;
       }
        if (t \ge t_dia)
           write_diagnostics(mass, pos, vel, acc, jerk, n, t, epot, nsteps,
```

```
einit, false, x_flag);
    t_dia += dt_dia;
}
if (t >= t_out){
    put_snapshot(mass, pos, vel, n, t);
    t_out += dt_out;
}
if (t >= t_end)
    break;
}
delete[] acc;
delete[] jerk;
}
```

Starting again with the argument list, we see that the mass array, as always, is defined as const, since we do not model a mechanism for mass loss for stars, nor do we (yet) allow collisions between stars, which could be followed by mergers that would produce a merger remnant with a mass equal to the sum of the masses of the two stars. The only place where we do not define the mass array as const is in the function get_snapshot, where the mass values are read in from the standard input stream. Note that the time t is passed by reference. In our current program, this is not necessary, since the value of t is not used in main(), where execution is halted immediately upon completion of the call to evolve(). However, in future extensions we may well add further commands in main(), and in that case it would be useful to have the value of the current time available.

As we have seen before, before we can enter the integration loop we have to start with an initial call to the function computing the accelerations and jerks. This function, get_acc_jerk_pot_coll() does what its name suggest: besides calculating accelerations and jerks, it also reports the value of the total potential energy of the system as well as the value of the time scale on which a 'collision' between particles can occur, i. e. a significant change of order unity in the local configuration of at least two particles. The latter information, stored in the variable coll_time, will be needed in the main integration loop in order to determine the size of the first time step. Accelerations and jerks are needed for the first part of the first integration time step, and the potential energy is used in the initial call to write_diagnostics(), following the first call to get_acc_jerk_pot_coll().

In addition, if the user has specified the init_out flag to be true, the input values of the N-body system are echoed as they are on the output stream; the default behavior is to wait with output until some integration steps have been taken. This is a sensible default, since in many cases we are only interested in one final output snapshot, which can then served as the input for a later invocation of the integrator. If we invoke our program with the same value for the snapshot output interval as the duration of the

run, we guarantee that only one final output will be made. An example usage of this type is:

```
|gravity> nbody_sh1 -d 0.01 -e 2 -o 40 -t 40 < data.in > data.out
```

Before entering the main integration loop, we schedule the next times for diagnostics and snapshot output, as well as the final halting time. The loop itself is an infinite loop, governed by the tautological while (true), which is obviously always the case. The standard C/C++ trick to define an infinite loop uses an empty for loop, in the form for(;;), but that expression is less transparent, whereas while (true) leaves no doubt as to it being an infinite loop. The only way to jump out of this infinite loop is at the end of the loop: when time progresses past the halting time t_end, the break statement causes control flow to continue past the loop.

The first time we enter the loop, the second while argument will be evaluated as true, unless one of the three values dt_dia, dt_out or dt_tot would be zero or negative, which would be nonsensical values. Ideally, we should check somewhere that all command line option arguments fall within reasonable ranges. Since in the present code we have already introduced so many new features, we will not include such a defensive programming style at this point. However, later on we will insist on checking all values which reach a program through an interface, such as presented by command line options. For now, we will live with the danger of a non-positive value for either dt_dia or dt_out, which combined with a positive value for dt_tot would lead to an infinite number of output attempts, without the time ever advancing.

With natural choices of parameters, the majority of loop cycles will not lead to any output. In those cases a new time step size is determined, and the function evolve_step() is called, which as the name implies will advance the system by one integration step, and in addition update the time by an amount dt. Sooner or later it will be time for output or for ending the run. In either case, the second while statement will evaluate as false, no integration time step will be taken and therefore the time will not be advanced either. Instead, the required output will be done and/or the integration will be finished altogether. If the run is not yet finished, the next cycle in the infinite loop will lead to another integration step, and so on.

Note the freeing up of memory for acceleration and jerk arrays, at the end of evolve(). As in the case of the memory allocation in main(), this is not strictly necessary, since the program is about to finish, but again it is certainly good form to include these statements here.

8.11 Taking a Single Integration Step

In the function evolve_step(), we encounter the first case where specific memory allocation and deallocation occurs more often than once during a run:

```
/*-----
 * evolve_step -- takes one integration step for an N-body system, using the
                  Hermite algorithm.
 */
void evolve_step(const real mass[], real pos[][NDIM], real vel[][NDIM],
               real acc[][NDIM], real jerk[][NDIM], int n, real & t,
               real dt, real & epot, real & coll_time)
{
   real (* old_pos)[NDIM] = new real[n][NDIM];
   real (* old_vel)[NDIM] = new real[n][NDIM];
   real (* old_acc)[NDIM] = new real[n][NDIM];
   real (* old_jerk)[NDIM] = new real[n][NDIM];
   for (int i = 0; i < n; i++)
       for (int k = 0; k < NDIM; k++){
         old_pos[i][k] = pos[i][k];
         old_vel[i][k] = vel[i][k];
         old_acc[i][k] = acc[i][k];
         old_jerk[i][k] = jerk[i][k];
   predict_step(pos, vel, acc, jerk, n, dt);
   get_acc_jerk_pot_coll(mass, pos, vel, acc, jerk, n, epot, coll_time);
   correct_step(pos, vel, acc, jerk, old_pos, old_vel, old_acc, old_jerk,
               n, dt);
   t += dt;
   delete[] old_pos;
   delete[] old_vel;
   delete[] old_acc;
   delete[] old_jerk;
}
```

As we have seen already in chapter 6, the Hermite code requires knowledge of the values of all four dynamical variables at the previous time step, indicated here by the prefix old. Since we do not want to introduce global variables, and since these variables are not needed outside the context of the current function, we allocate the memory in the first four lines, and free up those memory locations in the last four lines. If we now would omit those last four lines, the resulting memory leak could let us run into serious trouble. For example, taking a million time steps with a hundred-body system would cause us to loose 4*NDIM=12 words or 12*8=96 bytes for each particle for each time step, leading to a total memory loss of $96*10^2*10^6$ bytes or roughly ten Gbytes, which may well be larger than the core memory of the computer

at hand.

Again, it would be very good if we would check with each call to new whether there is still enough memory available. Since we do not do that here, a memory leak will suddenly cause the program to crash, without giving us any clue of where to look. Even using a debugger may not help, since the actual crash may well occur somewhere else, where a small amount of legitimate memory is requested, only to find out that all memory has just been exhausted elsewhere in the code. Once more, we will postpone but not neglect this type of defensive programming.

After the current values of the dynamical variables have been passed to the oldcopies, we take the first half of a Hermite pass, in a call to predict_step(), followed by a recalculation of accelerations and jerks, as well as potential energy and collision time scale. We are then ready to complete the Hermite step through a call to correct_step(), and update the time t.

8.12 The Predictor Step

The first half of a Hermite step is particularly simple, nothing more than a rather short Taylor series development:

Notice how much we can already read off from the way the arguments to predict_step() are declared: accelerations and jerks are passed as const variables, whereas positions and velocities are not. This implies that the latter two are updated, whereas the former

two are used to provide information for the update, without being changed themselves. This of course is exactly what happens.

8.13 The Corrector Step

The second half of a Hermite step is again a Taylor series development, this time to a higher order than in the predictor step, even though this is not obvious from the way it is written. We refer to the discussion in the beginning of chapter 6, where the Taylor series character of the corrector step is made explicit. Here is the code:

```
/*-----
   correct_step -- takes one iteration to improve the new values of position
                    and velocities, effectively by using a higher-order
                    Taylor series constructed from the terms up to jerk at
                    the beginning and the end of the time step.
void correct_step(real pos[][NDIM], real vel[][NDIM],
                const real acc[][NDIM], const real jerk[][NDIM],
                const real old_pos[][NDIM], const real old_vel[][NDIM],
                const real old_acc[][NDIM], const real old_jerk[][NDIM],
                int n, real dt)
{
   for (int i = 0; i < n; i++)
       for (int k = 0; k < NDIM; k++){
           vel[i][k] = old_vel[i][k] + (old_acc[i][k] + acc[i][k])*dt/2
                                  + (old_jerk[i][k] - jerk[i][k])*dt*dt/12;
          pos[i][k] = old_pos[i][k] + (old_vel[i][k] + vel[i][k])*dt/2
                                   + (old_acc[i][k] - acc[i][k])*dt*dt/12;
       }
}
```

8.14 Where All the Work is Done

We now arrive at the core function of nbody_sh1.C, where all the hard work is being done. In addition, this function is both the longest and the most complicated among the ten functions in the file. The main reason for the complexity is that we are trying to accomplish four things in one function, as the name indicates. While calculating accelerations and jerks are logically related, the calculation of the potential energy and the collision time is more a matter of convenience with little natural or logical relation

to the calculation of the first two. The main reason for bundling these four operations is efficiency. Here is the code:

```
/*-----
   get_acc_jerk_pot_coll -- calculates accelerations and jerks, and as side
                             effects also calculates potential energy and
                             the time scale coll_time for significant changes
                             in local configurations to occur.
                                           j
                                                            ji ji -->
                                       ----- v
                     r
                                j
                                       |--> |3 | ji
                                                           |--> |2 ji |
             |--> |3
                      ji
                                 jі
             r
                                       | r | |
                                                           r
             | ji |
                                       | ji | |__
                                                           | ji |
   note: it would be cleaner to calculate potential energy and collision time
         in a separate function. However, the current function is by far the
         most time consuming part of the whole program, with a double loop
         over all particles that is executed every time step. Splitting off
         some of the work to another function would significantly increase
         the total computer time (by an amount close to a factor two).
   We determine the values of all four quantities of interest by walking
   through the system in a double {i,j} loop. The first three, acceleration,
   jerk, and potential energy, are calculated by adding successive terms;
   the last, the estimate for the collision time, is found by determining the
   minimum value over all particle pairs and over the two choices of collision
   time, position/velocity and sqrt(position/acceleration), where position and
   velocity indicate their relative values between the two particles, while
   acceleration indicates their pairwise acceleration. At the start, the
   first three quantities are set to zero, to prepare for accumulation, while
   the last one is set to a very large number, to prepare for minimization.
        The integration loops only over half of the pairs, with j > i, since
   the contributions to the acceleration and jerk of particle j on particle i
   is the same as those of particle i on particle j, apart from a minus sign
   and a different mass factor.
void get_acc_jerk_pot_coll(const real mass[], const real pos[][NDIM],
                         const real vel[][NDIM], real acc[][NDIM],
                         real jerk[][NDIM], int n, real & epot,
                         real & coll_time)
{
   for (int i = 0; i < n; i++)
```

```
for (int k = 0; k < NDIM; k++)
            acc[i][k] = jerk[i][k] = 0;
    epot = 0;
    const real VERY_LARGE_NUMBER = 1e300;
    real coll_time_q = VERY_LARGE_NUMBER;
                                               // collision time to 4th power
                                               // collision time scale estimate
    real coll_est_q;
                                               // to 4th power (quartic)
    for (int i = 0; i < n; i++){
        for (int j = i+1; j < n; j++){
                                                   // rji[] is the vector from
            real rji[NDIM];
                                                   // particle i to particle j
           real vji[NDIM];
                                                   // vji[] = d rji[] / d t
            for (int k = 0; k < NDIM; k++){
                rji[k] = pos[j][k] - pos[i][k];
                vji[k] = vel[j][k] - vel[i][k];
           }
            real r2 = 0;
                                                   // | rji |^2
           real v2 = 0;
                                                   // | vji |^2
            real rv_r2 = 0;
                                                   // ( rij . vij ) / | rji |^2
            for (int k = 0; k < NDIM; k++){
                r2 += rji[k] * rji[k];
                v2 += vji[k] * vji[k];
                rv_r2 += rji[k] * vji[k];
            }
            rv_r2 /= r2;
                                                   // | rji |
            real r = sqrt(r2);
           real r3 = r * r2;
                                                   // | rji |^3
// add the {i,j} contribution to the total potential energy for the system:
            epot -= mass[i] * mass[j] / r;
// add the {j (i)} contribution to the {i (j)} values of acceleration and jerk:
            real da[3];
                                                   // main terms in pairwise
            real dj[3];
                                                   // acceleration and jerk
            for (int k = 0; k < NDIM; k++){
                da[k] = rji[k] / r3;
                                                               // see equations
                dj[k] = (vji[k] - 3 * rv_r2 * rji[k]) / r3;  // in the header
            }
            for (int k = 0; k < NDIM; k++){
                acc[i][k] += mass[j] * da[k];
                                                              // using symmetry
                acc[j][k] -= mass[i] * da[k];
                                                              // find pairwise
                jerk[i][k] += mass[j] * dj[k];
                                                             // acceleration
                jerk[j][k] -= mass[i] * dj[k];
                                                              // and jerk
           }
```

// first collision time estimate, based on unaccelerated linear motion:

```
coll_est_q = (r2*r2) / (v2*v2);
            if (coll\_time\_q > coll\_est\_q)
                coll_time_q = coll_est_q;
// second collision time estimate, based on free fall:
            real da2 = 0:
                                                             // da2 becomes the
            for (int k = 0; k < NDIM; k++)
                                                             // square of the
                da2 += da[k] * da[k];
                                                            // pair-wise accel-
            double mij = mass[i] + mass[j];
                                                            // eration between
            da2 *= mij * mij;
                                                             // particles i and j
            coll_est_q = r2/da2;
            if (coll_time_q > coll_est_q)
                coll_time_q = coll_est_q;
        }
    }
                                                     // from q for quartic back
    coll_time = sqrt(sqrt(coll_time_q));
                                                     // to linear collision time
}
```

Notice the distribution of const declarations here, which is just the opposite from what we saw in predict_step() and correct_step(). In the latter two accelerations and jerk were const while positions and velocities were updated. Here the roles are reversed. In addition, there are two variables that are called by reference, epot and coll_time, which enable the information about potential energy and collision time to flow back to the calling function evolve_step() and from there back to evolve(), where they are used, as we have seen above.

After preparing the proper initial values for the four variables of interest, we enter the $\{i,j\}$ loop running over all particle pairs. As we have seen in the previous two chapters, we first compute a number of auxiliary quantities before we are ready to calculate first the contribution of a pair of particles to the potential energy and then their mutual contributions to each others acceleration and jerk.

At the end of the loop, we compute the two different collision time step estimates, in the same way we discovered at the end of the previous chapter. The first estimate follows the approximate of unperturbed linear motion, extrapolating current separation and rate of change of separation in order to guess when the particles will change their relative configuration substantially. The second estimate neglects the current rate of change of the pairwise separation, estimating instead the free-fall time of the two particles, in case they would start off at rest. In practice, the smaller of the two estimates provides a reasonably safe estimate for the time scale on which significant changes in configuration can occur.

8.15 Closing Logo

At the very end of our file, we add a simpler version of the gravitylab logo that we encountered at the top of the file:

It contains the name of the file, for consistency, and it guarantees that no part of the file has been truncated in a process of copying, editing or transmission over the net. While such mishaps are very rare nowadays, they still can occur occasionally, and it seems prudent to mark the intended end of the file. Meanwhile, our intrepid observer has changed directions from which to observe the world.